UNIVERSITY OF SASKATCHEWAN

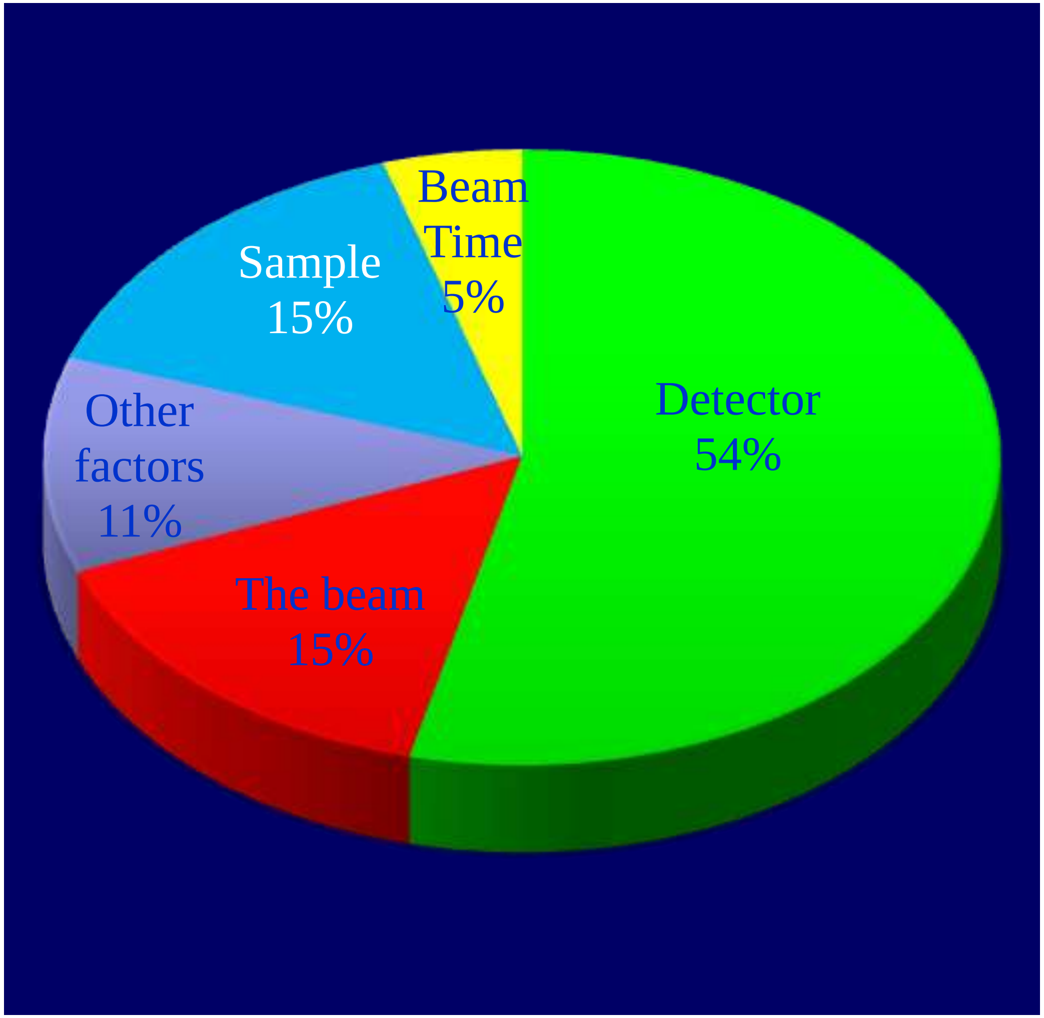
MONASH University

Detectors for Synchrotron Radiation

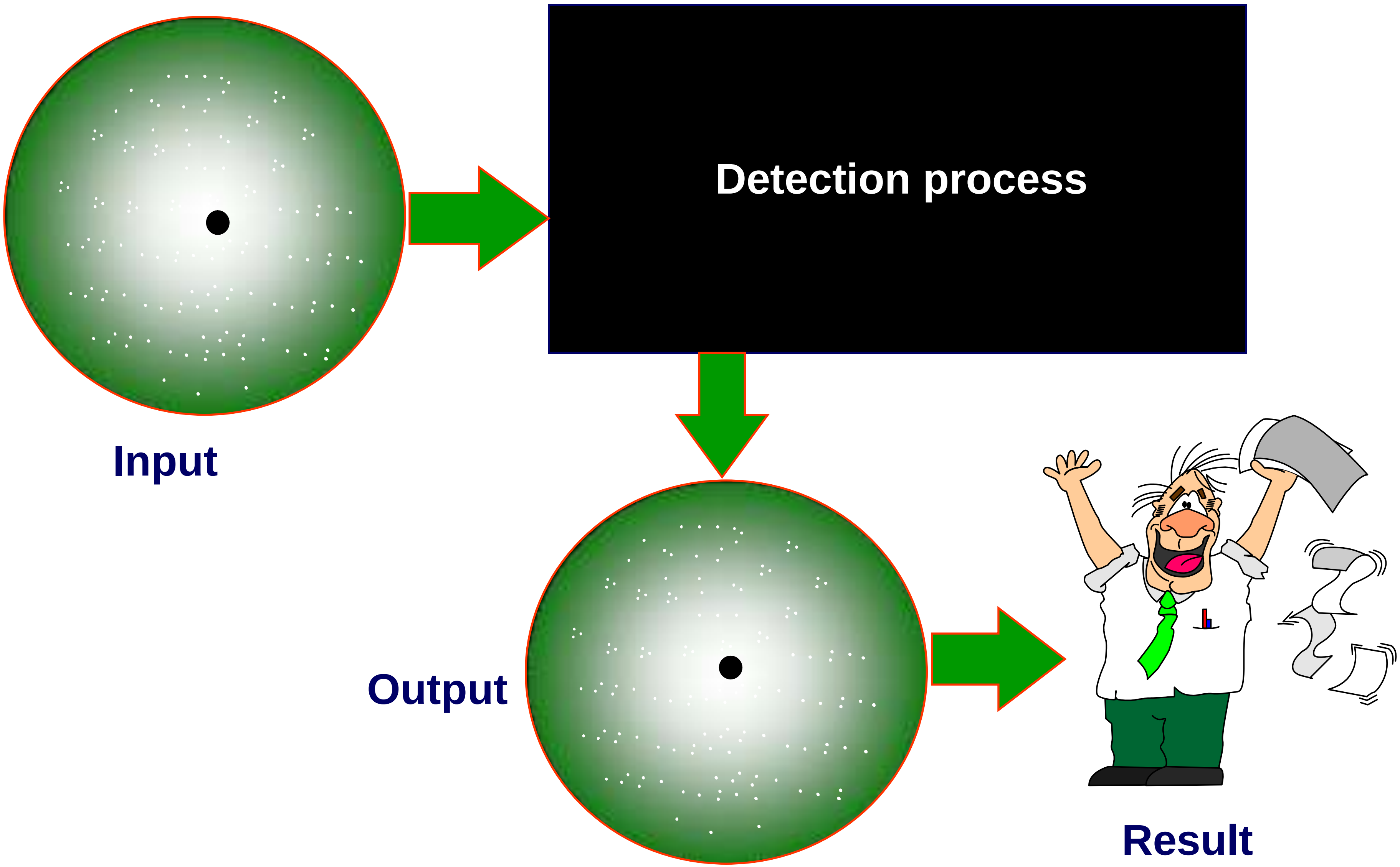
Rob Lewis

Factors Limiting Science

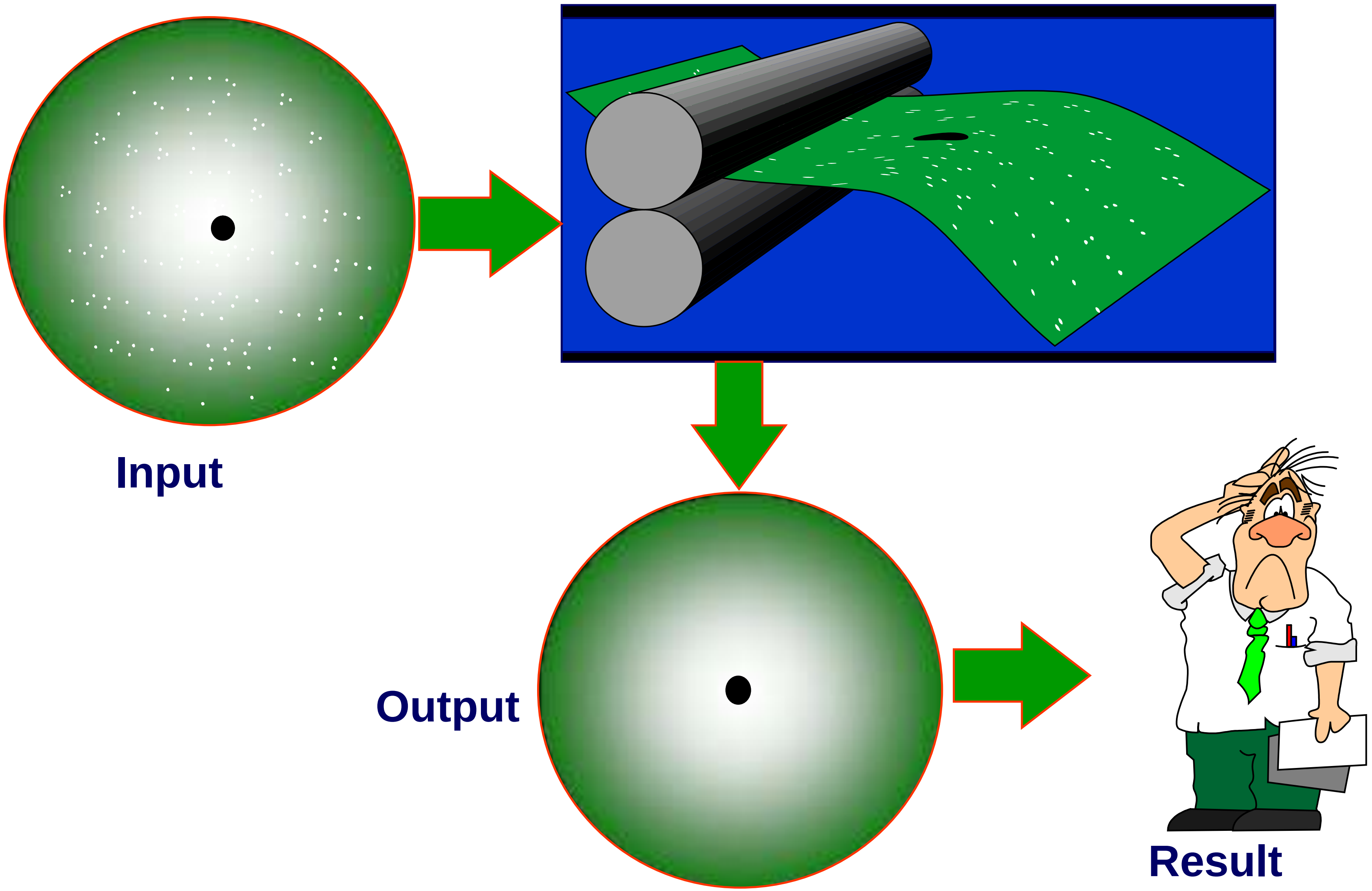
- Detectors are an oft-neglected but crucial part of an experiment
- They often limit the science



Scientist's View of Detector



The Truth!



Detection Mechanisms

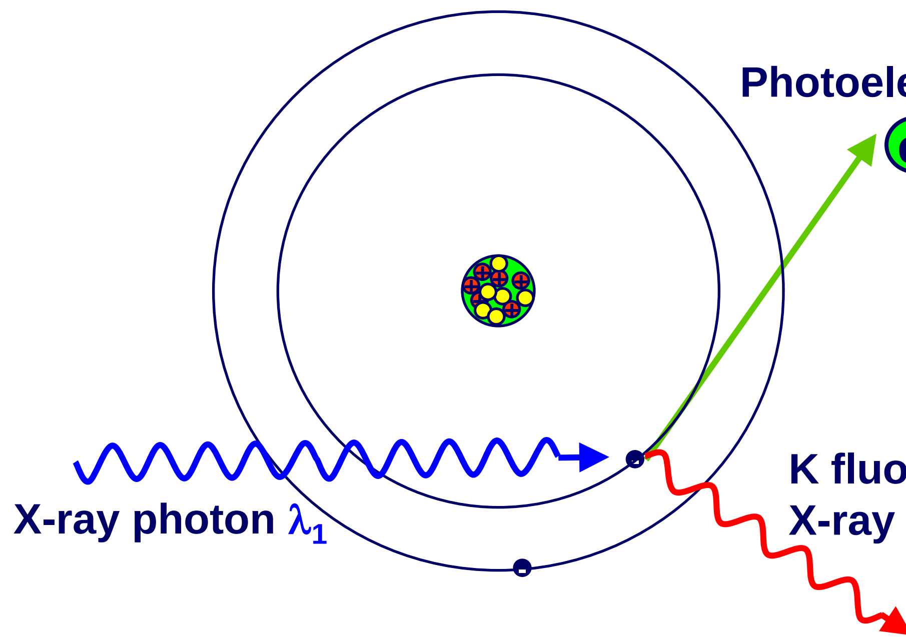
- There are many means of detection. All require the interaction of photons with matter
- Examples include
  - Gas ionisation
    - Photons produce electrons and ions which are then detected
    - E.g. Ion chambers, proportional counters
  - Photoelectric effect
    - Photons eject electrons from a solid creating a current which is measured
    - E.g.. Beam monitors
  - Generation of electron hole pairs
    - Photons produce electrons and holes in a semiconductor which are then detected
    - E.g.. CCD
  - Fluorescence, scintillation and F centres
    - Photons produce prompt fluorescence or F centres
    - E.g. Image plates and Scintillation counters
  - Chemical effect
    - Photons create a chemical change such as dissociating Ag halide
    - E.g. Film

Albert Einstein

Germany and Switzerland
Kaiser-Wilhelm-Institut
(now Max-Planck-Institut)
für Physik
Berlin-Dahlem, Germany
1879 - 1955

Nobel prize in physics 1921
"for his services to Theoretical Physics, and especially for his discovery of the law of the photoelectric effect"

Photoelectric Effect



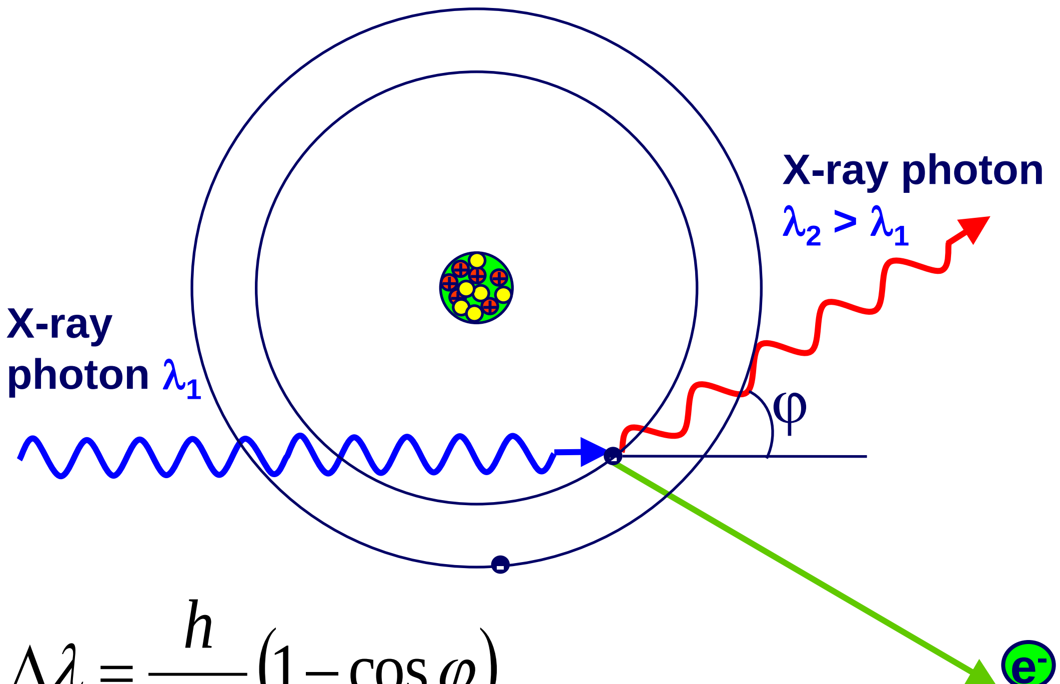
Arthur Holly Compton

University of Chicago
Chicago, IL, USA
1892 - 1962

Nobel prize in physics 1927

"for his discovery of the effect named after him"

Compton Effect

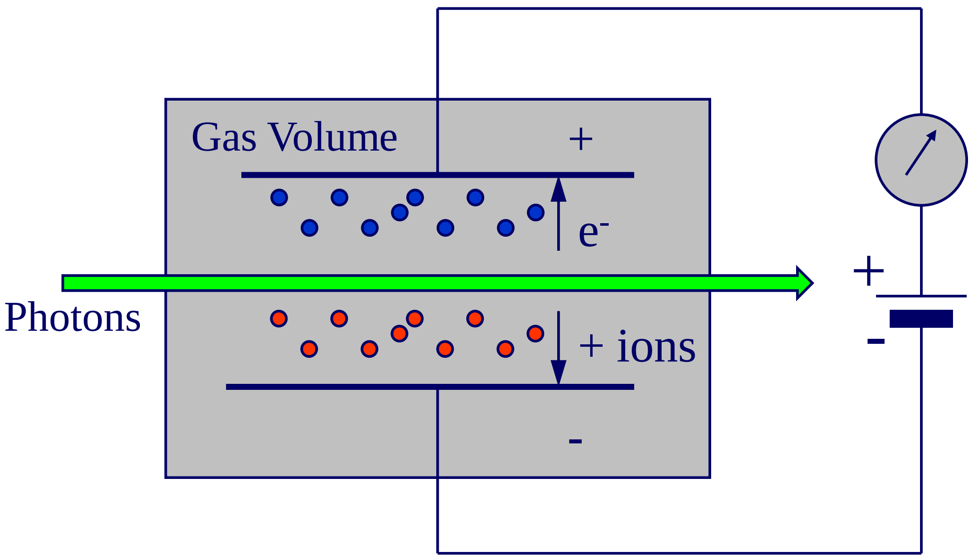


$$\Delta\lambda = \frac{h}{m_0 c}(1 - \cos\varphi)$$

An Example Detector

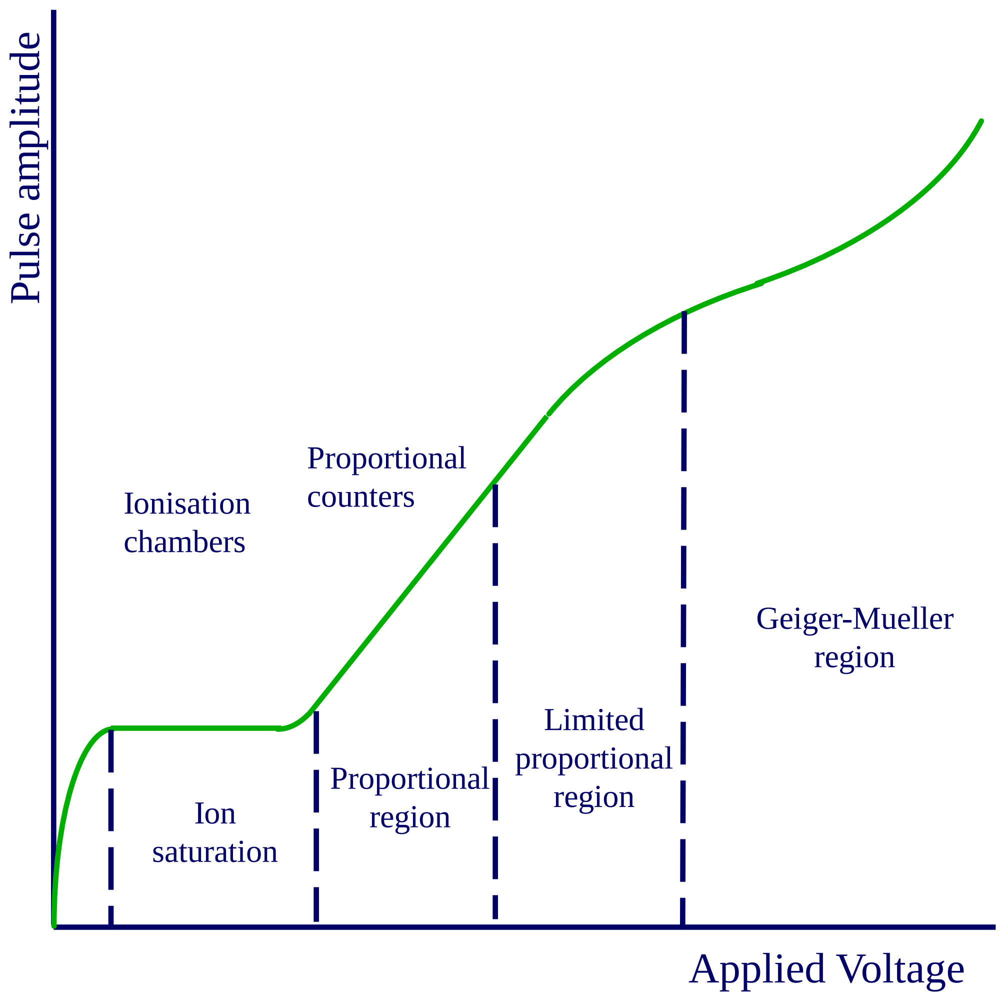
Echidna

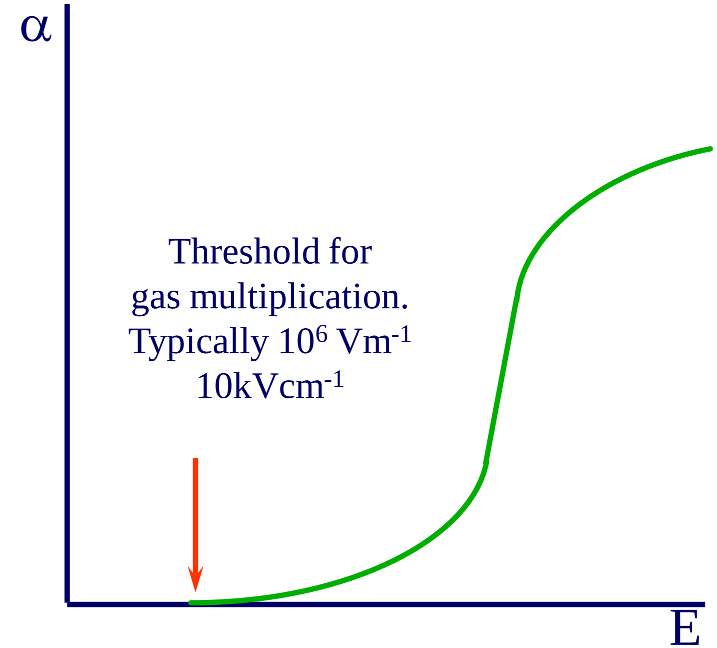
Ionisation Chamber



- Very simple device
- Approximately 1 e^- ion pair per 30eV deposited
- Important that recombination low as possible
  - Higher voltages required at higher rates since more carriers
  - Diffusion losses caused by separation of carriers minimised by higher voltages
  - Plates too close cause electron losses
- Ion chambers are sensitive to pressure and temperature

Operation regions of gas filled detectors





n is number of charges

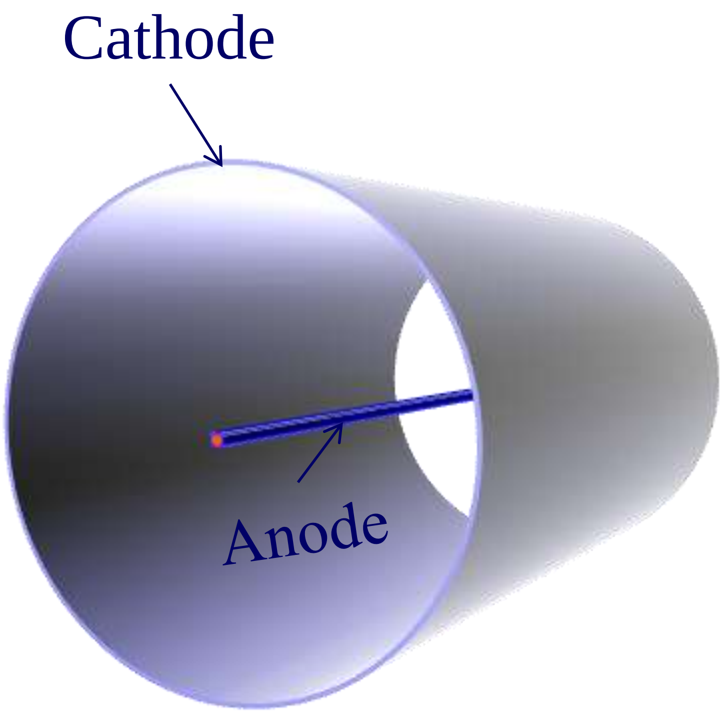
x is distance

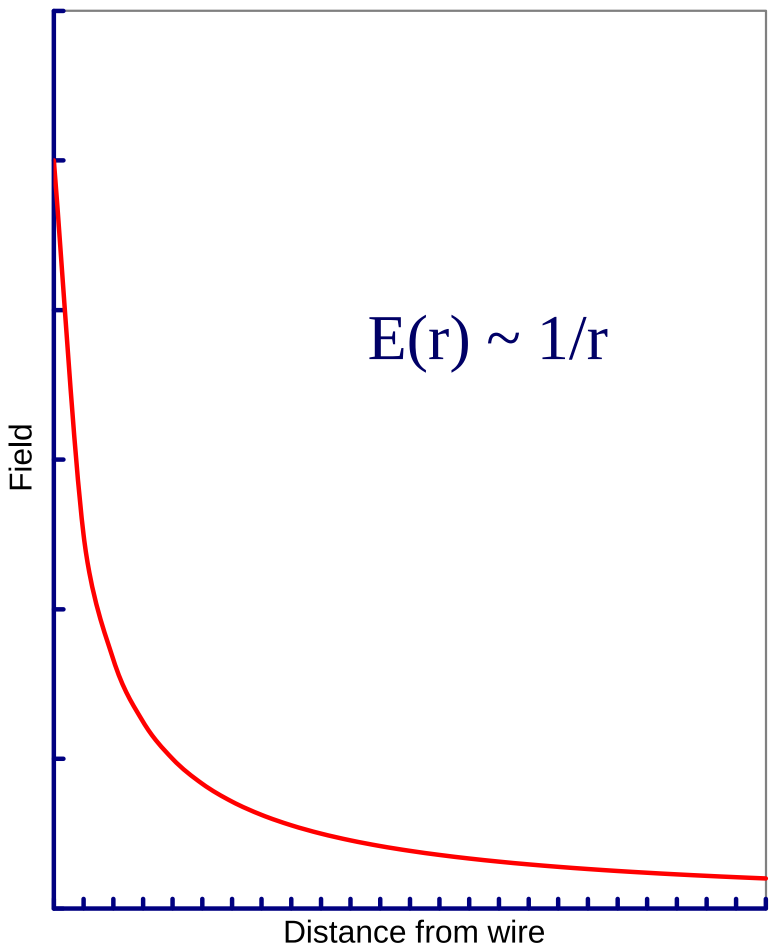
α is the first Townsend coefficient

$$\frac{dn}{n} = \alpha dx$$

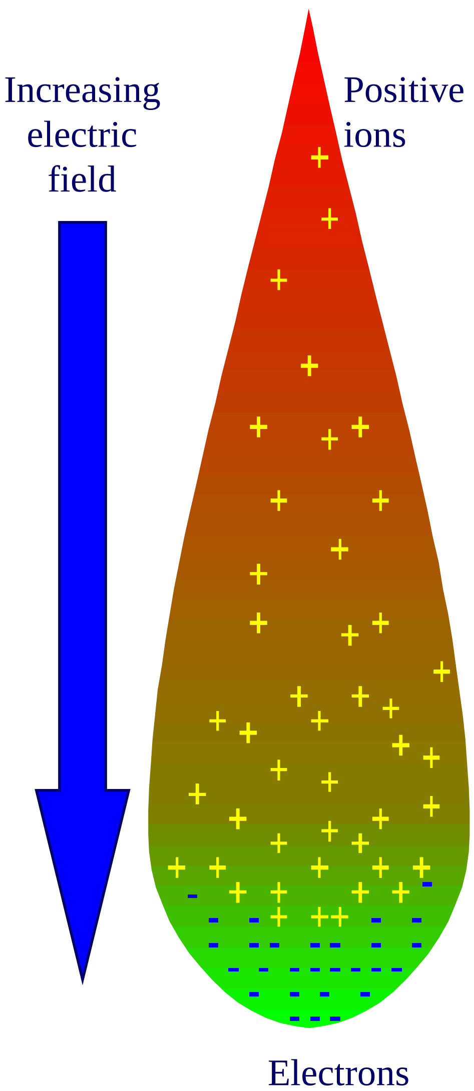
$$n(x) = n(0)e^{\alpha x}$$

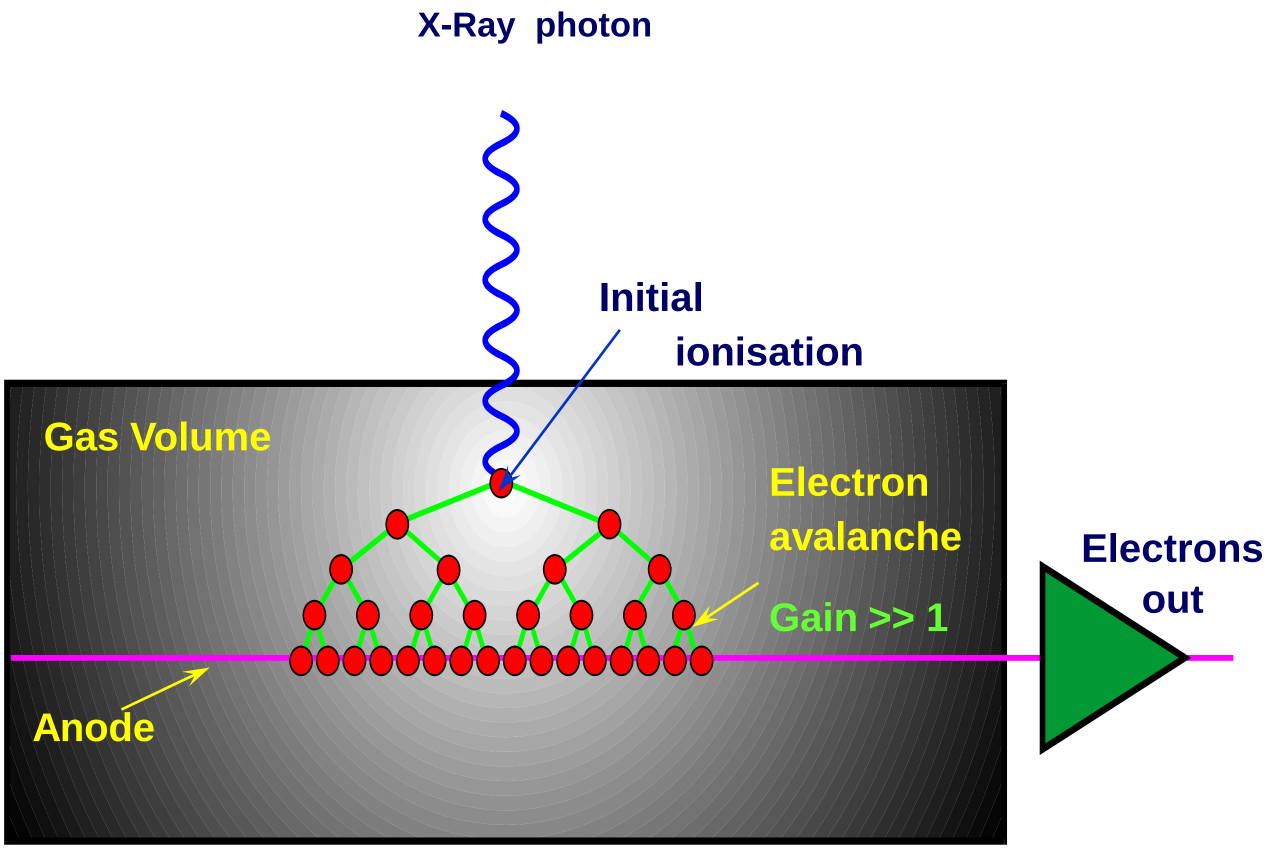
Field Variation



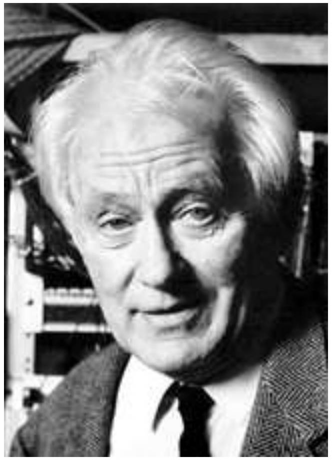


Avalanche & Proportional Counter





Georges Charpak



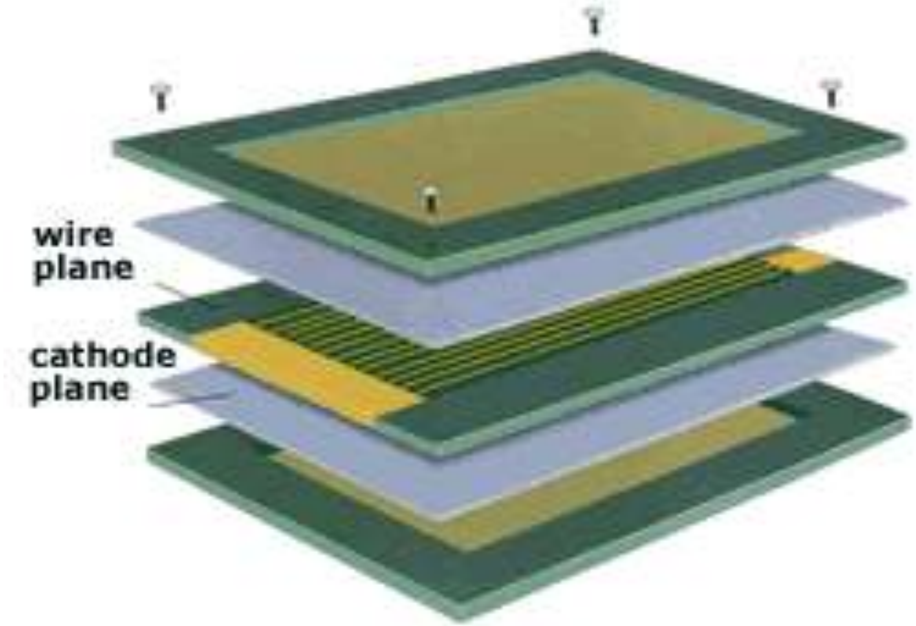
France
École Supérieure de
Physique et Chimie
Paris, France; CERN
Geneva, Switzerland

b. 1924
(in Dabrovica, Poland)

Nobel prize in physics 1992

"for his invention and development of particle detectors, in particular the multiwire proportional chamber"

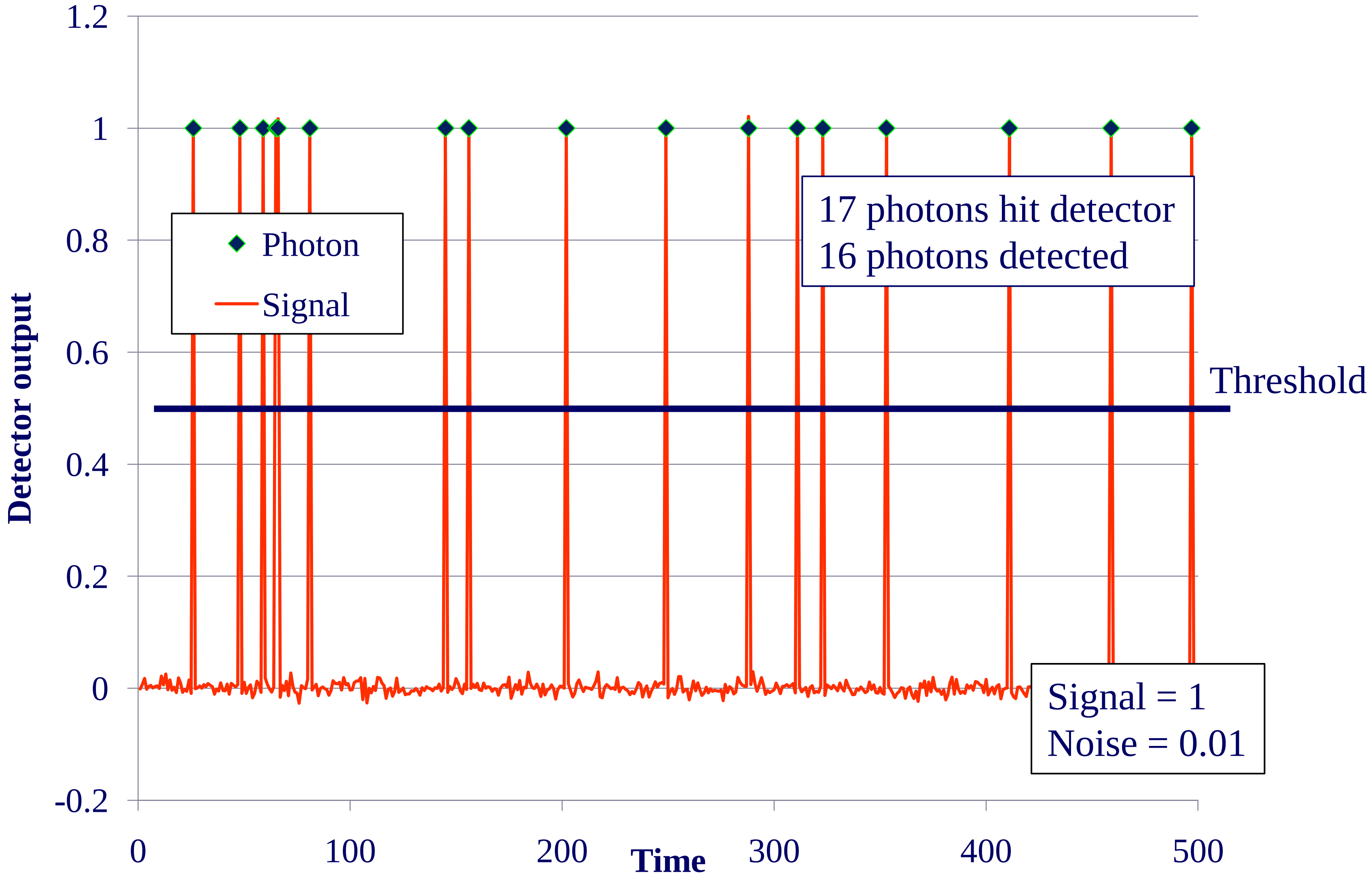
Multi-wire Proportional Counter



Counting and Integrating

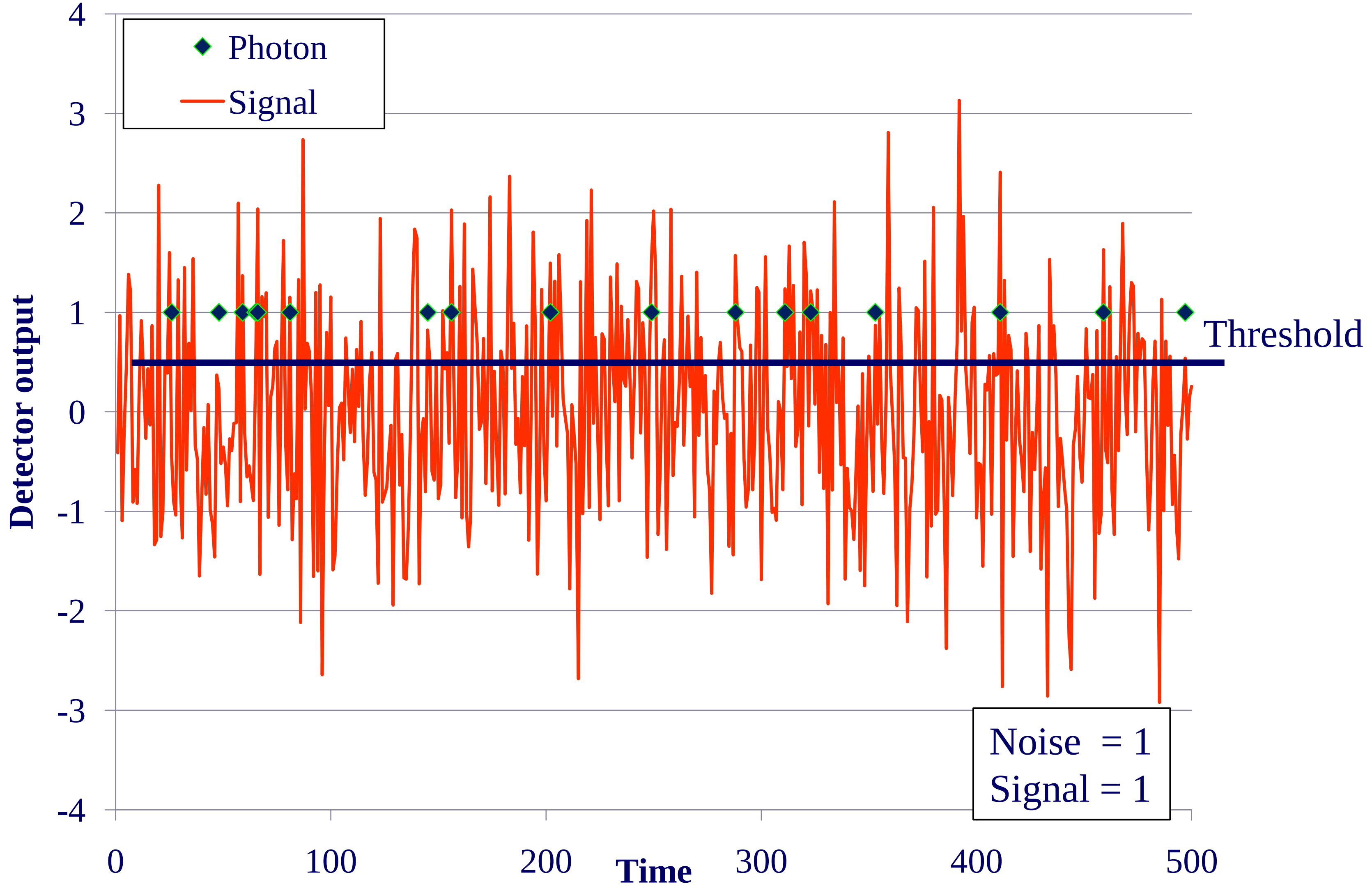
- If there is sufficient signal produced by the interaction of a photon or a particle in the detector then it is possible to operate the detector as a counter
- It’s all about signal to noise ratio!

SNR = 100



1.2
1
0.8
0.6
0.4
0.2
0
-0.2
Detector output
Photon
Signal
17 photons hit detector
16 photons detected
Threshold
Signal = 1
Noise = 0.01
0
100
200
300
400
500
Time

SNR = 1

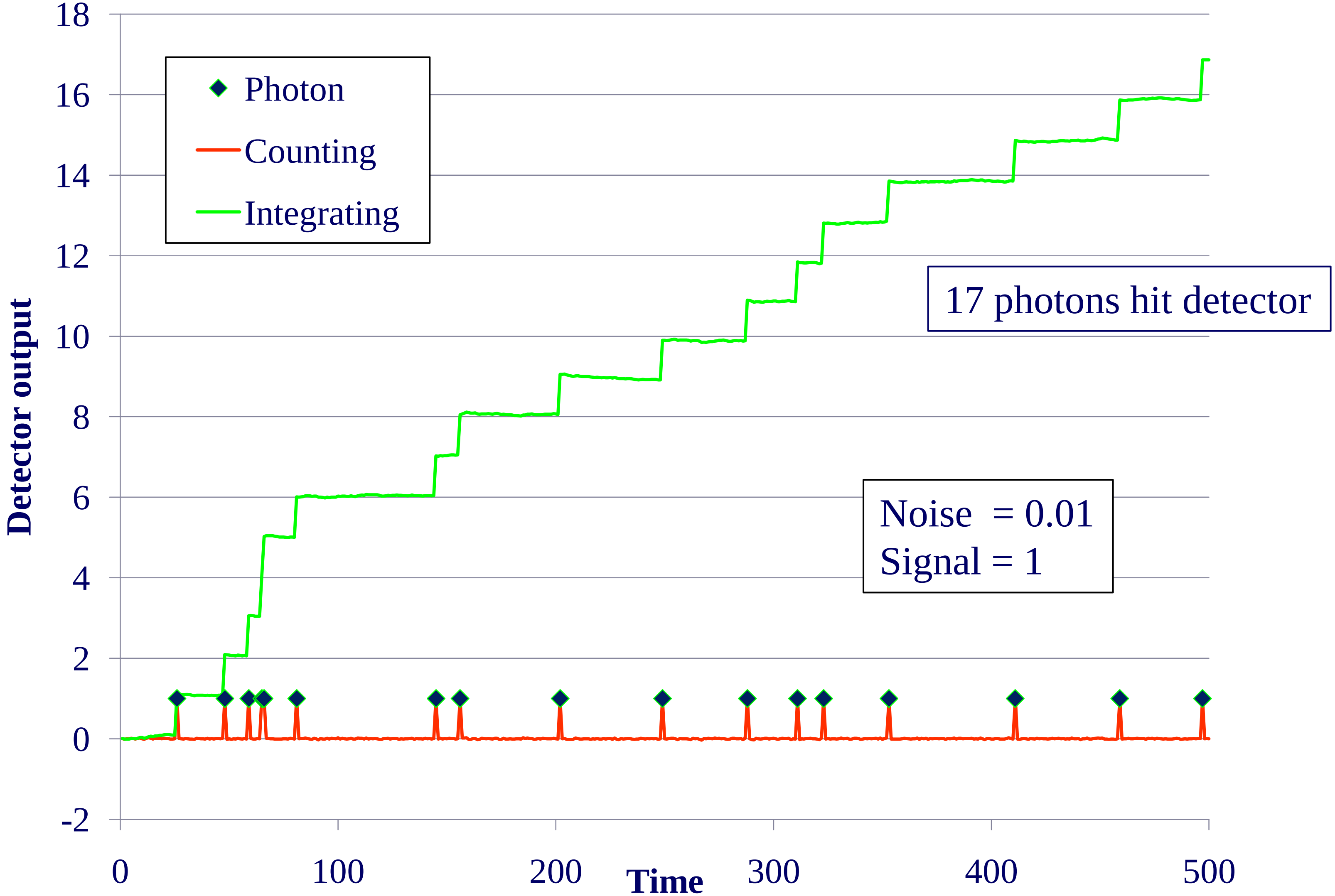


Photon
Signal
Threshold
Noise = 1
Signal = 1
Detector output
Time
4
3
2
1
0
-1
-2
-3
-4
0
100
200
300
400
500

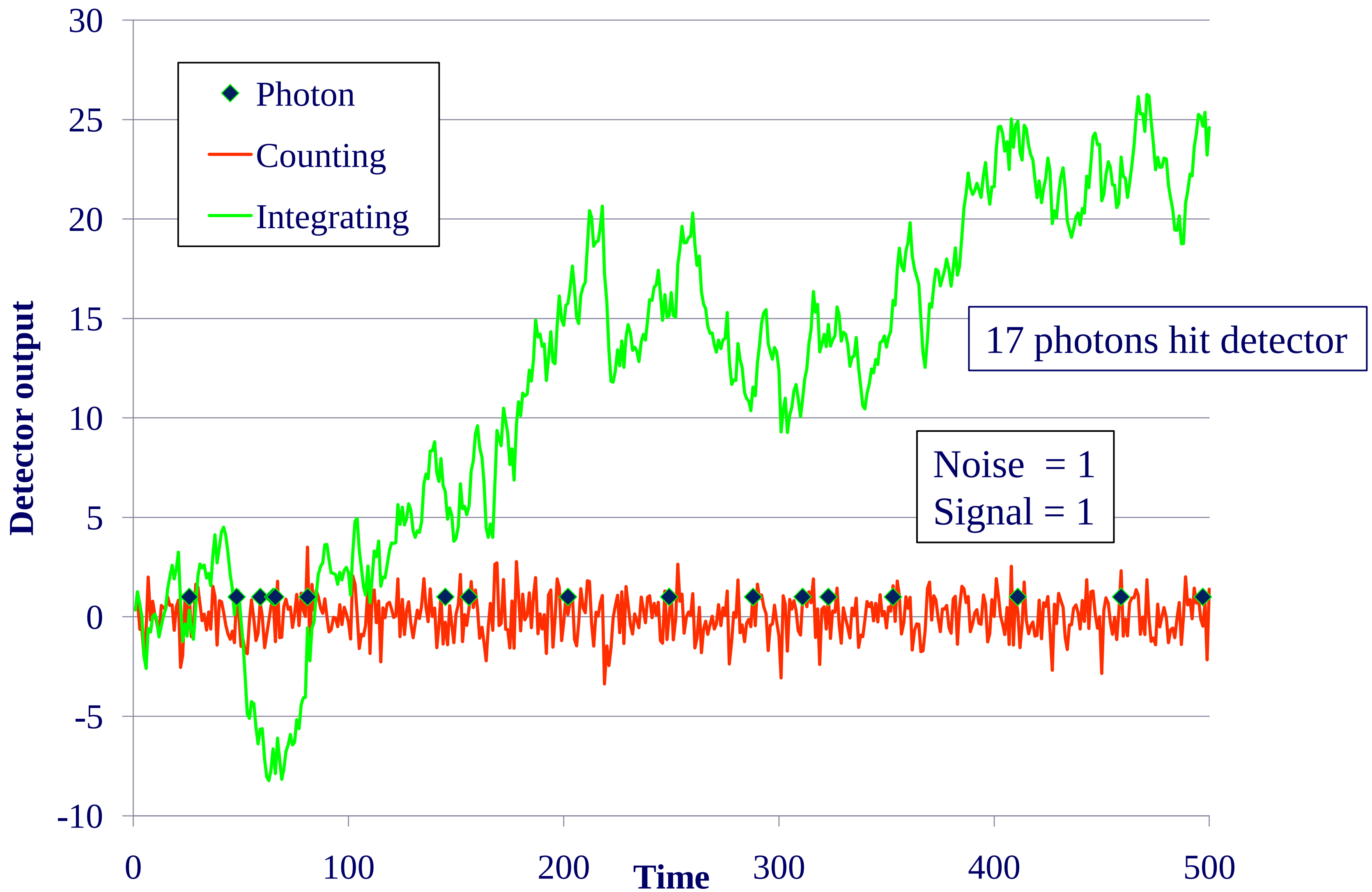
Counting and Integrating

- Usually SNR is insufficient and we have to accumulate many photons/particles before the signal becomes measurable

Counting & Integrating SNR =100

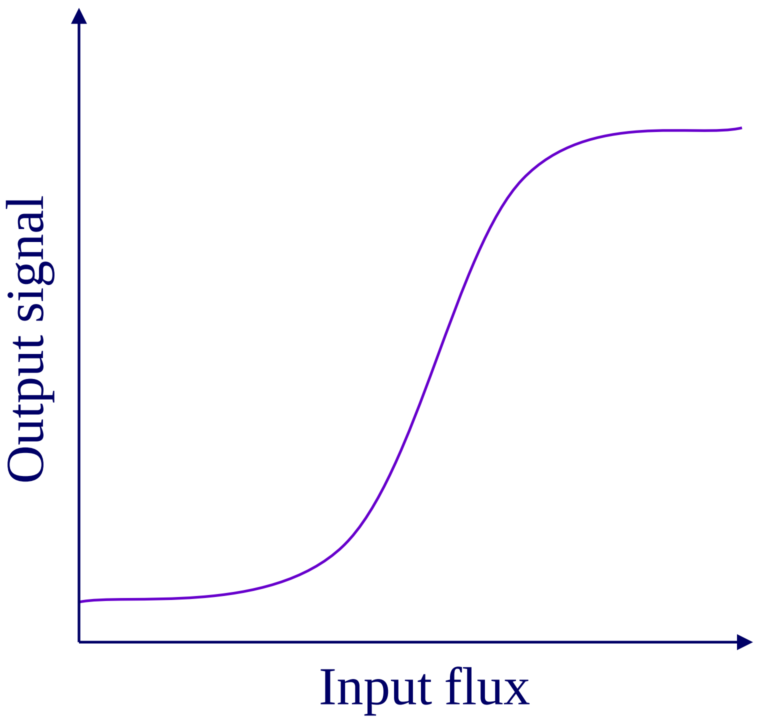


Counting & Integrating SNR = 1



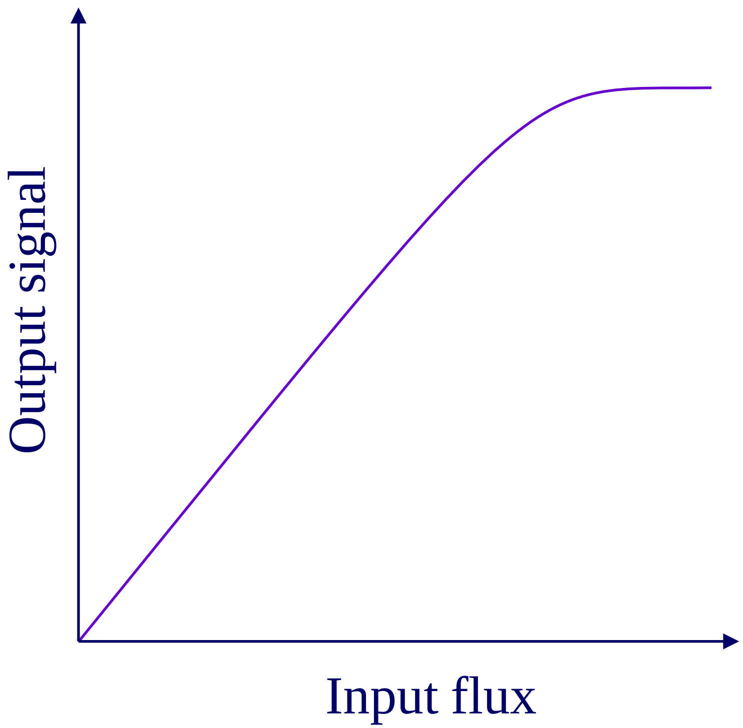
Integrating Detectors

- Mode
  - Measures deposited energy at end of integration period
- Characteristics
  - High input flux capability
  - Read noise dominates at low signal (“fog level”)
  - Dead time between frames
  - 2×20 keV phts = 1×40 keV photon i.e. Cannot perform simultaneous spectroscopy and positioning
  - Examples: Image plates, CCDs



Photon Counting Detectors

- Mode
  - Detects every photon as it arrives. Only active pixels read
- Characteristics
  - Quantum limited, Detector noise often negligible
  - No dead time between frames
  - Can measure position and energy simultaneously
  - Limited input flux capability
  - Examples: Prop counters, Scintillators

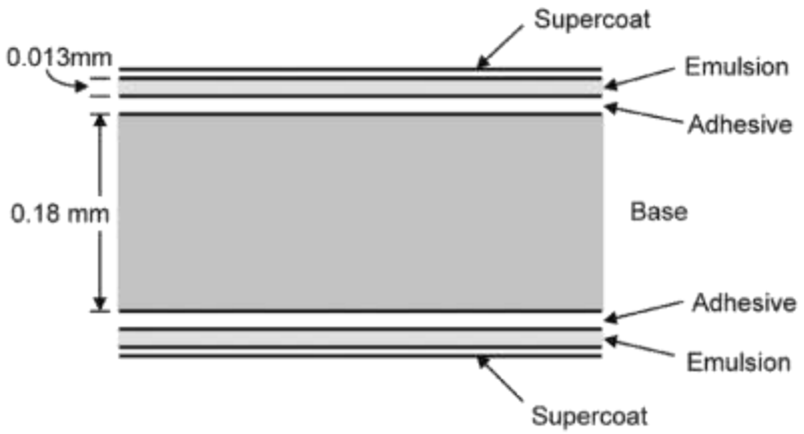


Types of Detectors

Crimson Rosella and King Parrot

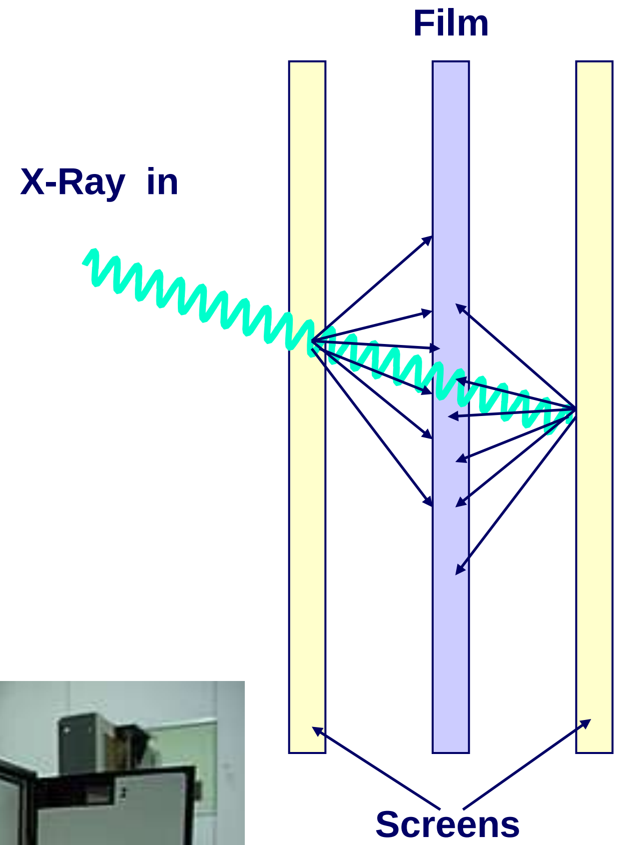
X-ray Film

- Active Ingredient
  - Small crystals of silver halide ~1.0 - 1.5μm
  - Typically 90-99% silver bromide and 1-10% silver iodide.
  - Suspended in the gelatin of the film emulsion.
  - Crystals have a cubic lattice with many point defects and free silver ions
- Exposure
  - A photon liberates an electron from a bromide ion
  - The electron travels until trapped at a defect
  - A free silver ion is attracted to the negative charge and combines (is reduced) to form an atom of metallic silver (which is optically black).
  - The single silver atom acts as an electron trap for another electron which then attracts another atom of silver which is then reduced to metallic silver. This process continues while the exposure to light continues.
- Conventional film is usually coated with emulsion on only one side, radiographic film is usually double coated (on each side of the base) to be used with intensifying screens.



Intensifying Screens

- An intensifying screen converts x-ray energy into light energy
- X-rays are absorbed by the phosphor
- The phosphor becomes excited & fluoresces emitting UV and/or visible light
- For every x-ray photon absorbed, hundreds of light photons are emitted
- The use of intensifying screens inevitably means that certain degree of unsharpness will be introduced into the image in comparison to non-screen film



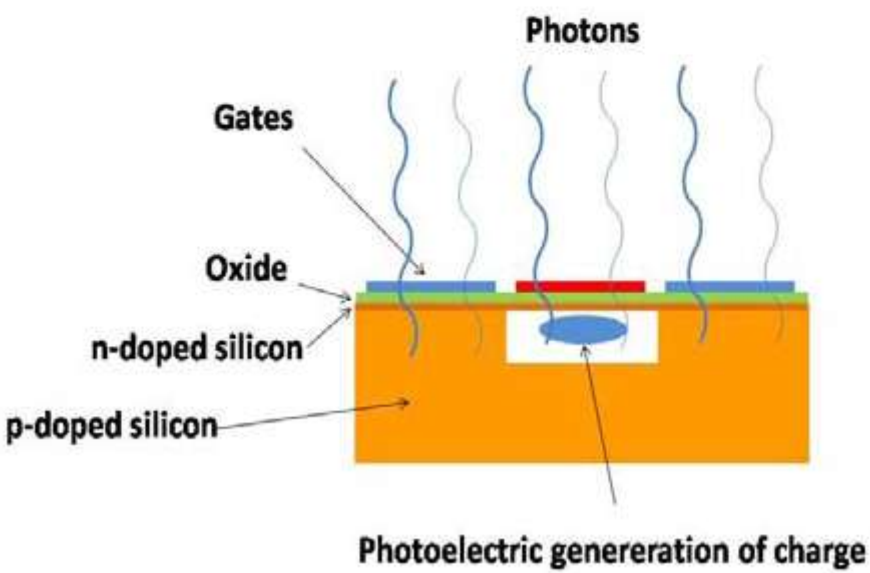
Willard S. Boyle & George E. Smith

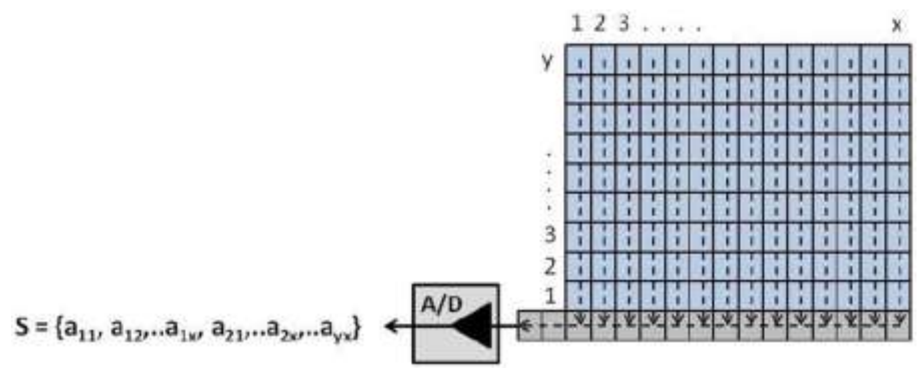
Bell Laboratories
Murray Hill, NJ, USA

Nobel prize in physics 2009

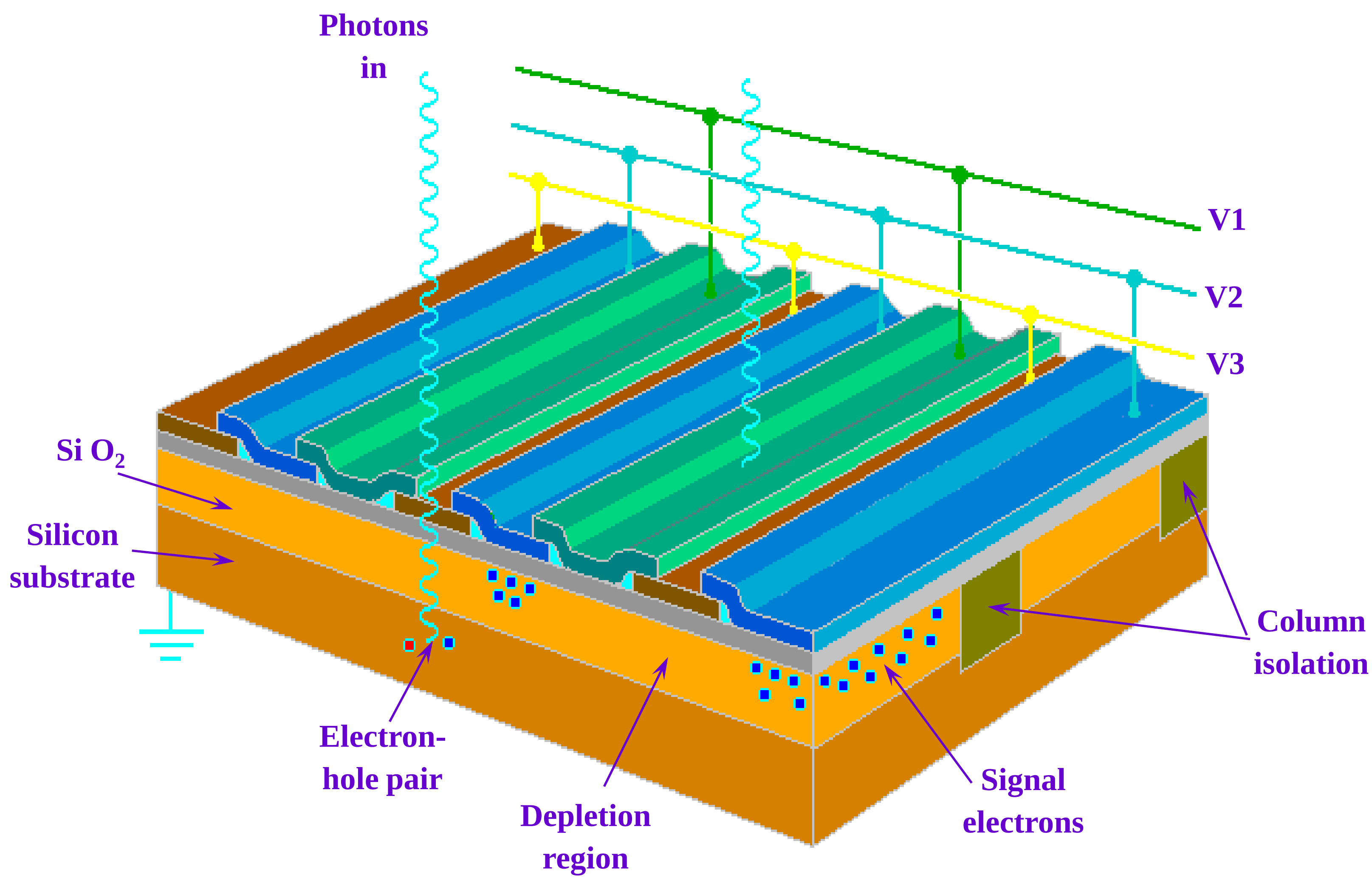
"for the invention of an imaging semiconductor circuit – the CCD sensor"

CCD

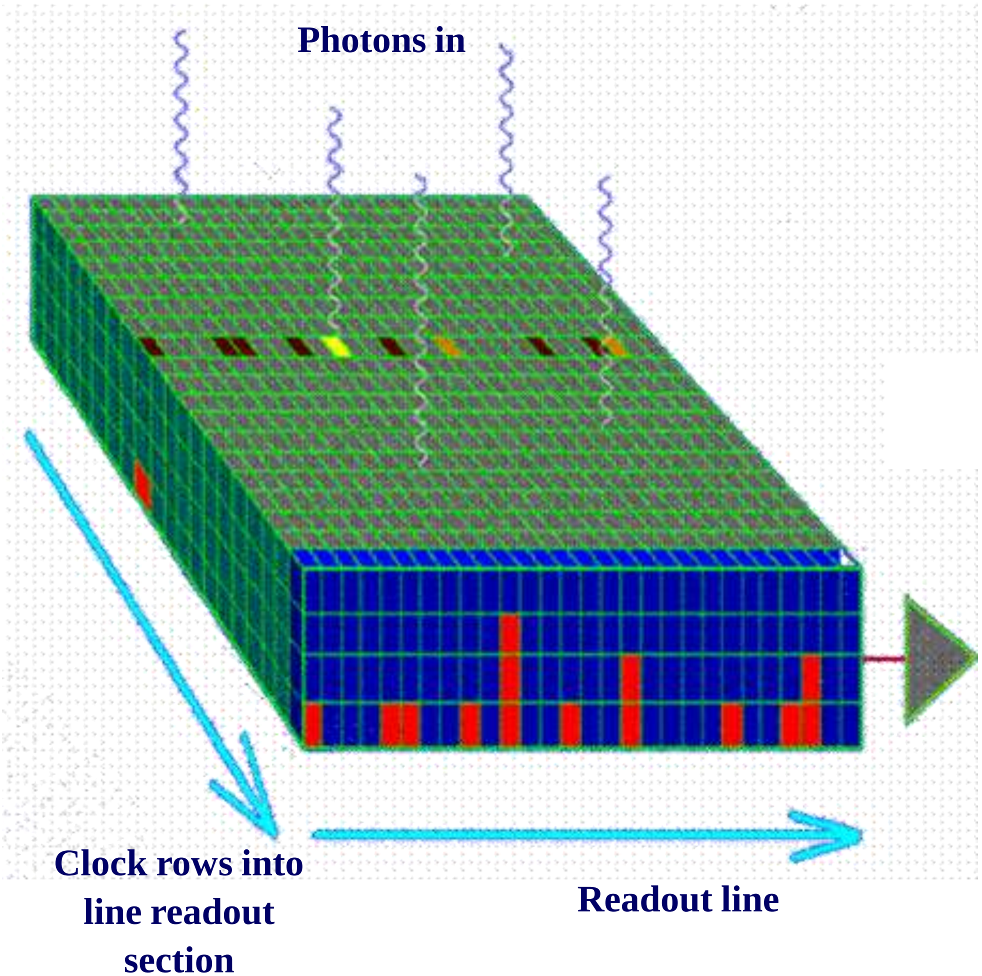


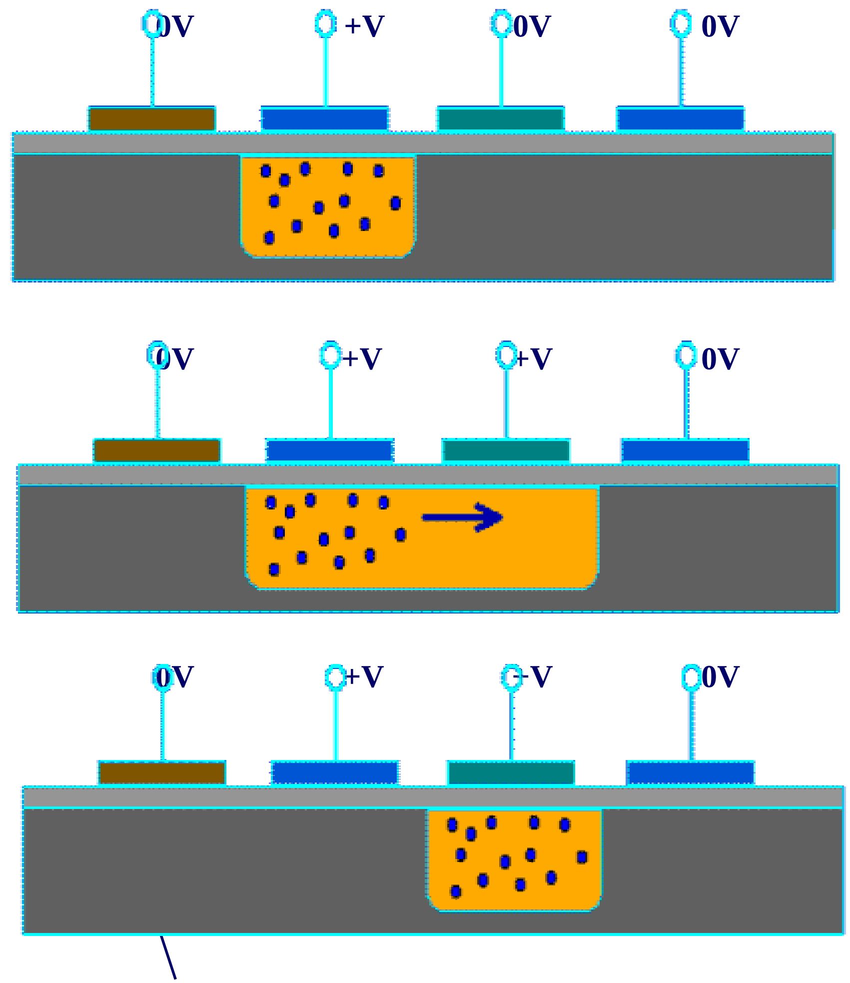


Charge Coupled Device



CCD Readout

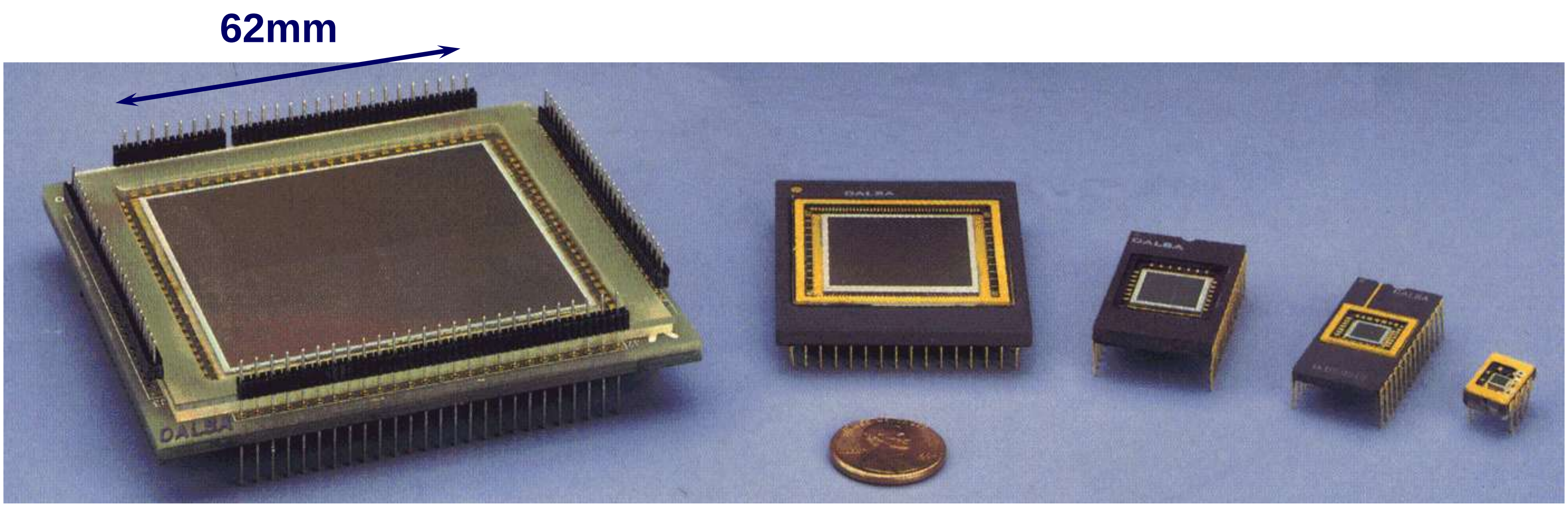

Photons in
Clock rows into line readout section
Readout line


0V
+V
0V
0V
0V
+V
+V
0V
0V
+V
0V

CCD Readout

- Charge is moved from pixel to pixel by clocking
- Each pixel has a limited capacitance (well depth) typically 10^4-10^5 e^-
- This limits dynamic range for direct detection
  - 10keV photon creates ~ 3000e^- so saturation = ~ 10 photons
- Speed of clocking is restricted by line capacitance and charge transfer efficiency
  - Size of CCD restricted by this
- Noise can be reduced by cooling
- Amplifier usually on chip
  - Heats up that part of chip

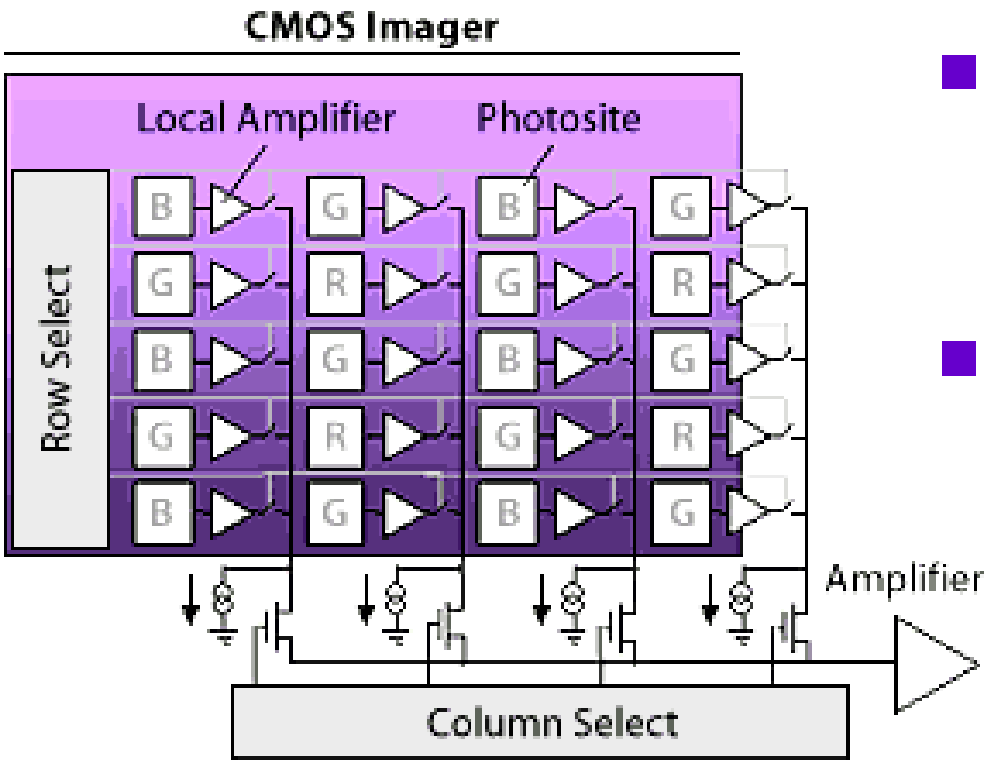
CCDs



Although sizes > 50mm are available, the read speed is slow to preserve low noise and cte (line capacitance becomes very high)

Shutter required

Complimentary Metal-Oxide Semiconductor (CMOS)



- A readout amplifier transistor on each pixel converts charge to voltage
- Allows random access to pixels, similar to the row-column memory cell access in RAM

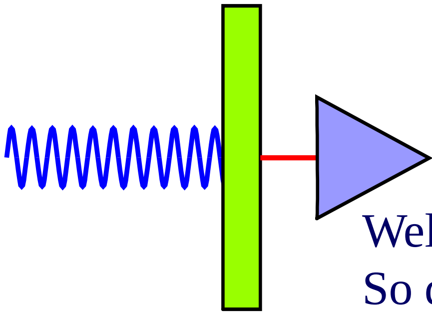
CMOS vs CCD

- Traditionally CCD higher sensitivity and lower noise
- Modern lithography means they are now similar
- CMOS sensors can have much more functionality on-chip than CCDs
  - On chip image processing, edge detection, noise reduction, and analog to digital conversion
- CMOS lower power → less heat → less noise

Use with X-rays

Direct detection

Gain ~ 2000e$^-$ / 8keV x-ray



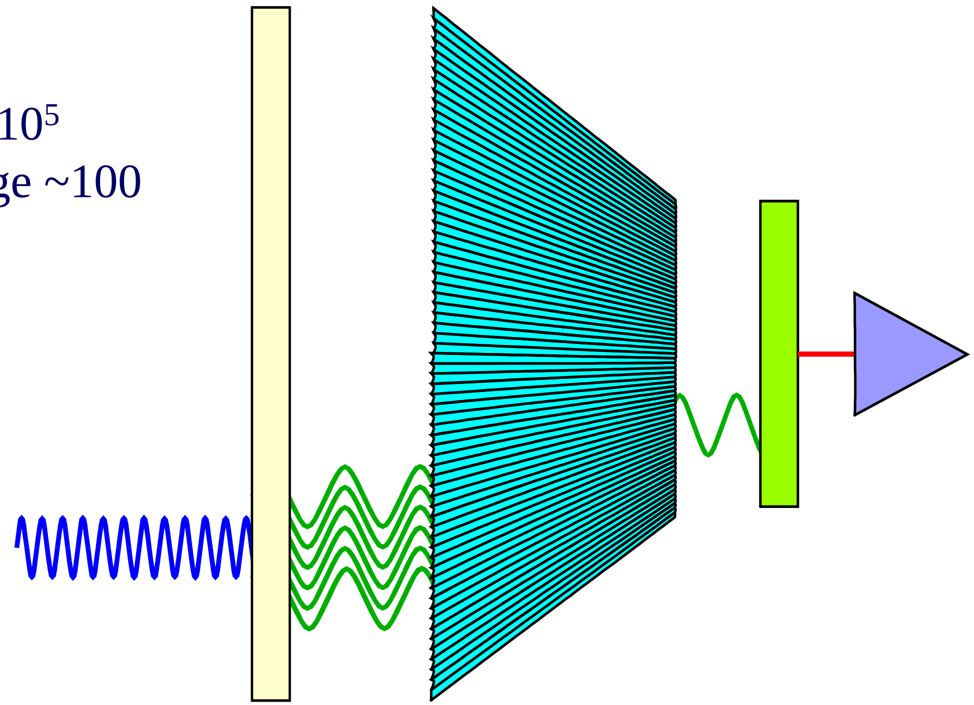
Well depth = 2×10^5

So dynamic range ~100

Phosphor coupled with reducing optics to CCD

Phosphor gain >> 1

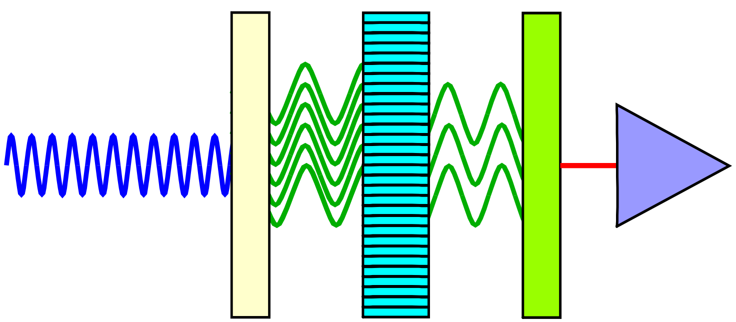
Optics Gain << 1



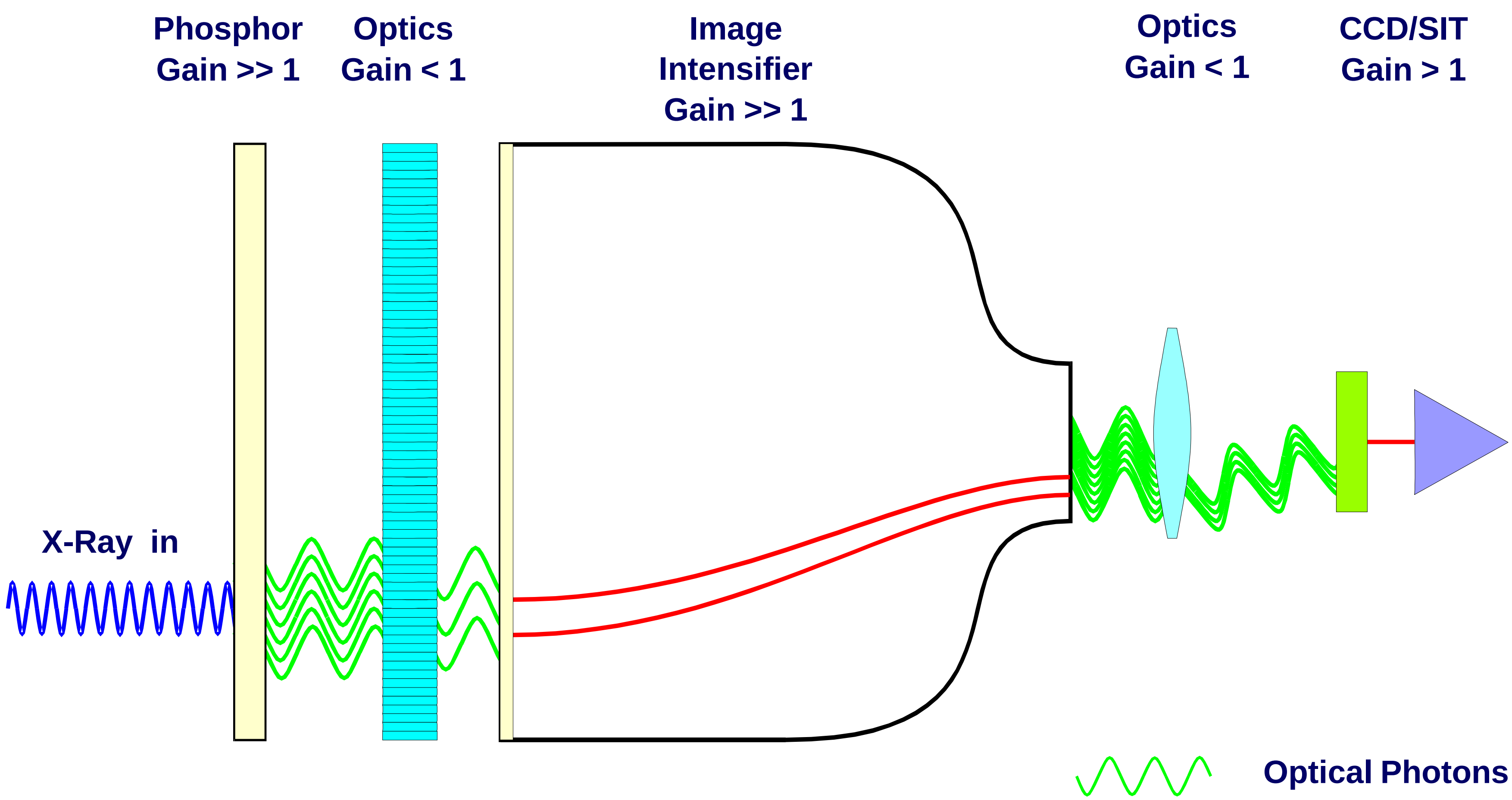
Phosphor coupled 1:1 to CCD

Phosphor gain >> 1

Optics Gain < 1



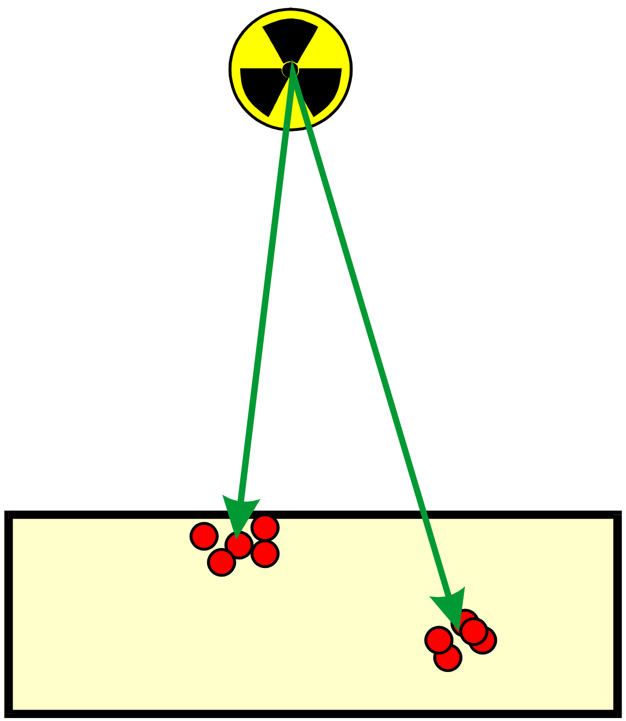
TV detector with IIT



Computed Radiography-Image Plate

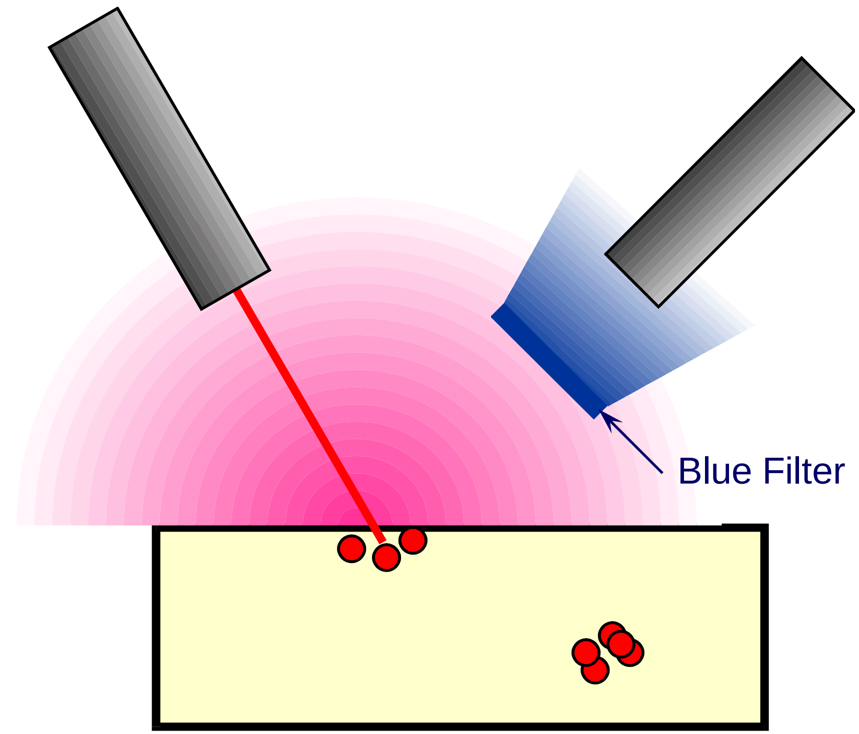
Exposure

Creation of F centres
Gain >> 1

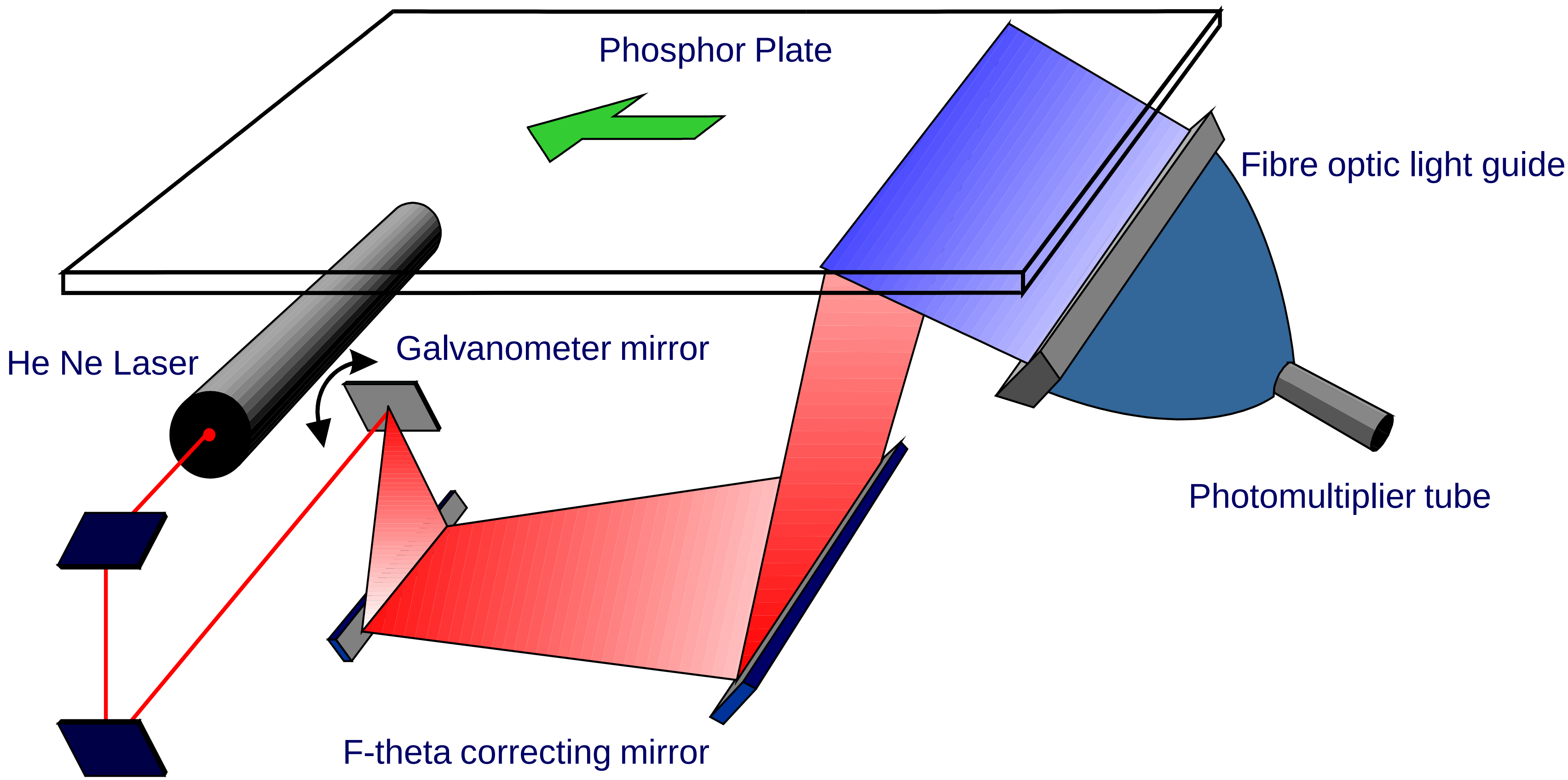


Scanning

Stimulation of PSL	Collection of PSL	PMT Amplification
Gain < 1	Gain < 1	Gain > 1

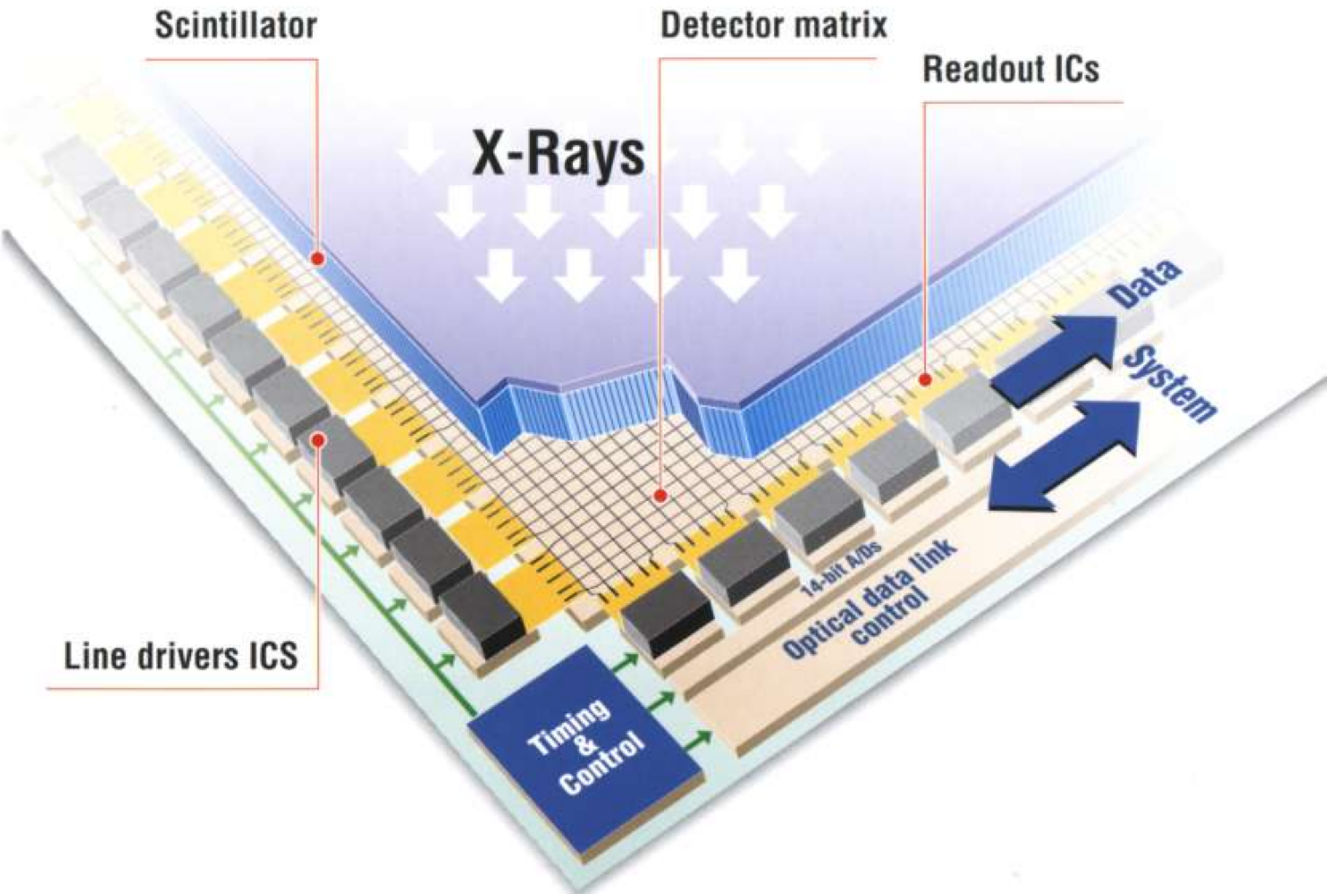


X-Y Flat bed Scanner



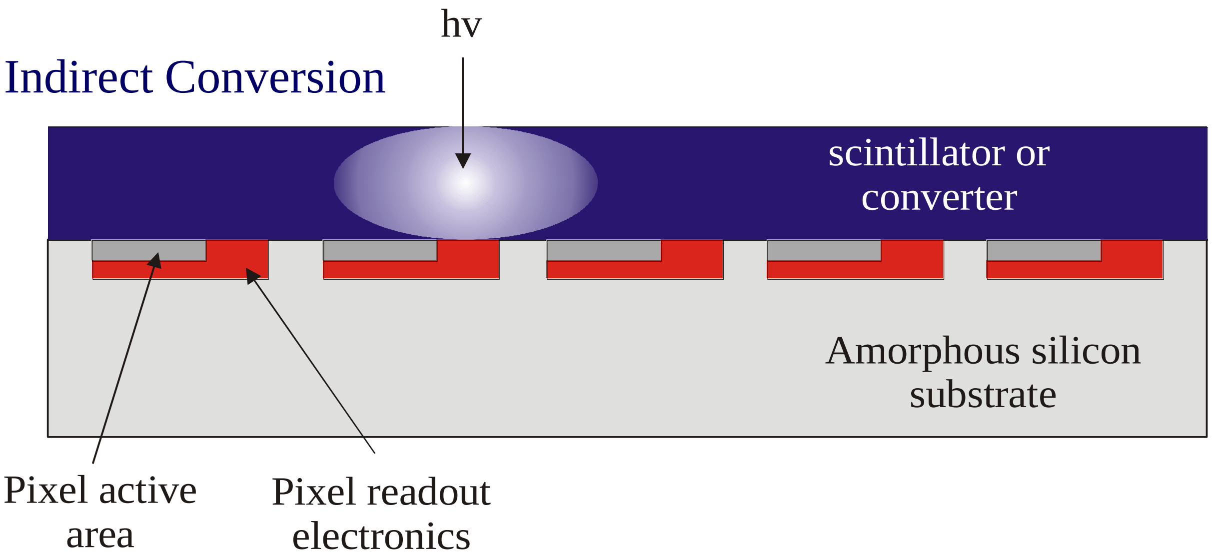
Distributed Light Collection

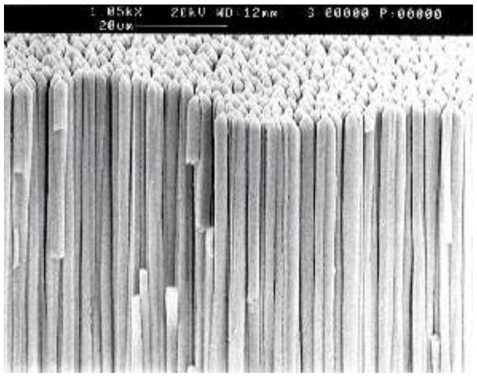
TFT Flat panel Detector



a-Si:H TFT arrays

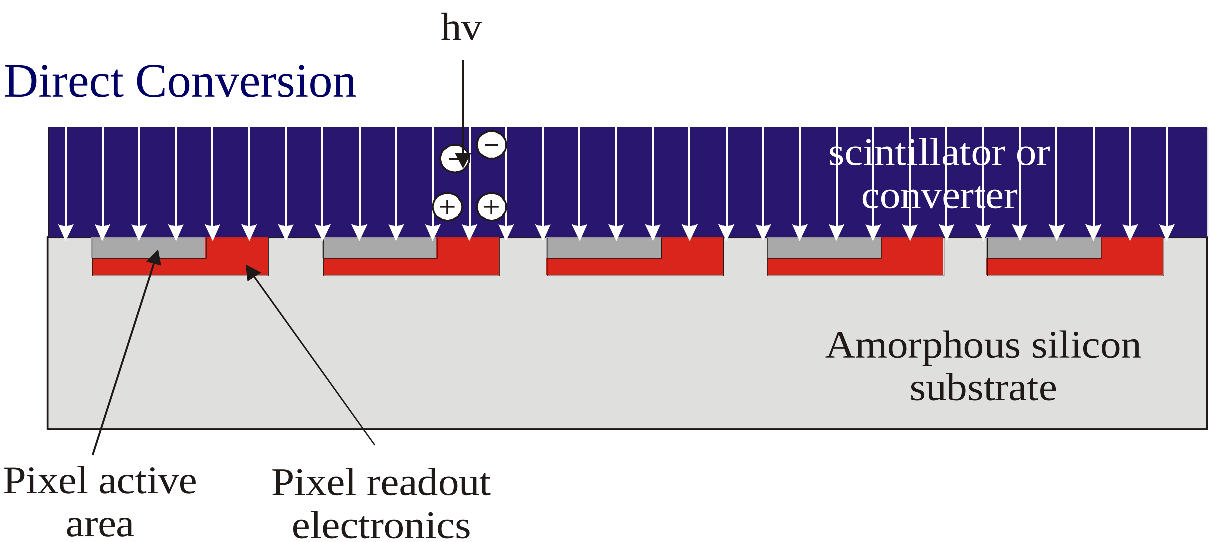
Indirect Conversion



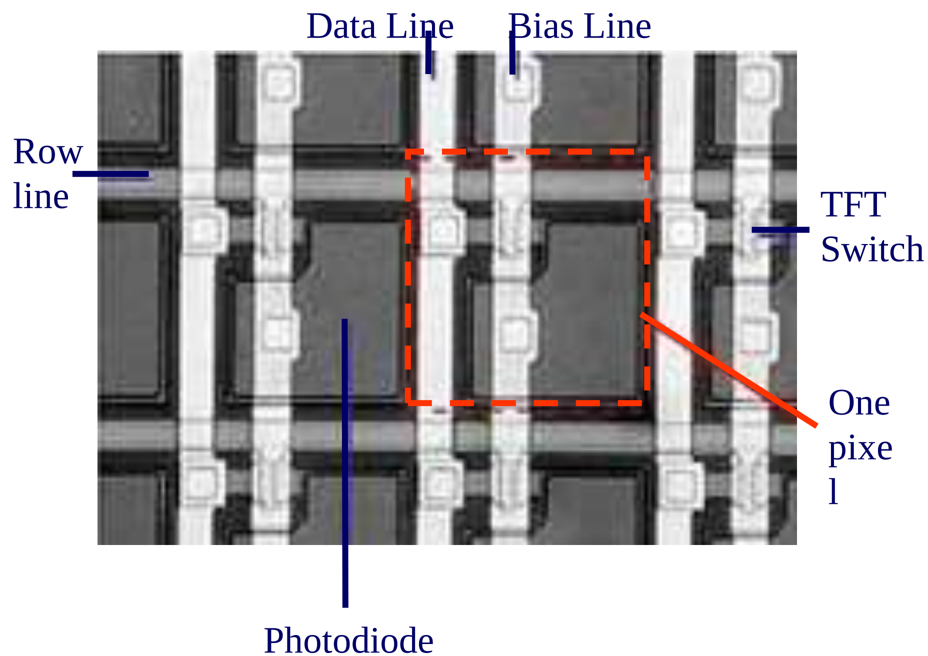


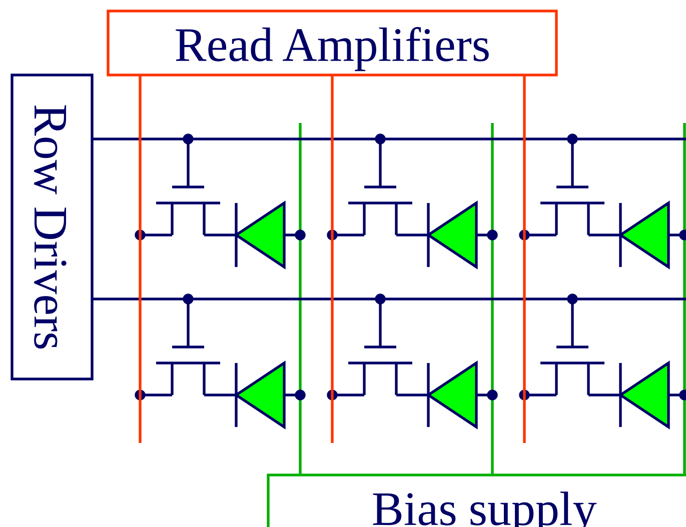
Needle diameter 6μm

Direct Conversion

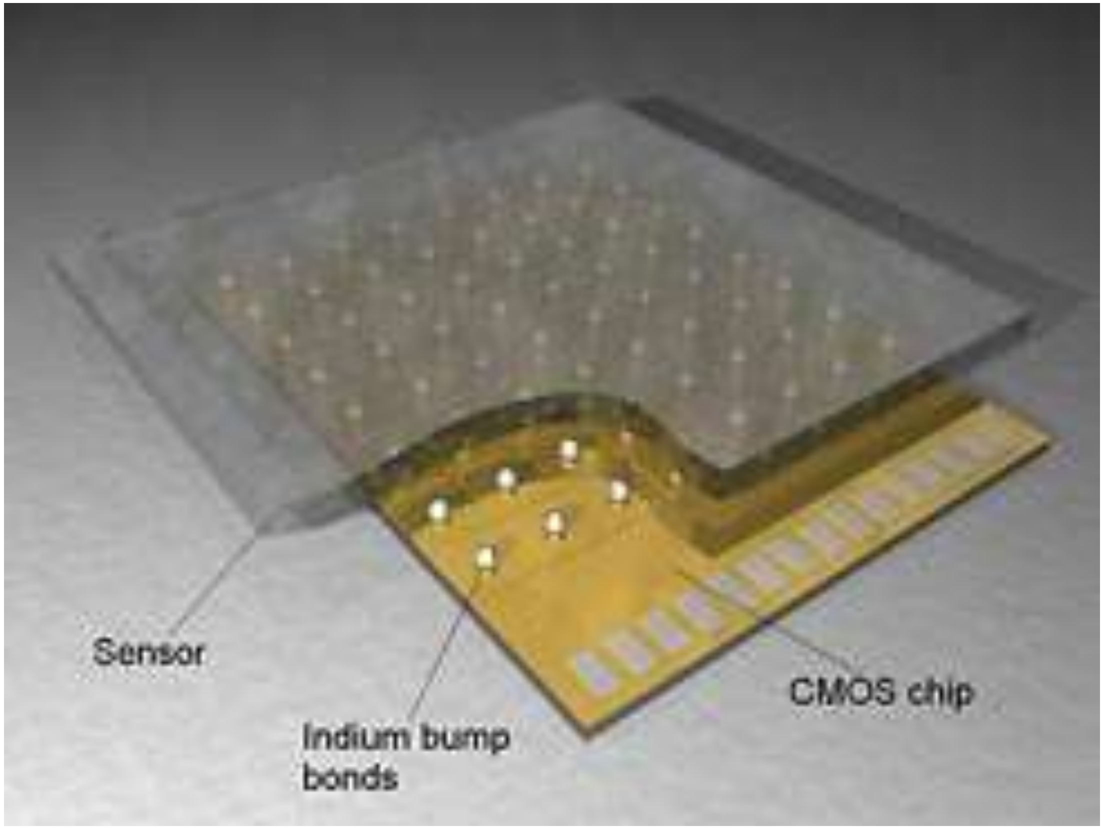


a-Si:H Array dpiX - Flashscan 30





PILATUS 6M Detector



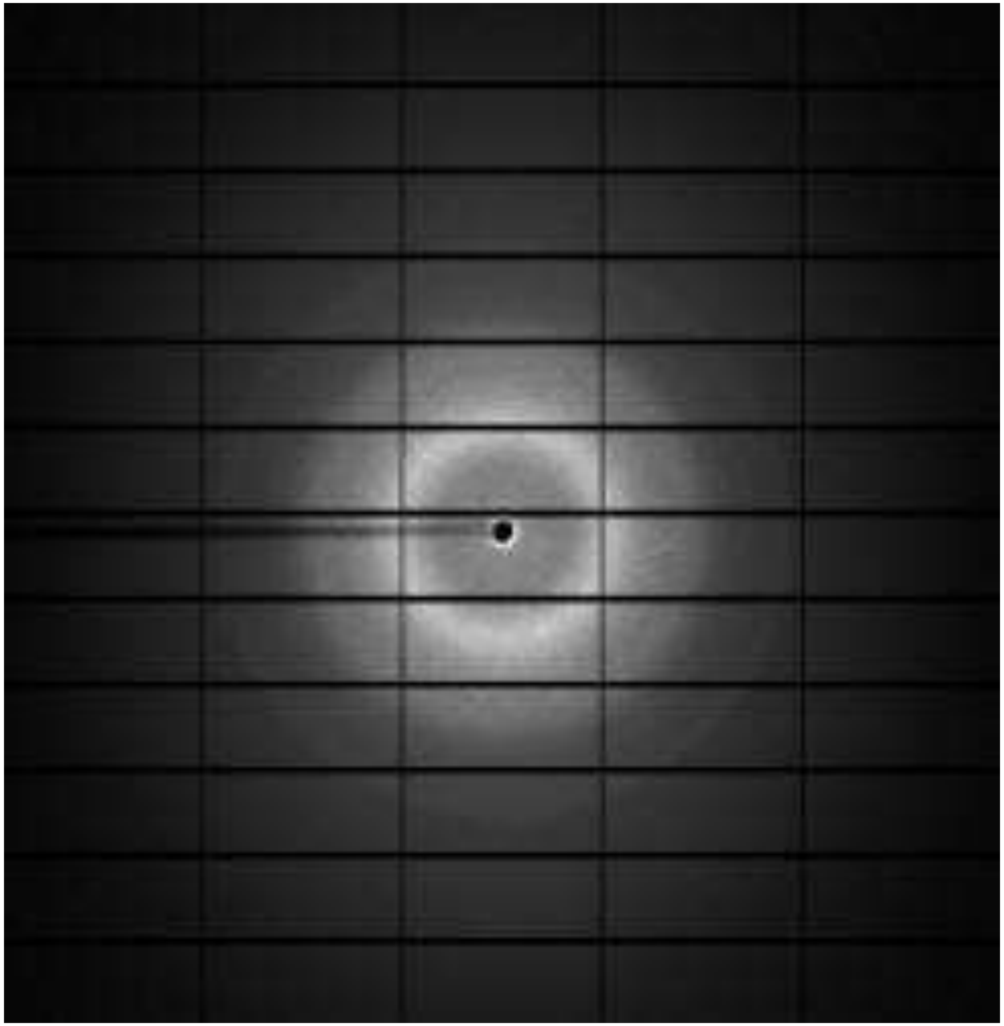
Ch. Brönnimann, E. Eikenberry, B.Schmitt, M. Naef, G. Hülsen (SLS); R. Horisberger, S. Streuli (TEM); Ch. Buehler (LOG); F. Glaus (LMN); M. Horisberger (LNS)

PILATUS 6M Detector

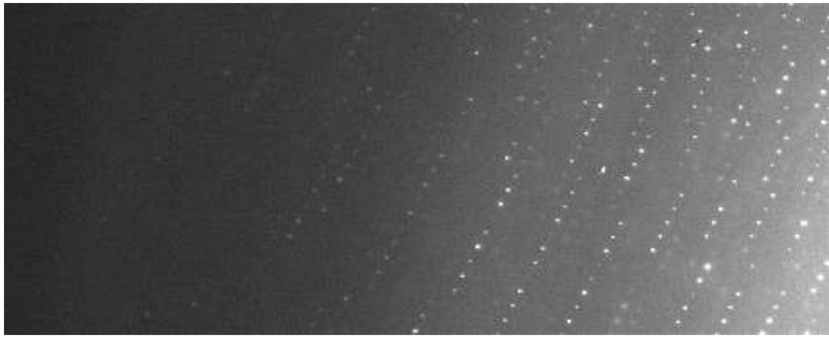
- Sensor 5 x 12 = 60 modules
  - Reverse-biased silicon diode array
  - Thickness 320 µm
  - Pixel size 172 x 172 μm^2
- 2463 x 2527 = 6,224,001 pixels
- Area 431 x 448 mm^2
- Intermodule gap x: 7 pixels, y: 17 pixels, 8.4% of total area
- Dynamic range 20 bits (1:1,048,576)
- Counting rate per pixel > 2 x 10^6 X-ray/s
- Energy range 3 – 30 keV
- Quantum efficiency (calculated)
  - 3 keV: 80%
    8 keV: 99%
    15 keV: 55%
- Energy resolution 500 eV
- Adjustable threshold range 2 – 20 keV
  Threshold dispersion 50 eV
- Readout time 3.6 ms
- Framing rate 12 Hz
- Point-spread function 1 pixel

PILATUS 6M Detector

DECTRIS
Next Generation X-Ray Detectors



- X-ray diffraction image recorded from a ferritin crystal (energy=16 keV, distance = 204 mm).



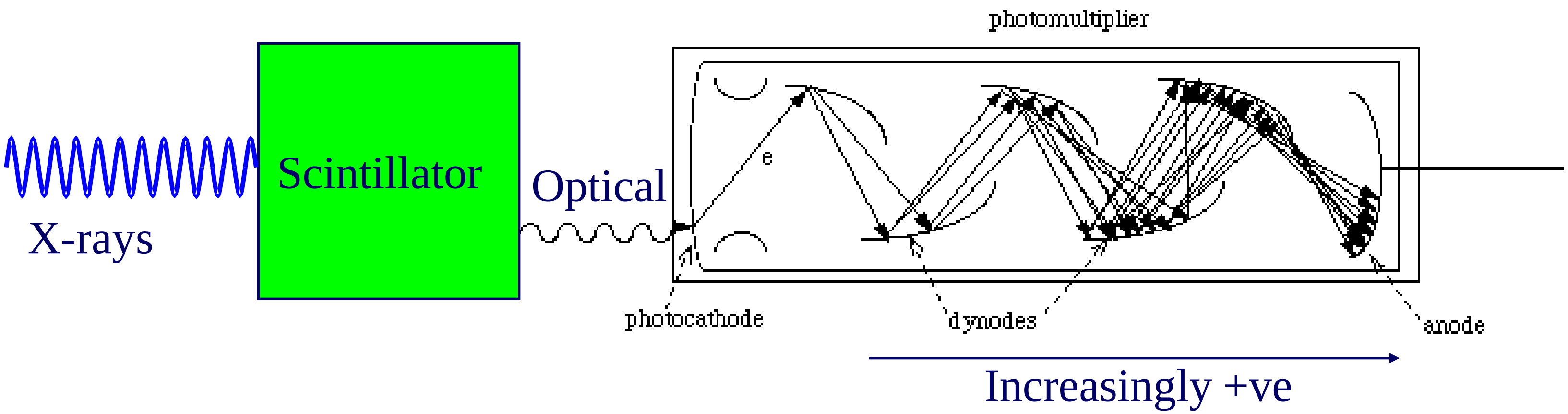
Spectroscopic Detectors

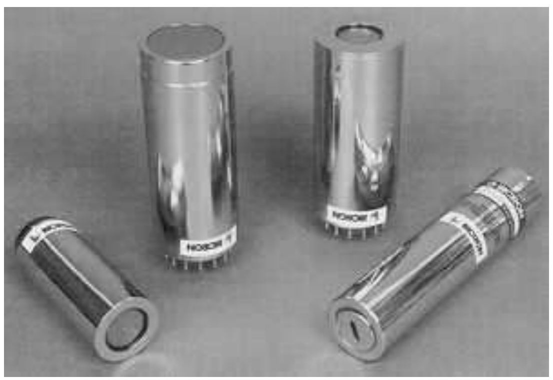
Rainbow Lorikeets

Spectroscopic Detectors

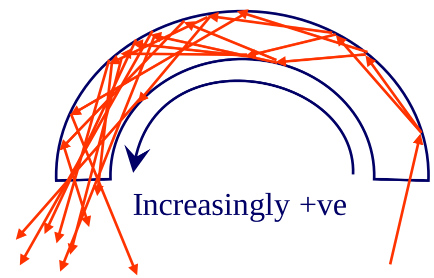
- For quantitative work, most are counting detectors that measure the size of individual energy deposits
- Alternative ids the use of filters as in optical colour cameras

Electron multipliers & Scintillators





Channeltron is a similar with distributed dynode



Micro-channel plates are mutlichannel channeltrons with each channel being an electron multiplier.

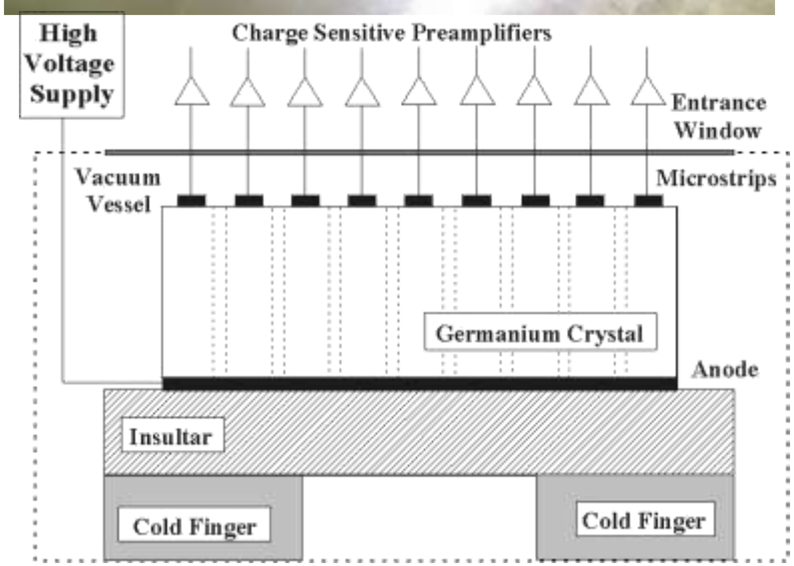
Multi Channel Spectoscopic Detectors

Canberra Ultra-LEGe detector

WRULEAD (Windowless, Retractable, Ultra Low Energy Array Detector) works down to 300eV

Multichannel devices up to 30 channels at 3×10^5 cts s^{-1} channel^{-1} have been built

SPring-8 128 channel Ge strip



- Ge
  - 55.5×50.5×6mm
- Strips
  - Number 128
  - Width 300µm
  - Interstrip 50µm
  - Length 5mm
- Readout
  - Single channel 100ns
  - 32 channels 3.2ms
- Max expected count rate
  - 14kcps

Spectral Resolution

- Average number of carriers, $N = E/w$

  where w is energy to create electron hole/ion pair
- Poisson statistics $\sigma = 1/\sqrt{N}$

  $= (E/w)^{-1/2} = (w/E)^{1/2}$
- $\Delta E/E$ fwhm $= 2.355\sigma$

  $= 2.355(w/E)^{1/2}$

- For Ge, w= 3eV so at 10keV $\Delta E/E \sim 4\%$
- For NaI, w= 30eV so at 10keV $\Delta E/E \sim 13\%$

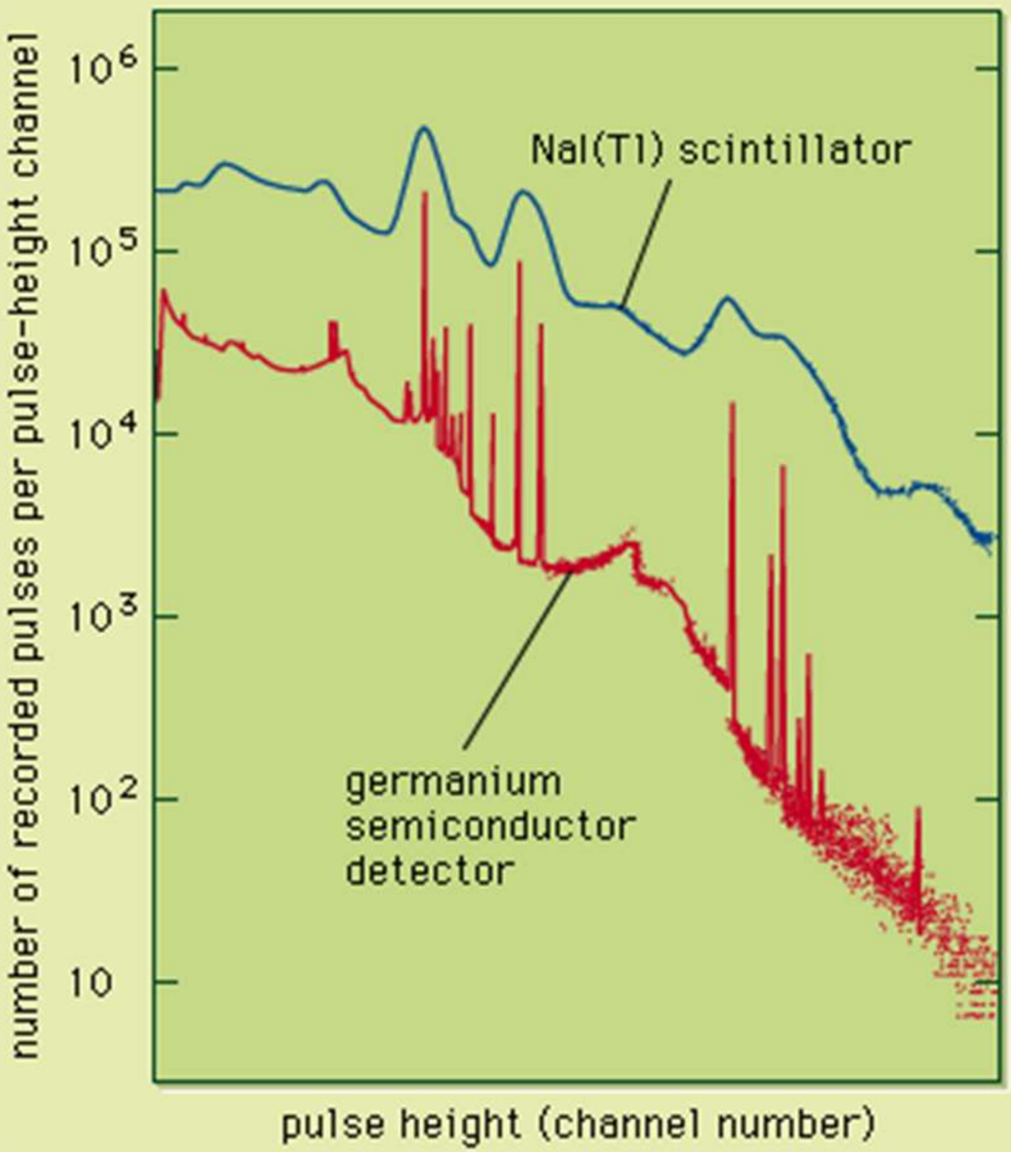
Fano Factor

- If all energy from photon or particle were converted into carriers there would be no variance
- Poisson statistics assume only a small fraction of energy goes into charge creation
- Reality is somewhere in between so we introduce Fano factor F
- Fano factor is defined as

$$F = \frac{\sigma^2}{\mu}$$

  where σ^2 is the variance and μ is the mean number of carriers
- For a Poisson process, the variance equals the mean, so $F = 1$
- Examples
  - Si: 0.115
    Ge: 0.13
    GaAs: 0.10
    Diamond: 0.08
- Observed relative variance = F x Poisson relative variance

Scintillator vs Germanium



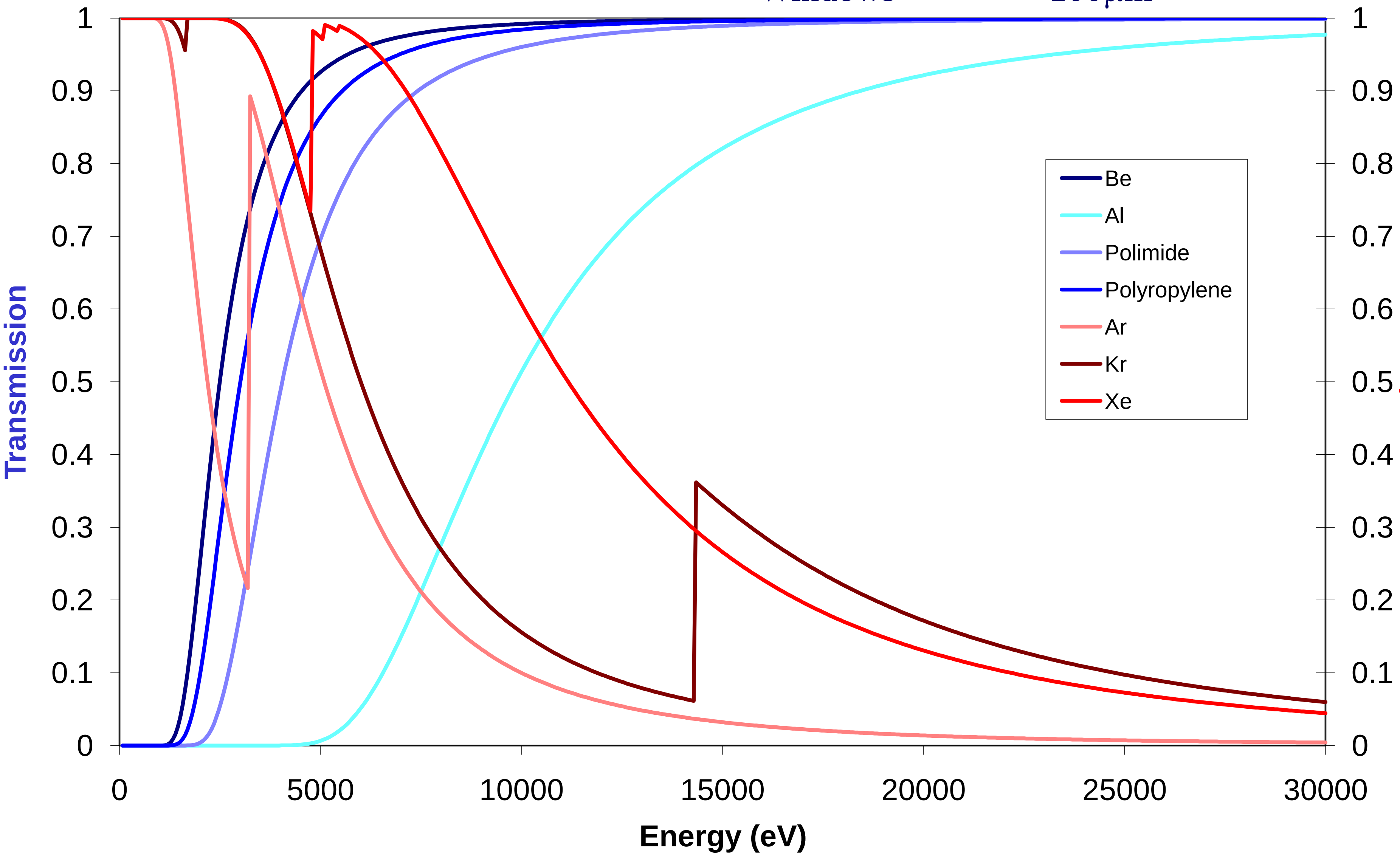
The top spectrum is from a scintillation detector, and the bottom is from a germanium semiconductor detector. The superior energy resolution of the germanium is evident from the much narrower peaks, allowing separation of gamma-ray energies that are unresolved in the scintillator spectrum.

Things to Look Out For

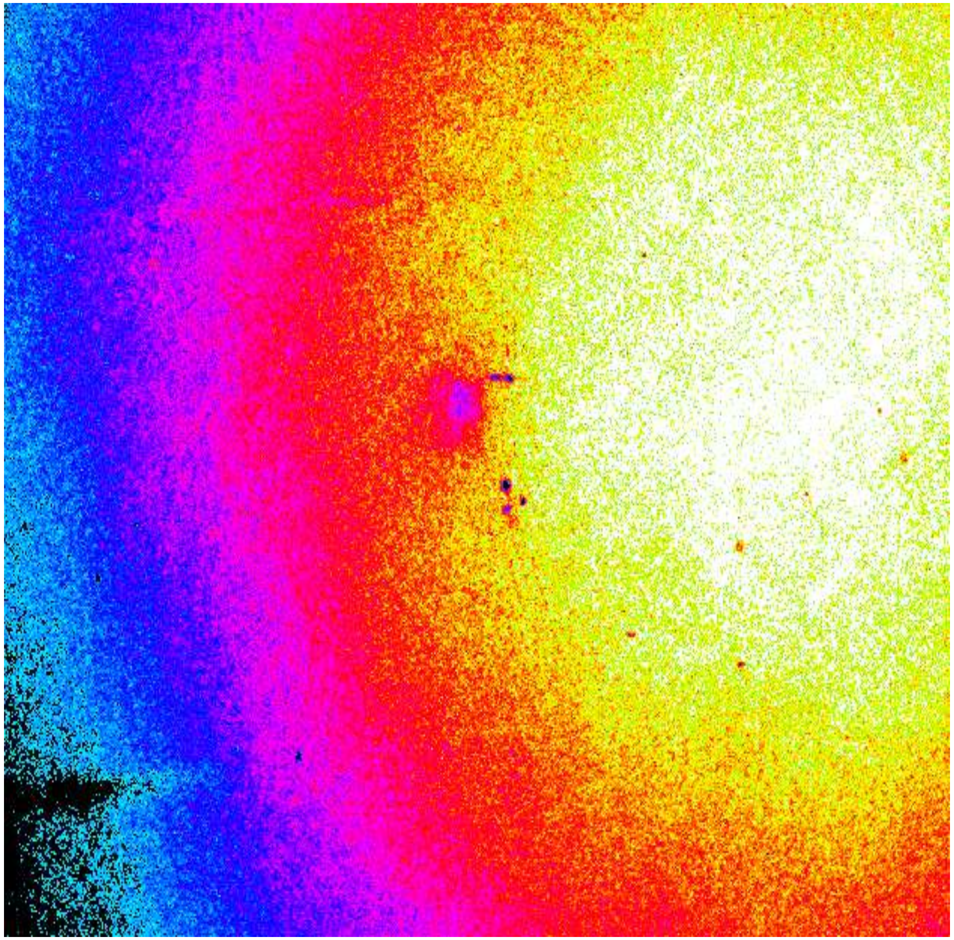
Crocodile

Efficiencies

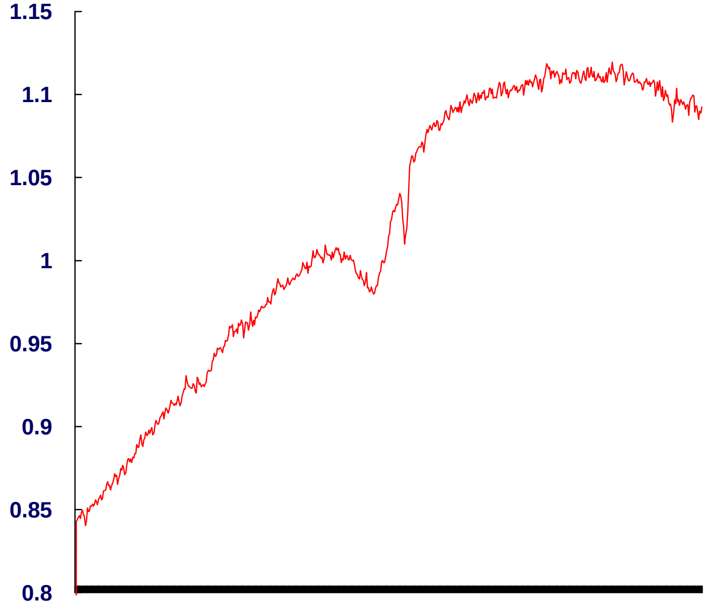
Gas thickness 1cm
Windows 100µm



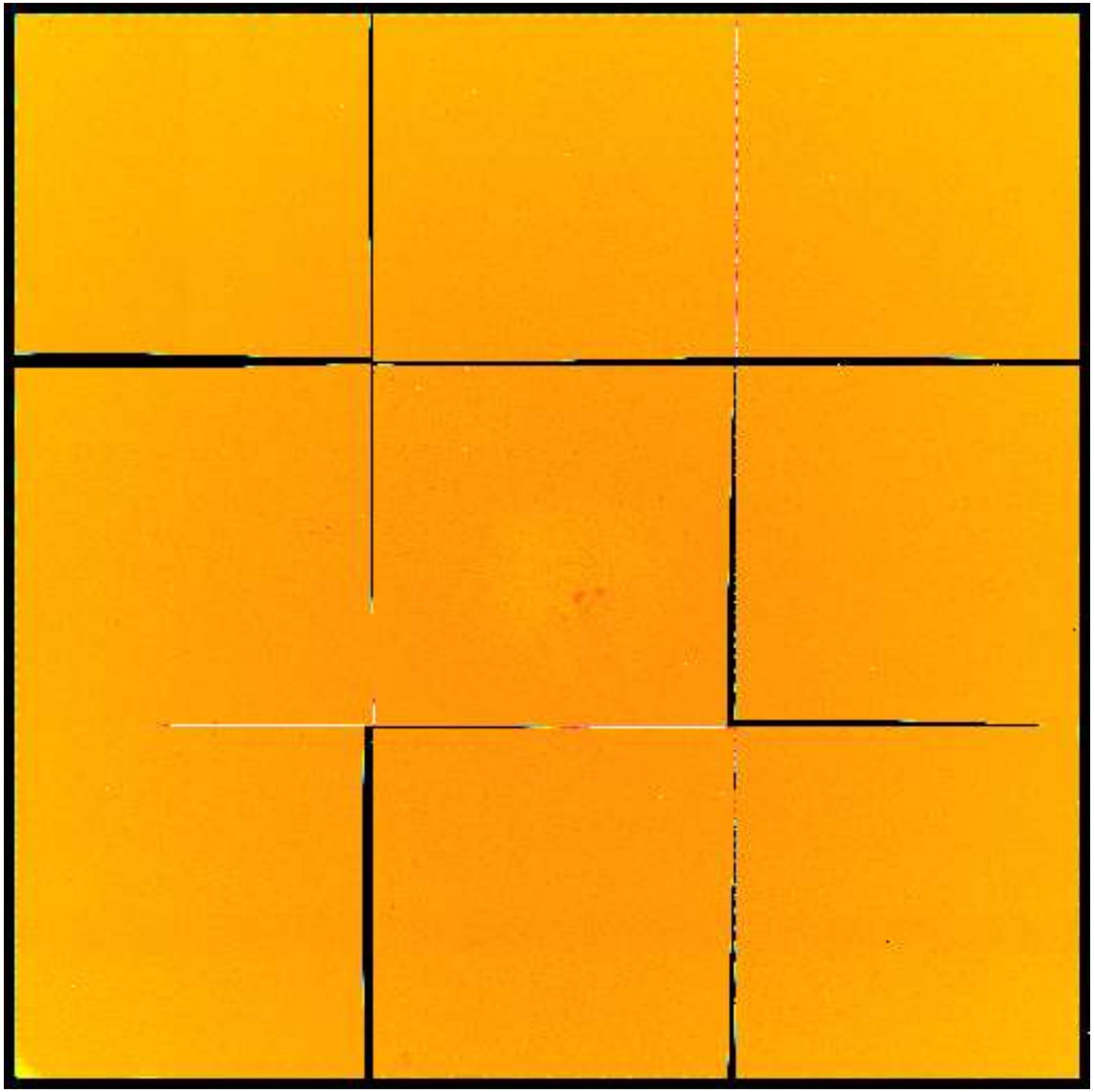
Response to Uniform Illumination



ESRF TV Detector
Thompson IIT & CCD



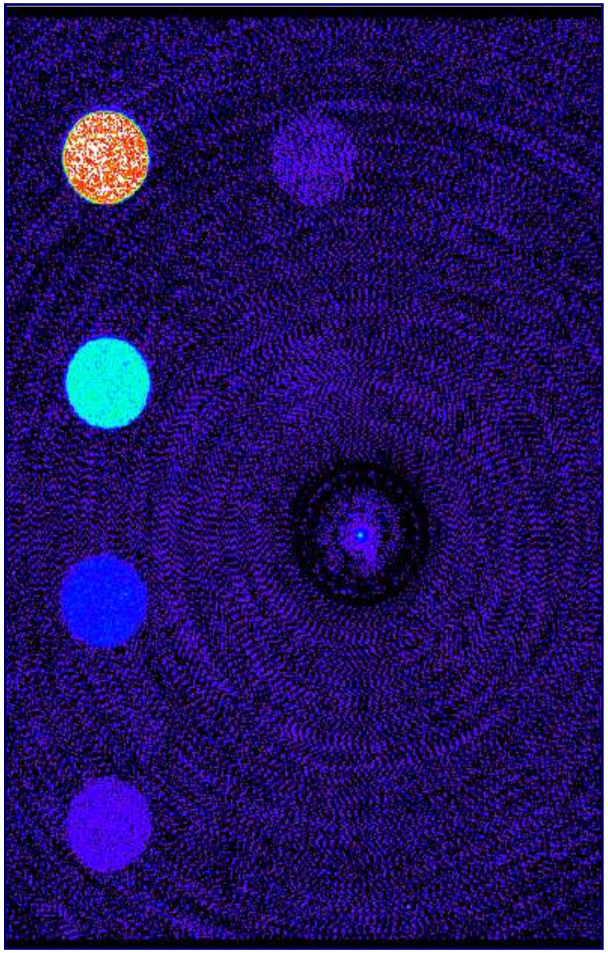
Gaps



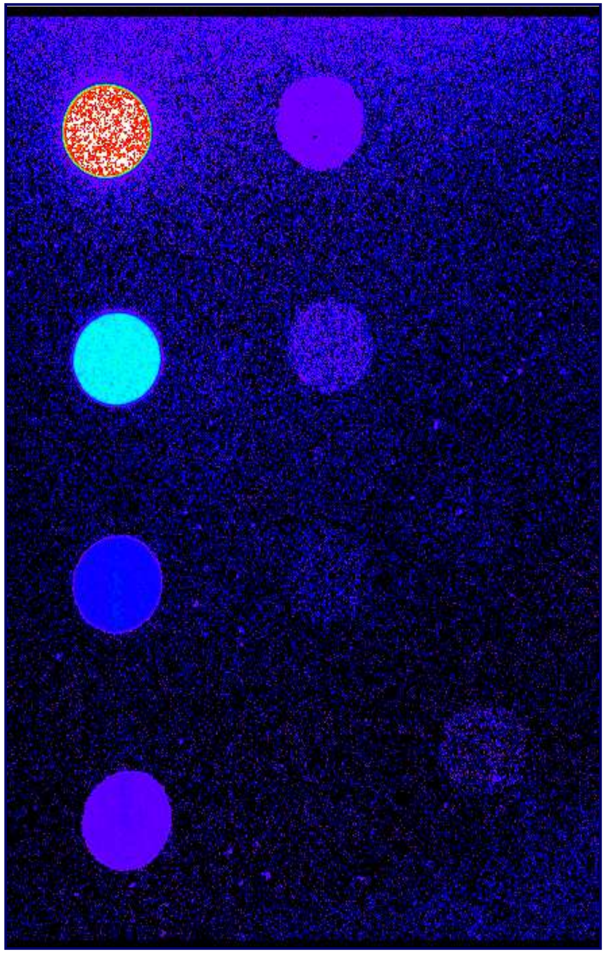
Spec	**0.2mm max**
Worst gap	**2.97mm**
Pixels in gaps	**513922** **5.45%**

Graded Absorber Comparison

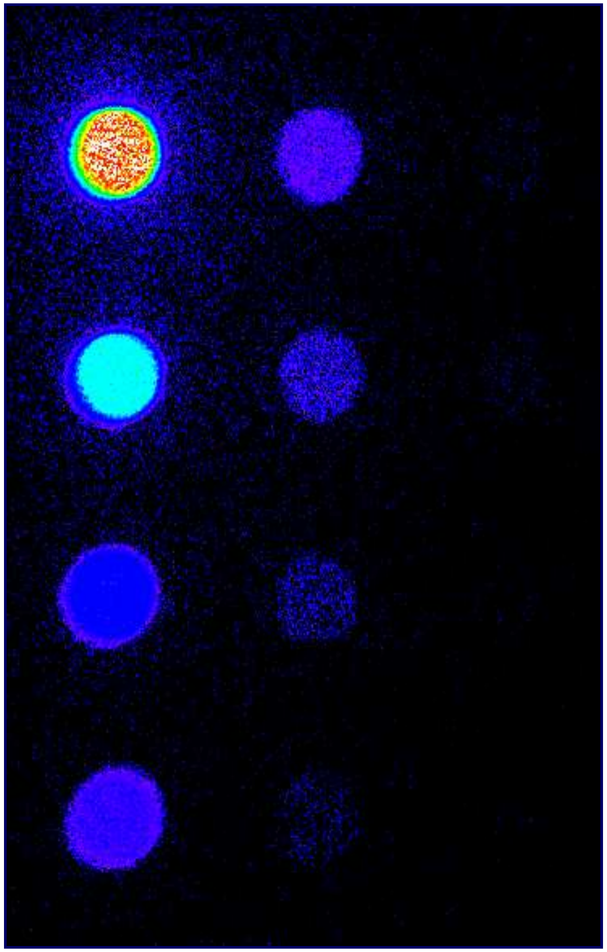
Mar Image Plate



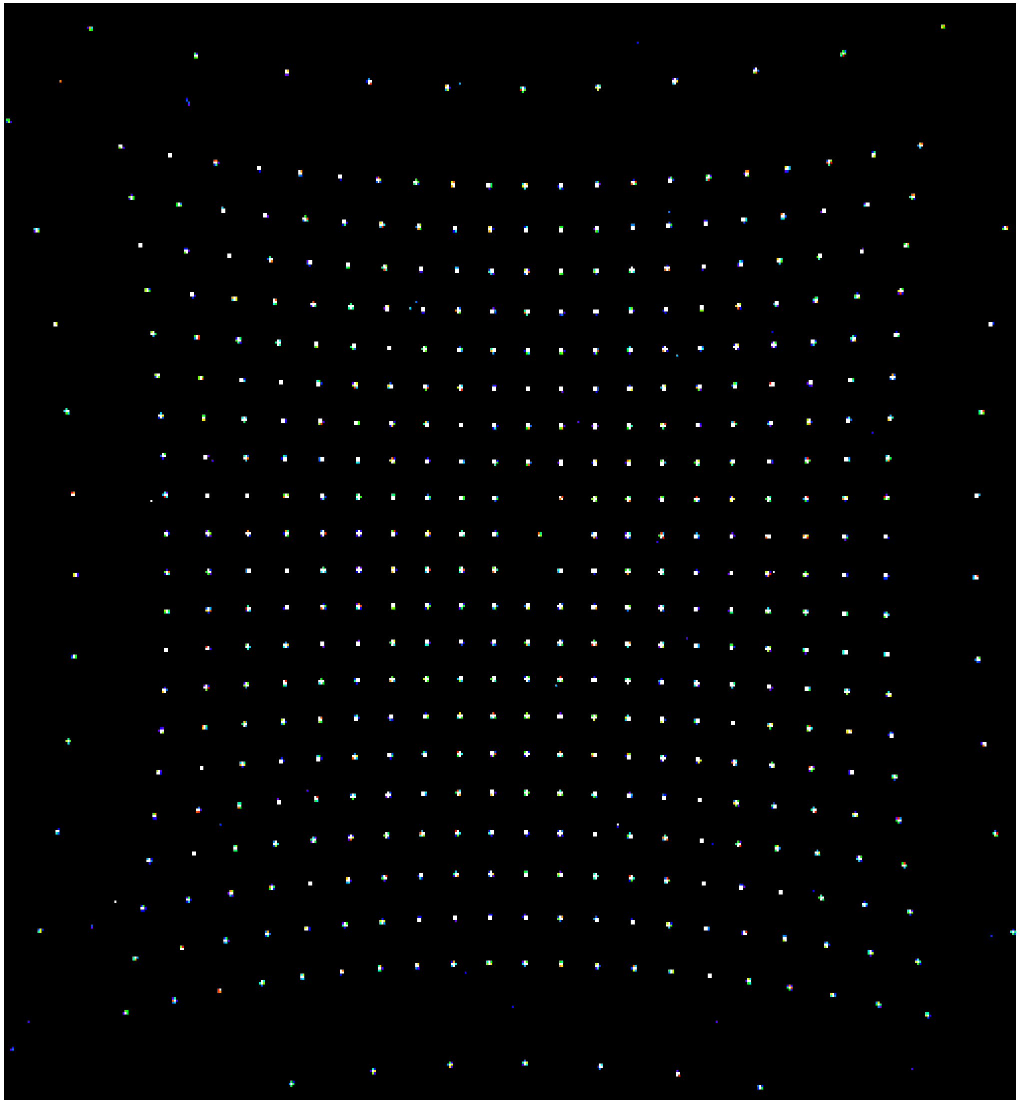
ESRF-Thompson IIT / CCD



Daresbury MWPC

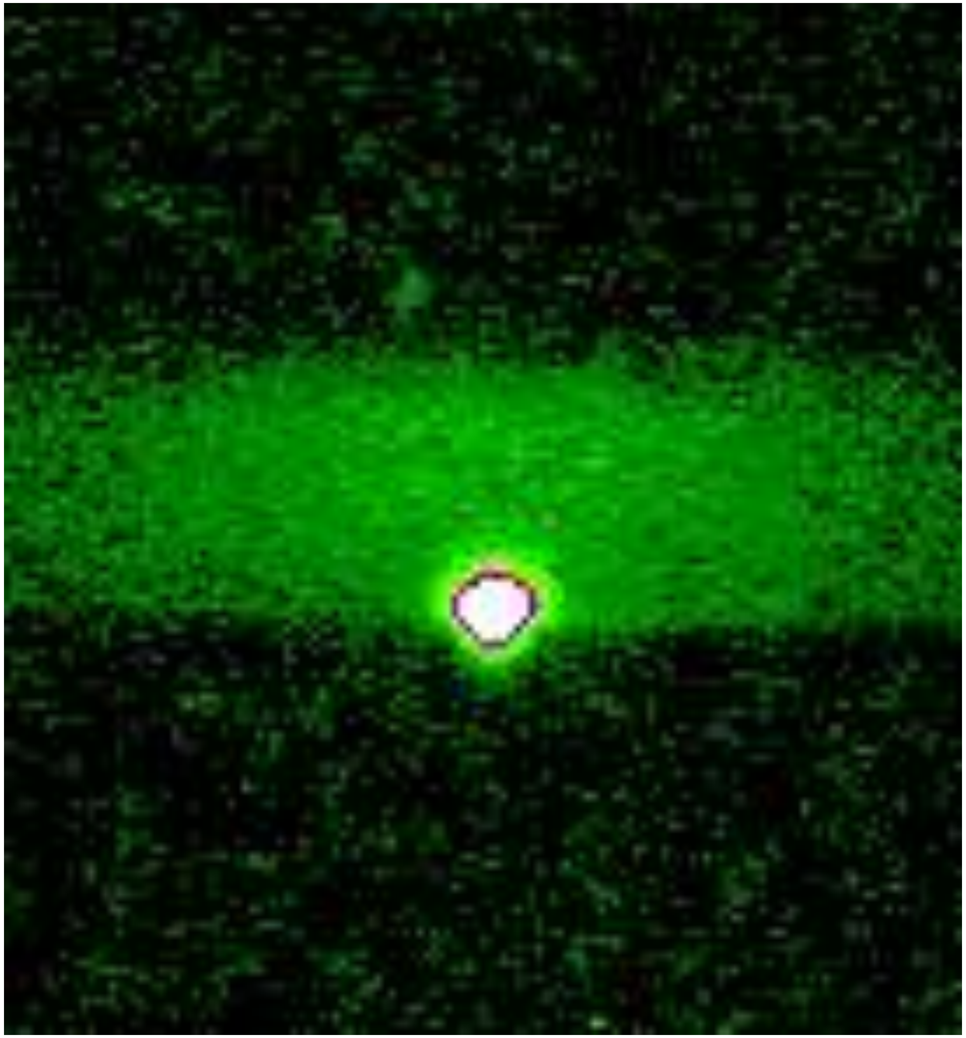


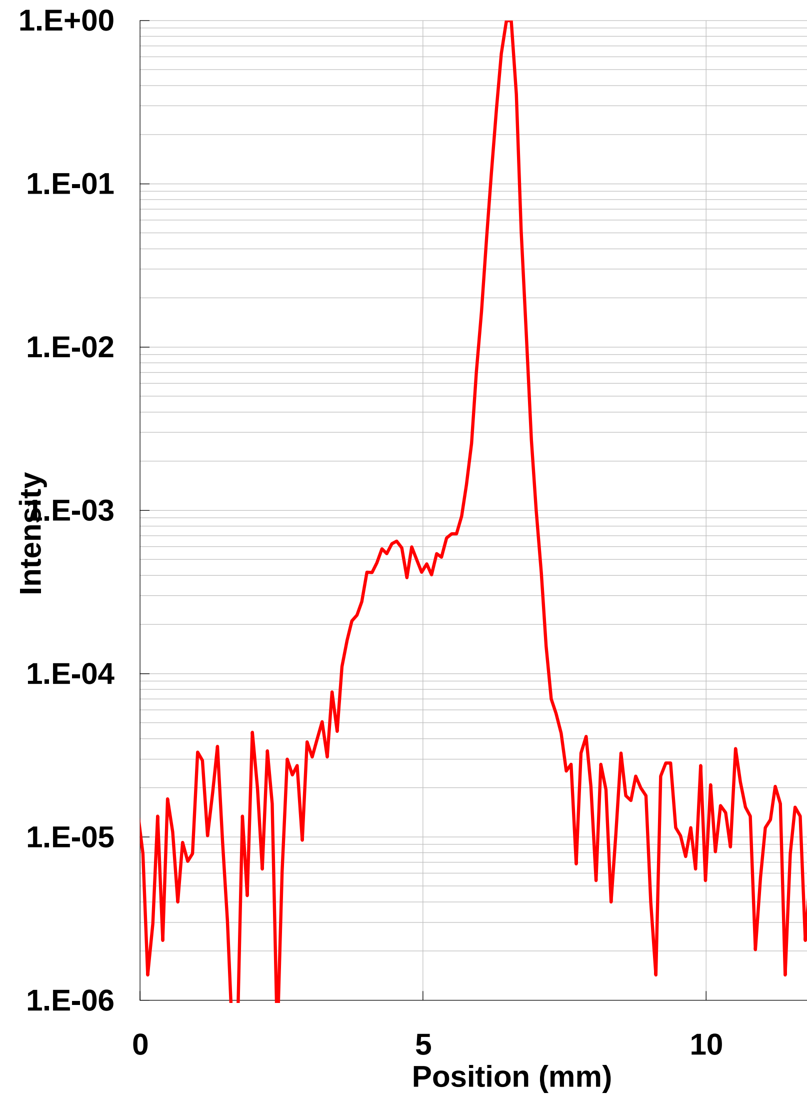
Spatial distortion



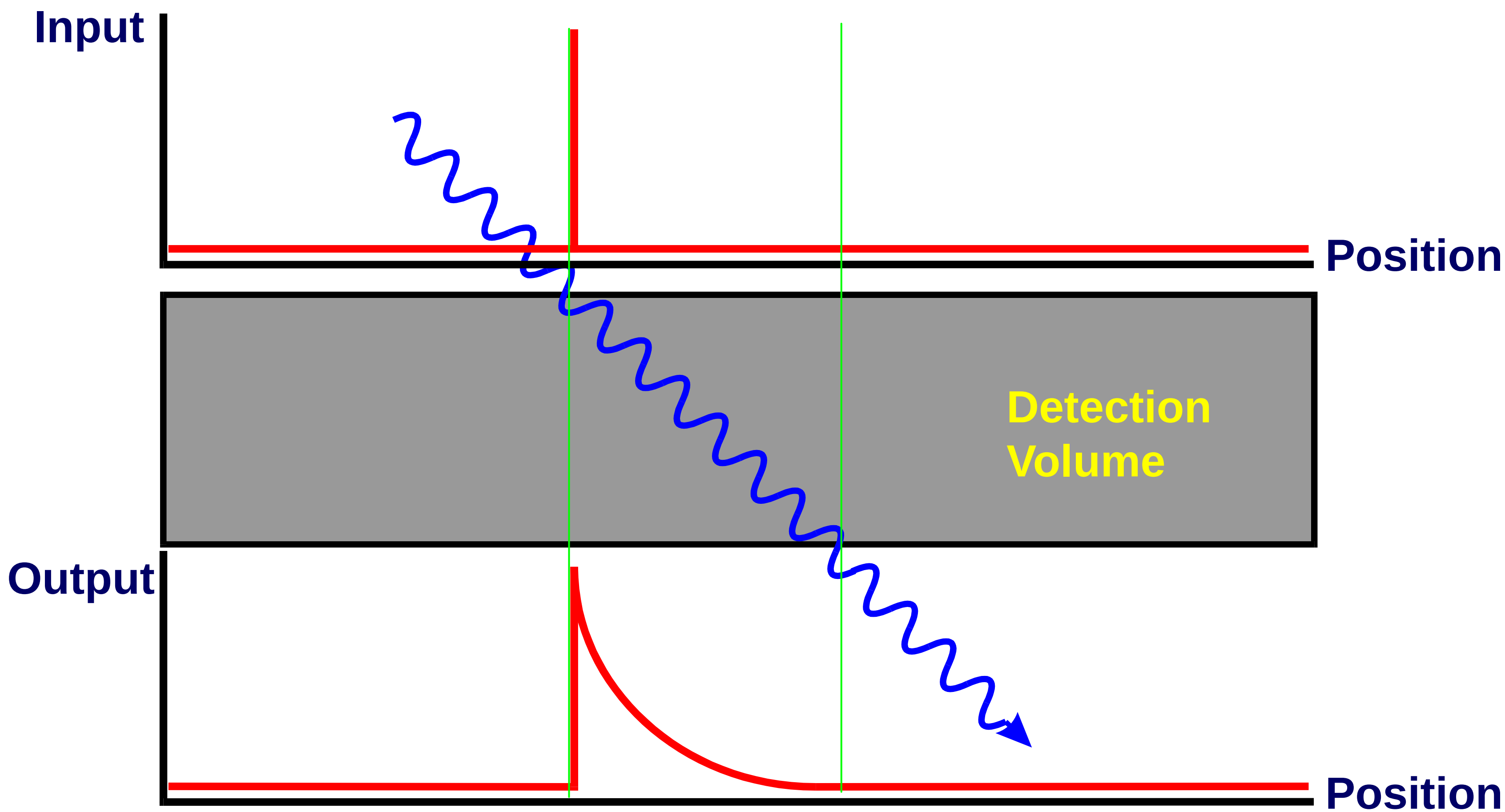
ESRF Image intensifier detector

IPlate Single Peak PSF



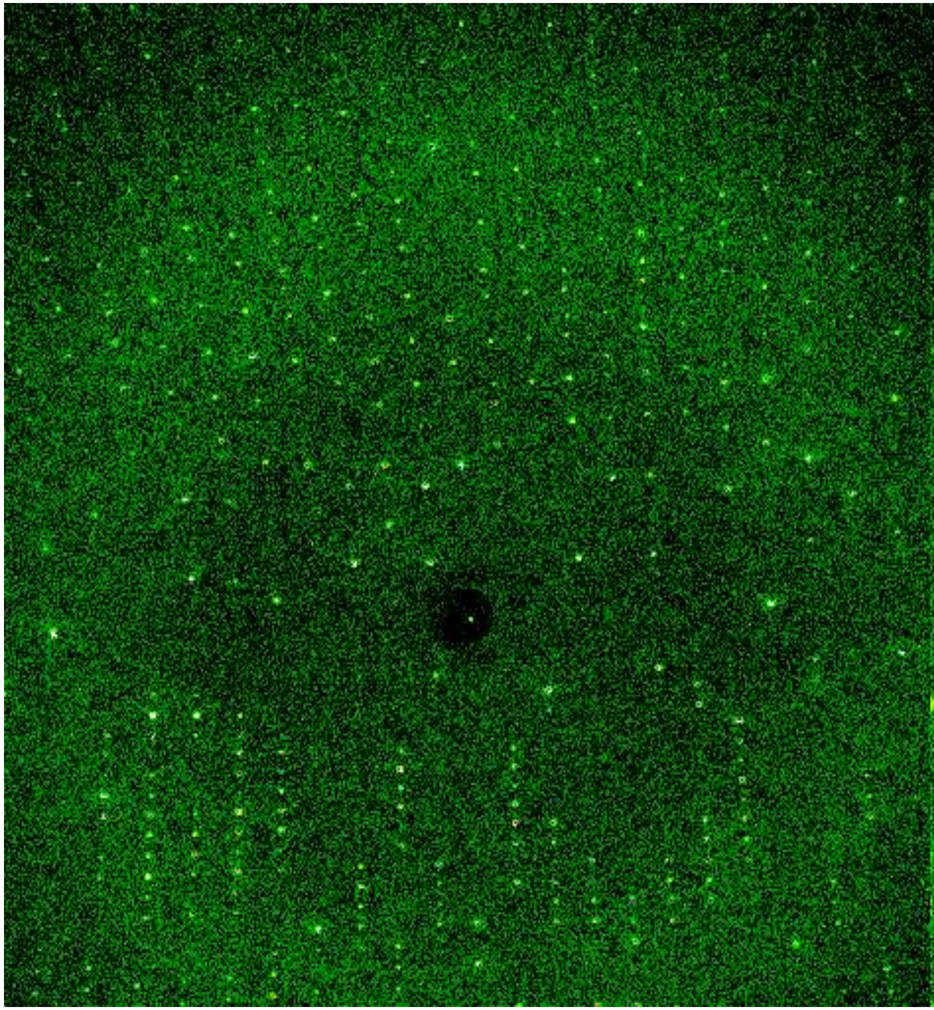


Parallax Broadening

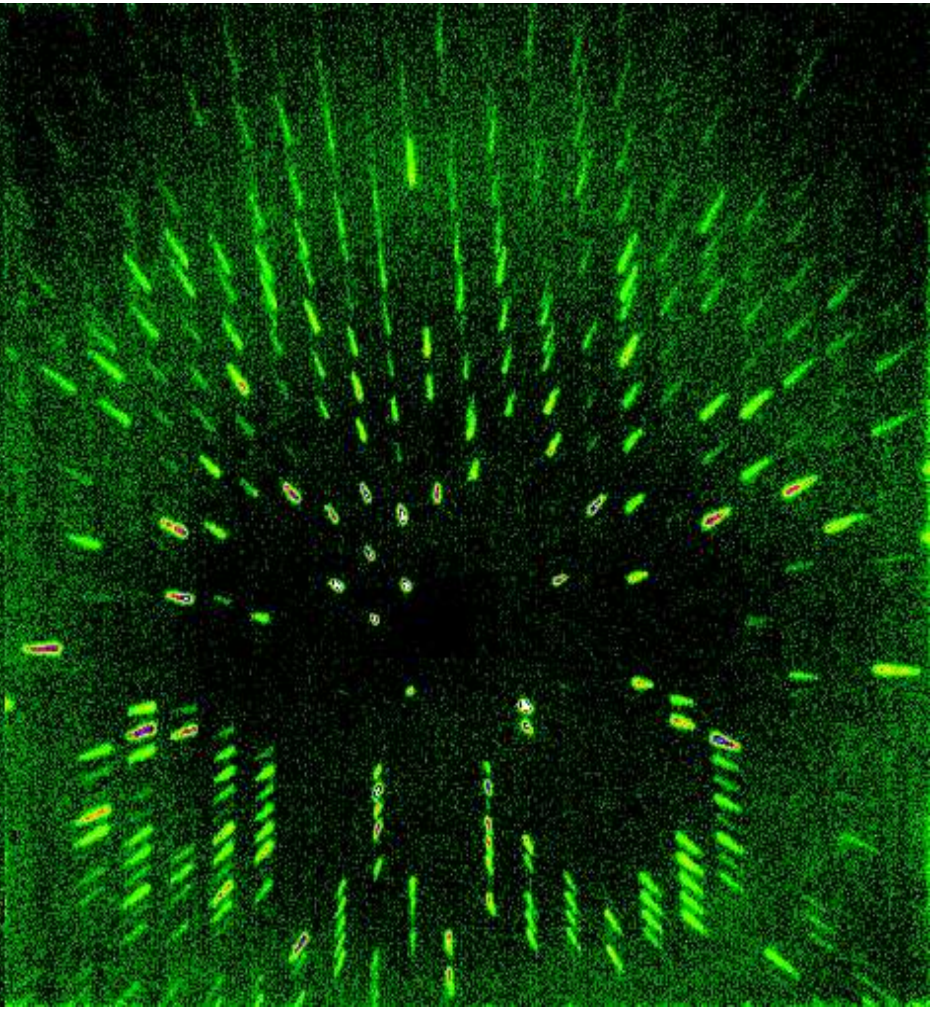


Parallax Effect

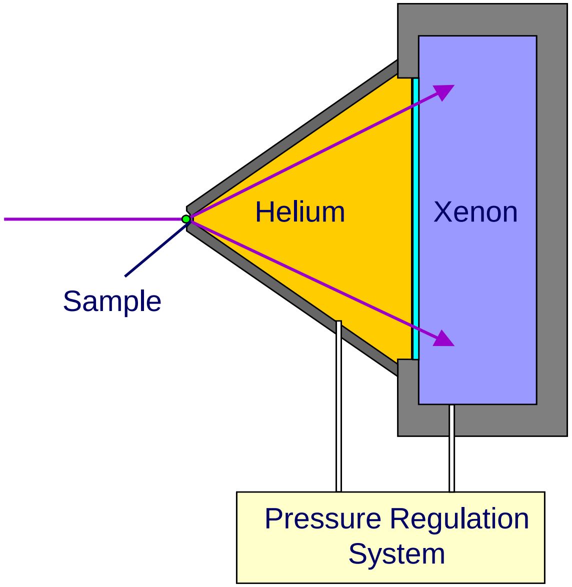
Image Plate



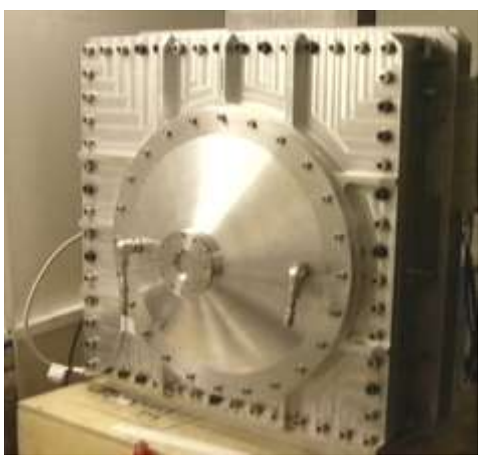
Gas Proportional Counter

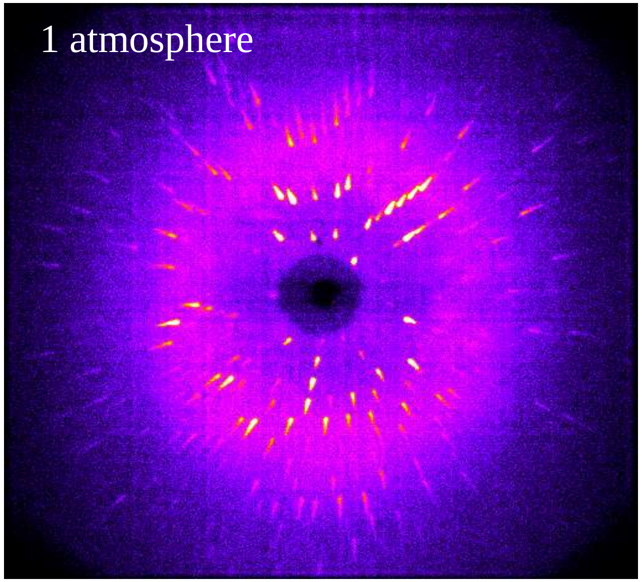


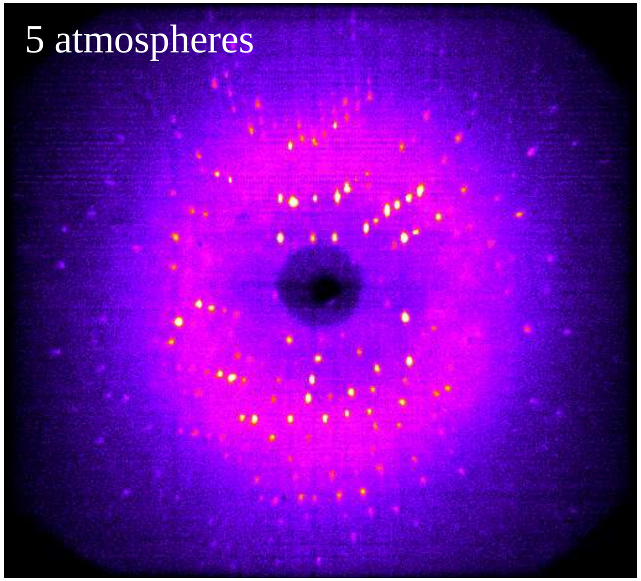
Daresbury High Pressure MWPC



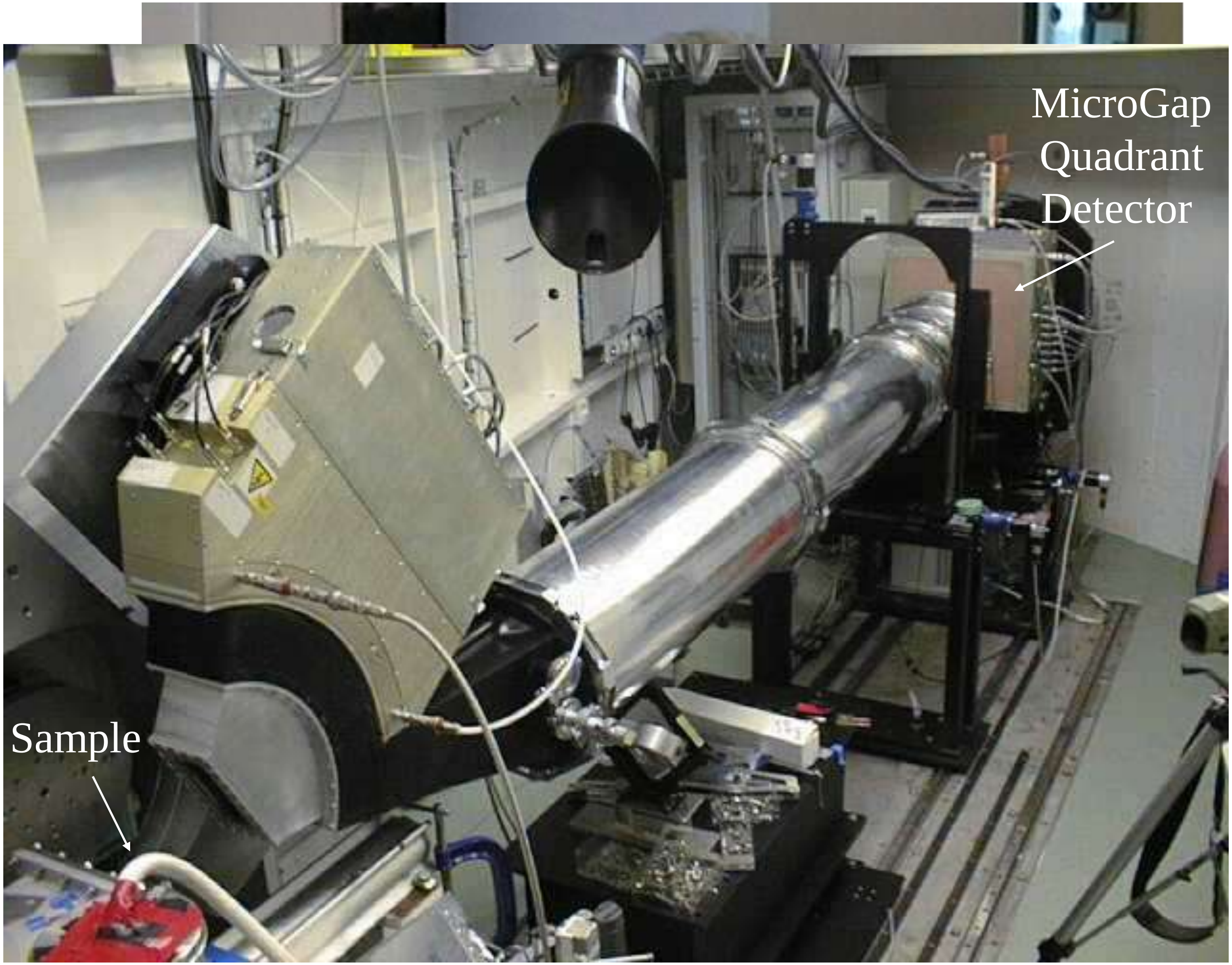
Force on 28 x 28 cm window at 5 bar = 4 tonnes
Force on window of 1 x 1 cm at 5 bar = 5 kg



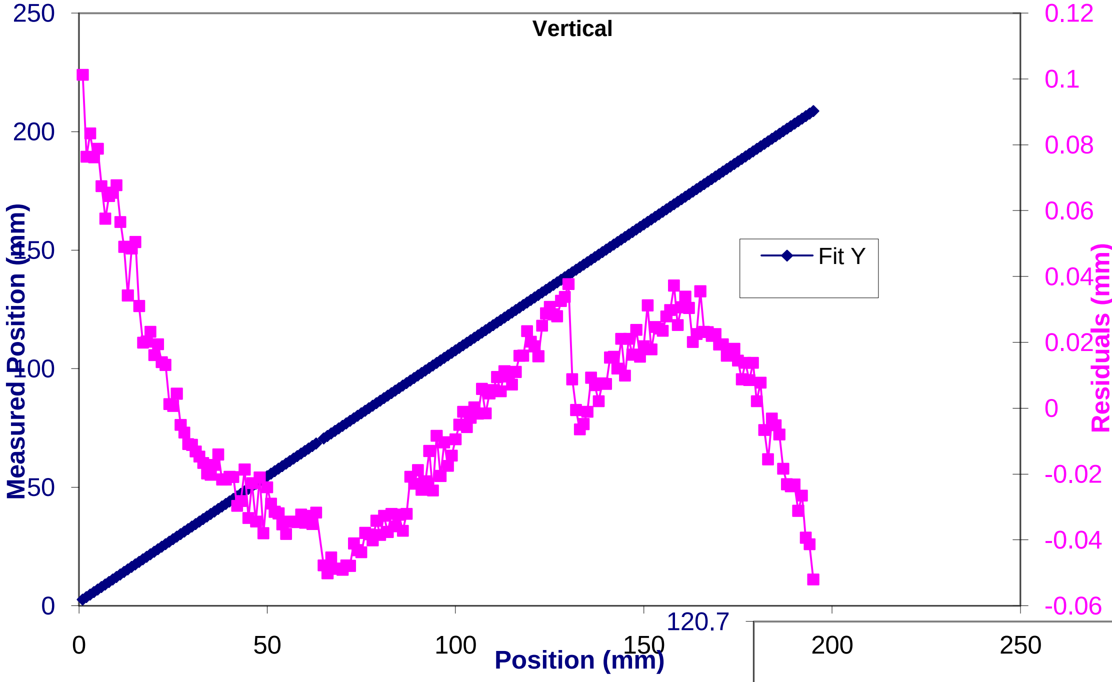


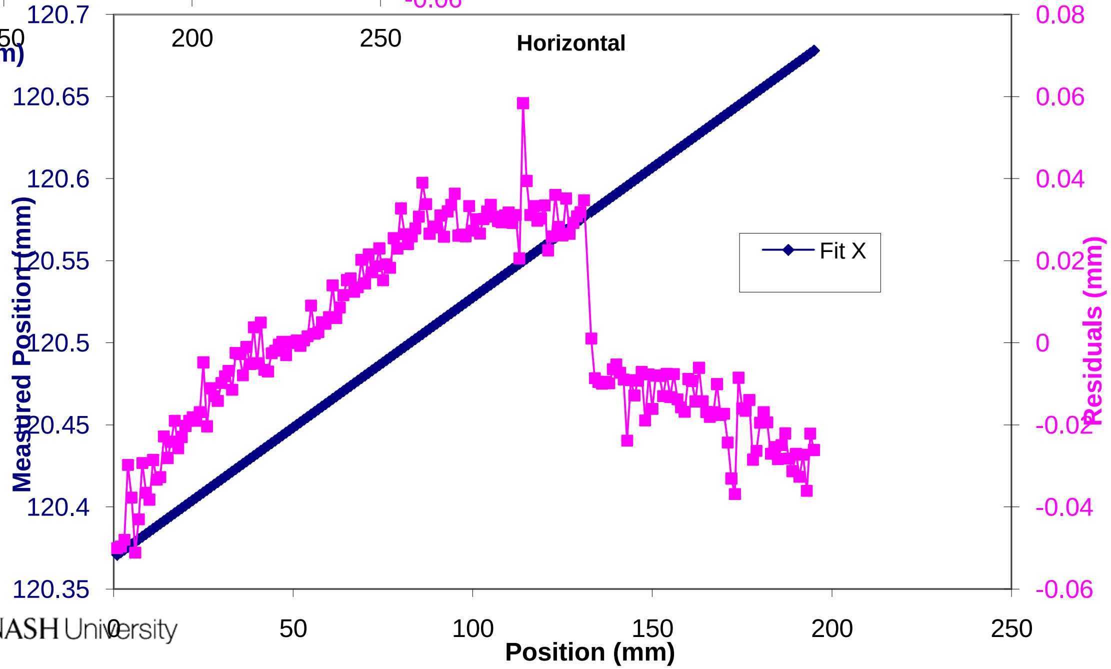


RAPID2 SAX WAX

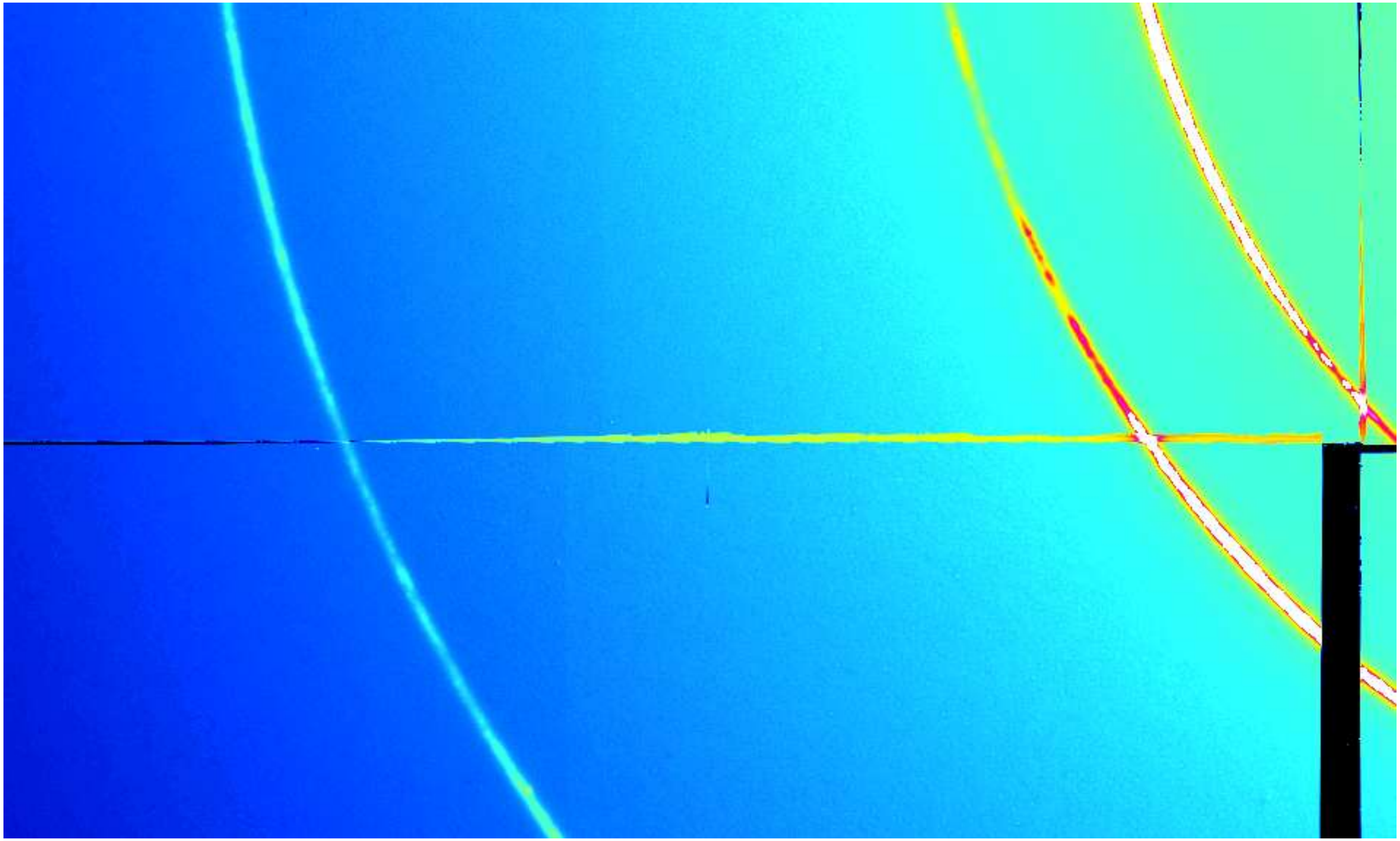


Geometric Distortion

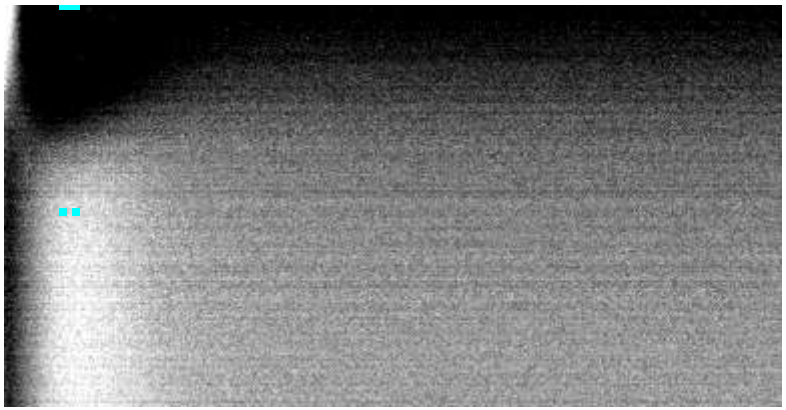


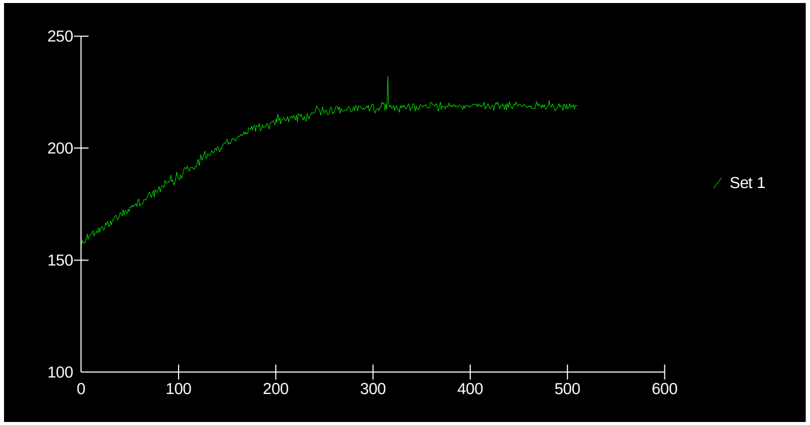


Overlaps



Dark Currents





Flat and Dark Correction

For each image, two correction images must be recorded.

1. A flat field (uniform illumination of the detector)

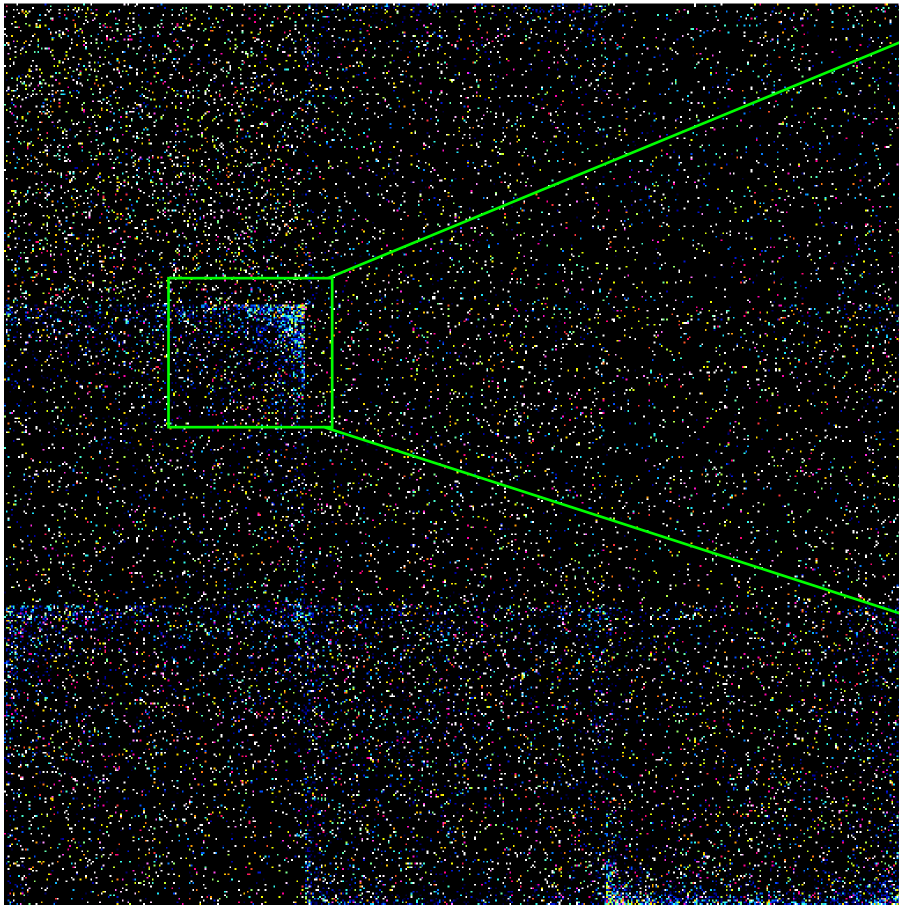
2. A dark image (no irradiation of detector)

Both must be recorded with the same exposure time as the original image since dark current is a function of exposure time.

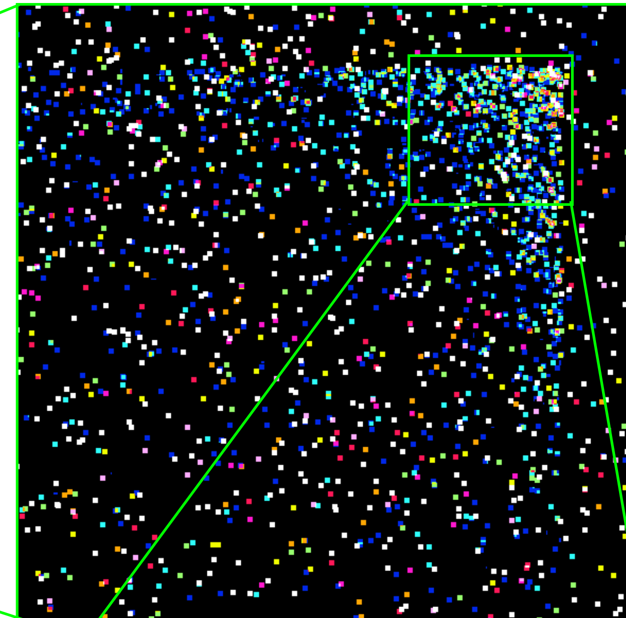
Then apply the following correction

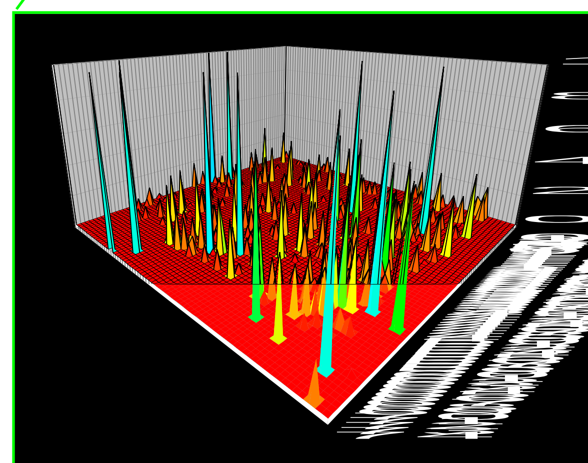
$$Corrected = \frac{(image - dark)}{(flat - dark)}$$

Dark Current



Pixels above the 0.2 photons pix^{-1} specification



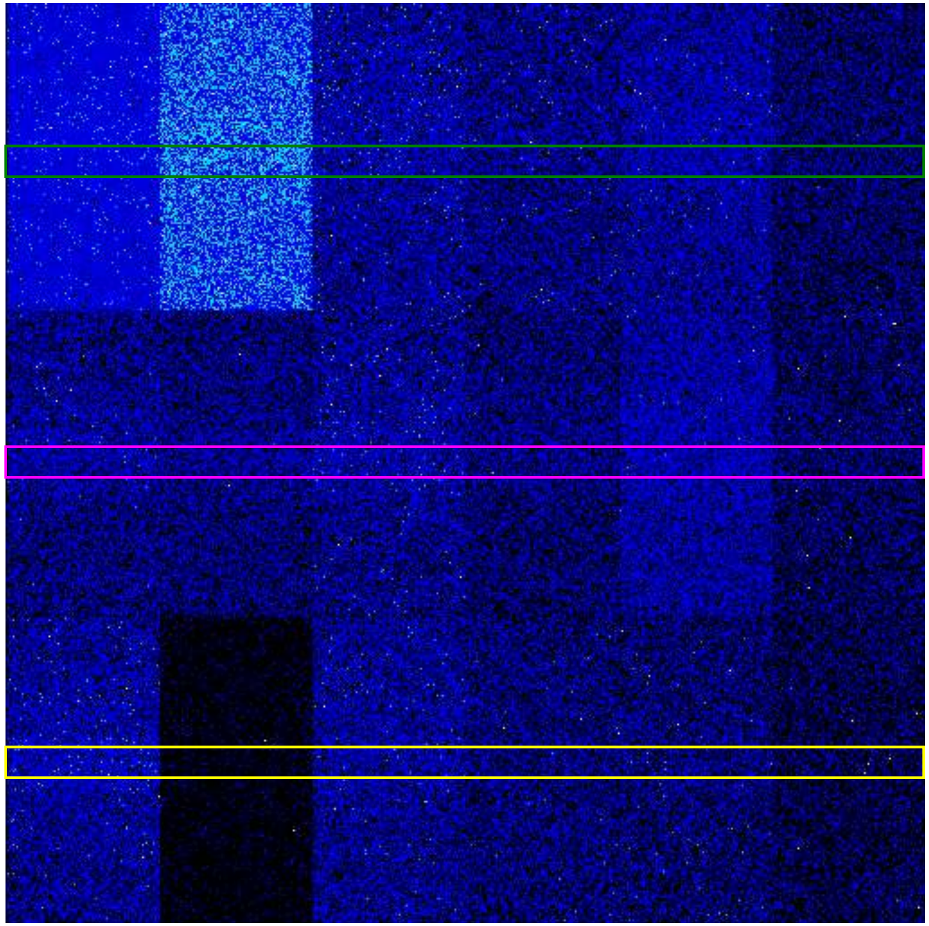


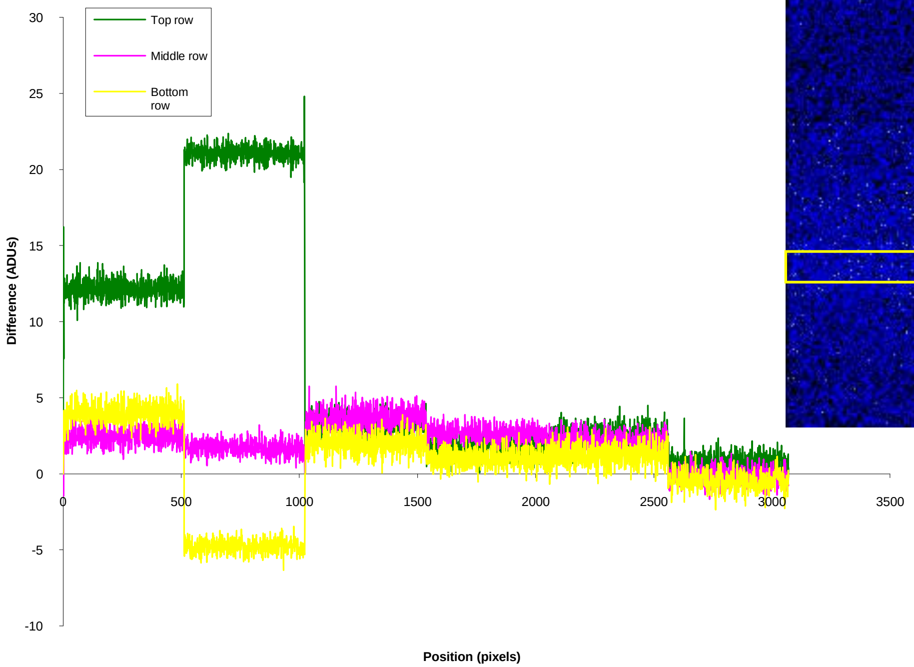
Number failing 2 measurements 5-2000s

Mean	44764	0.47%
Min	40822	0.43%
Max	48706	0.52%

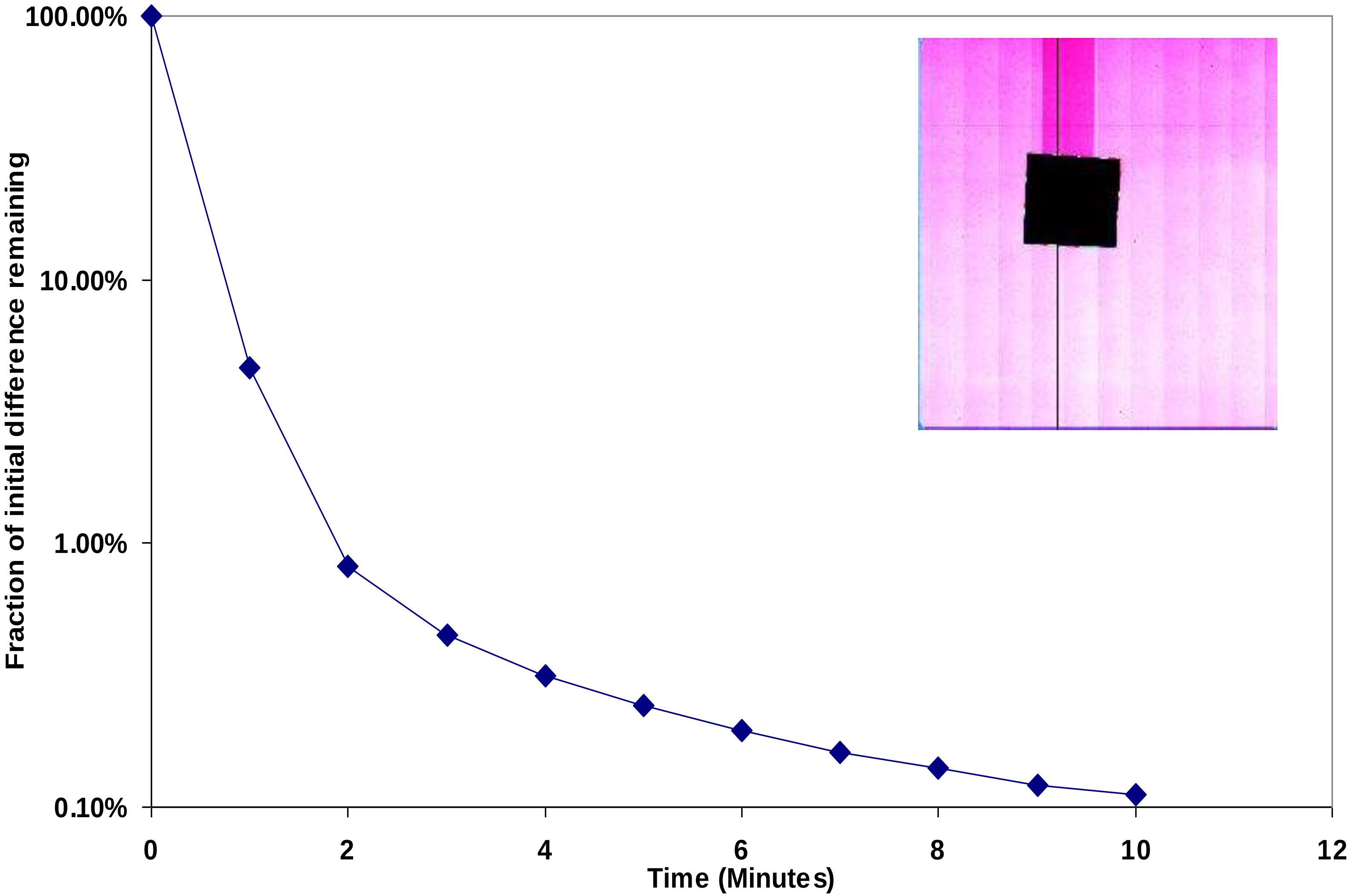
nb. 14300 pixels not common to both

Subtraction of dark images





Flashscan 30 - Image Lag



Radiation Damage (Medipix)

- Damage occurred at 40Gy or 1.3×10^{10}pht/mm^2 in the readout chip
- At 13 keV photon energy
  - Strong diffraction spots typically 10^5 phts/s or 10^6 phts/mm2/s
    - Damage requires ~ 8hours exposure
  - Direct beam (10^{10}–10^{13} photons/mm^2/s)
    - Damage in less than a second.

dpiX Flashscan 30 PaxScan 4030

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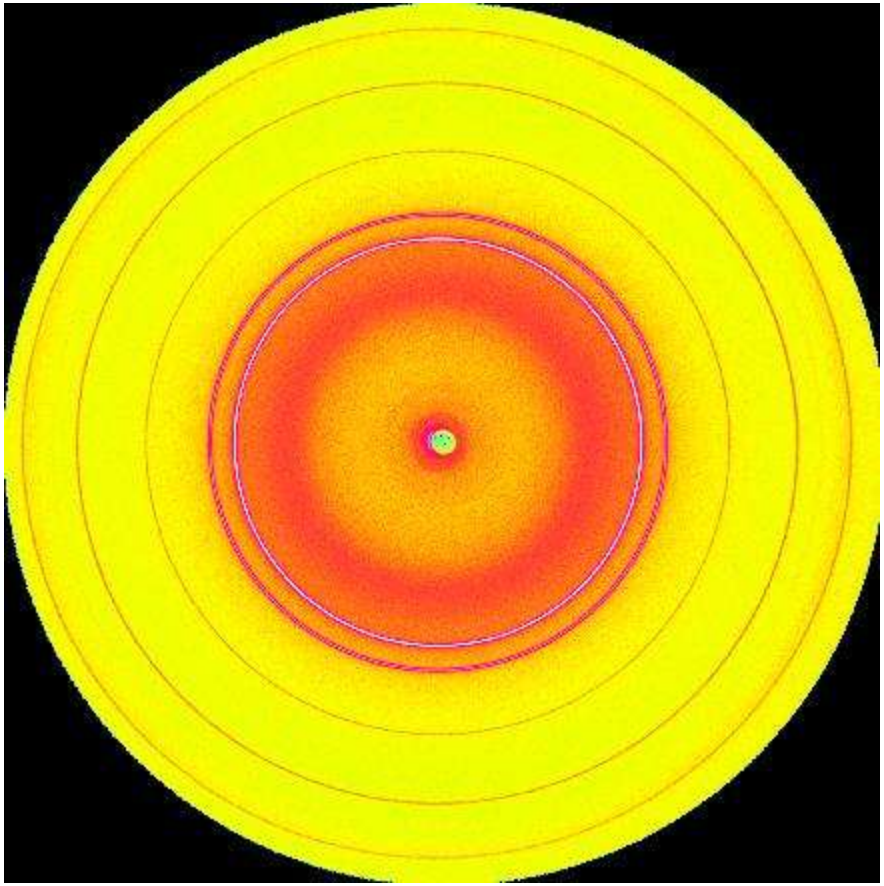
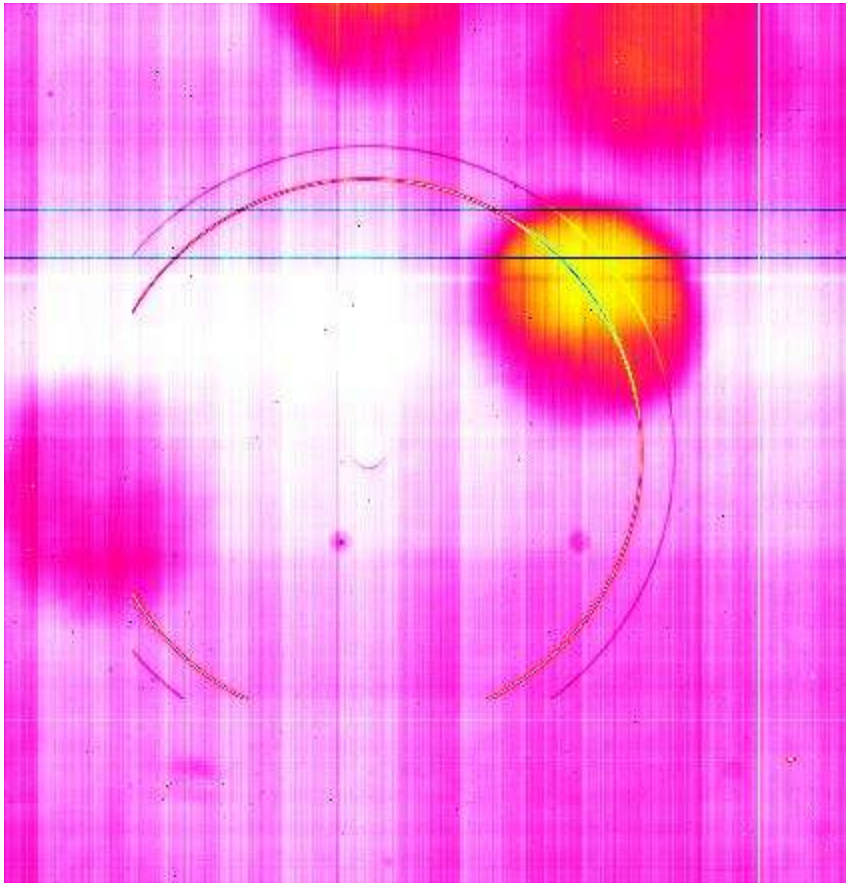
Flashscan 30 - Performance

Mar Image Plate



t_{int}=30s

Flashscan-30



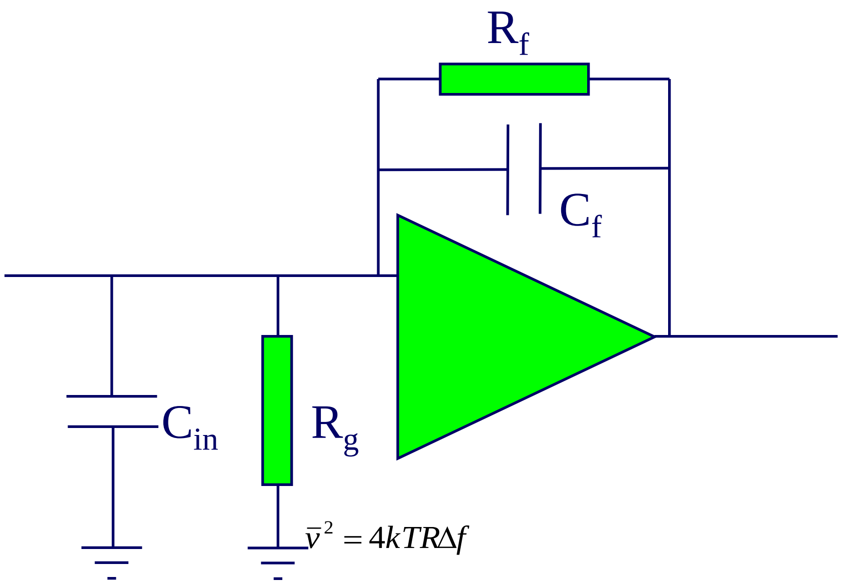
t_{int}=190s

Electronics Issues

Koalas

Albino Kookaburra

Amplification



- Voltage mode
  - Output ∝ input voltage
  - Effect of R_f dominates C_f
- Current mode
  - Output ∝ input current
  - Low input impedance
- Charge mode
  - Output ∝ input charge
  - C_f dominates R_f

- In almost all cases we require amplification
- Amplifier-detector interaction is critical
- Most important element is the input, often a FET
- Noise is the major issue
  - Thermal or Johnson Noise
    - Brownian motion of electrons
    - No current flow required
    - White noise
  - Shot Noise
    - Fluctuations in current
    - White noise

$$\bar{i}^2 = 2q_e\bar{I}\Delta f$$

Equivalent Noise Charge

- Low noise is no use if signal is low
- Introduce ENC which is that signal charge that will produce the same output as the RMS noise

$$ENC^2 = \exp(2)\left[\frac{kT}{2R_g}\tau + \frac{eI_D}{4}\tau + \frac{kT(C_{in})^2}{2g_m\tau}\right]$$

Where

- k = Boltzman's constant
- T = temperature
- e = the electronic charge
- R_g = Load resistance and/or feedback resistance
- g_m = transconductance of input FET. (Links current in to voltage out)
- τ = Rise time of amplifier
- C_{in} = input / stray and feedback capacitance

- Note that ENC is directly related to energy resolution
- FWHM(keV) = 2.355×10^{-3} ENC/ew where w is the energy per electron

Noise Dependence

$$ENC^2 = e^2\left[\frac{kT}{2R_f}\tau + \frac{q_e I_D}{4}\tau + \frac{kT\left(C_{in}\right)^2}{2g_m\tau}\right]$$

- τ optimum at

$$\tau_{opt} = \left[\frac{kT/2g_m}{\left(kT/2R_f\right)+\left(q_e I_D/4\right)}\right]^2 C_{in}$$

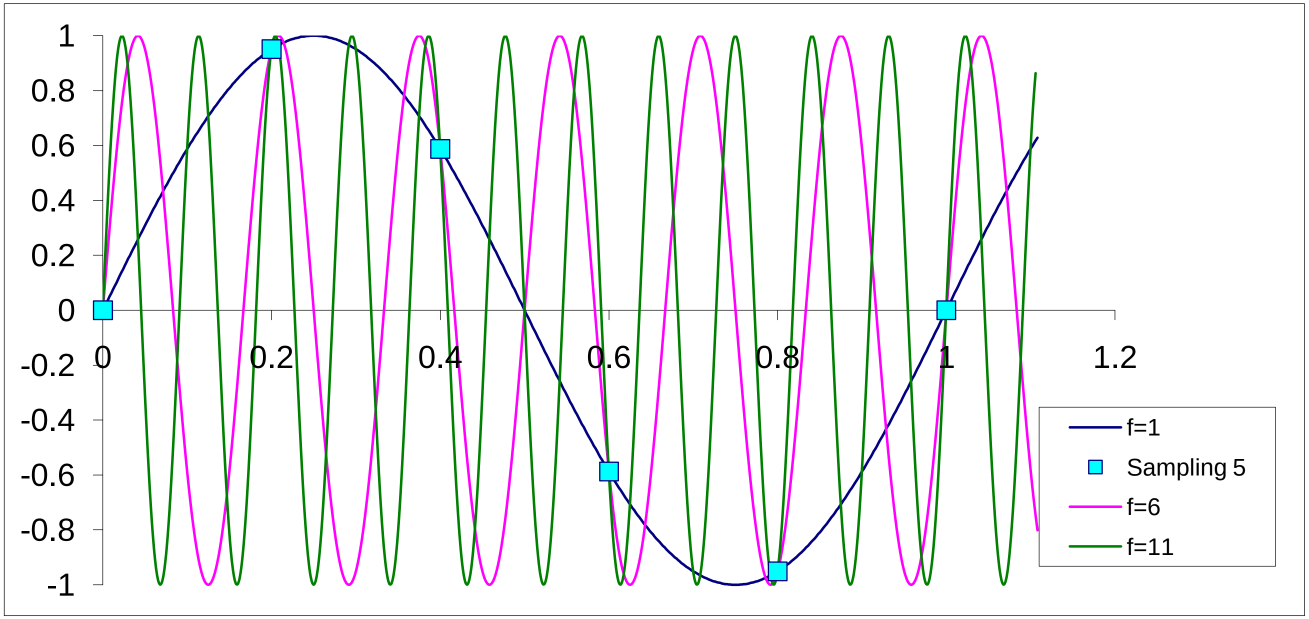
- Choosing optimum τ gives best noise performance but may not be fast enough
- We often have to sacrifice energy resolution for speed

Optimum τ

$$ENC_{\min}{}^2 = 2\exp(2)\left[\left(\frac{kT}{2R_g}\right) + \left(\frac{eI_D}{4}\right)\left(\frac{kTC_{in}{}^2}{2g_m}\right)\right]^2$$

- R_g as large as possible ~ $10^{10}\Omega$
- I_D (leakage) as small as possible
  - For Ge cooling is vital
- Low T is good
- C_{in} as small as possible (note that this includes C_f)
- g_m as large as possible but this affects C_{in}

Sampling & Aliasing



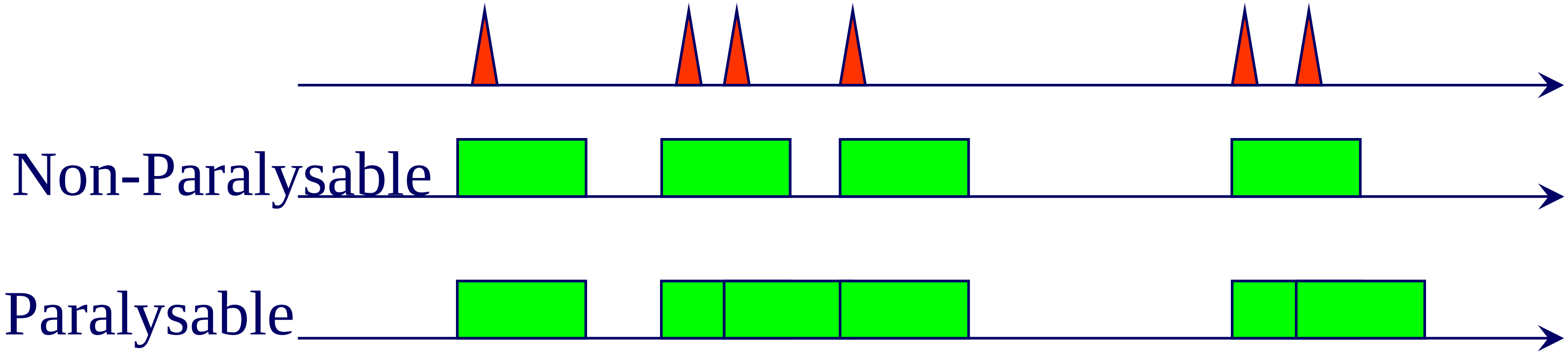
- Shannon's Theorem and Nyquist Criterion
  - The highest frequency that can be measured is twice the sampling frequency
- If the input is not band limited to frequencies less than $f_s/2$, then aliasing will occurs at frequencies $f \pm nf_s$
  - where f = signal frequency, fs = sampling frequency, n = integer
- If you have 100μm pixels, ideal PSF > 200μm

Synchrotron Detectors

- A synchrotron source is used primarily when sensitivity is an issue
  - Signal too weak
  - Time resolution too poor
  - Sample too small
- More intensity can help this but…
- It places a major strain on detectors and

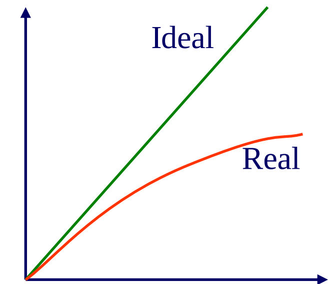
Flux is a major issue!

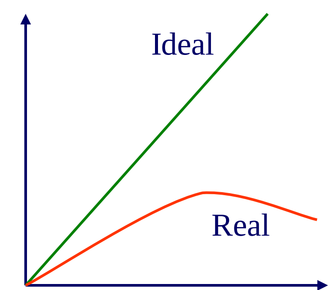
Dead Time



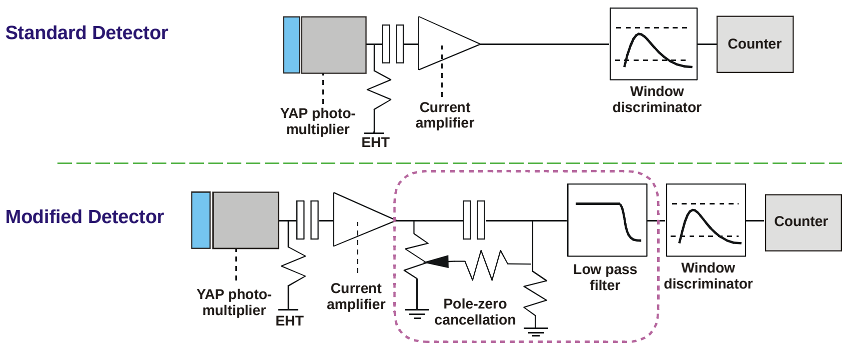
R_i=input rate, R_d=detected rate, τ dead time

- Non-paralysable
  - Fraction of time detector is dead = $R_d \tau$
  - Live time is therefore = $1 - R_d \tau$
  - Input rate = $R_i = R_d/(1 - R_d \tau)$
- Paralysable
  - R_d = Probability of getting no event within τ of an event
  - Probability of n events in time t is $$P(n,t) = \frac{e^{-R_i t}(R_i t)^n}{n!}$$
  - Detected rate $$R_d = P(0,\tau) = R_i e^{-R_i \tau}$$





EDR Detector for Powder Diffraction

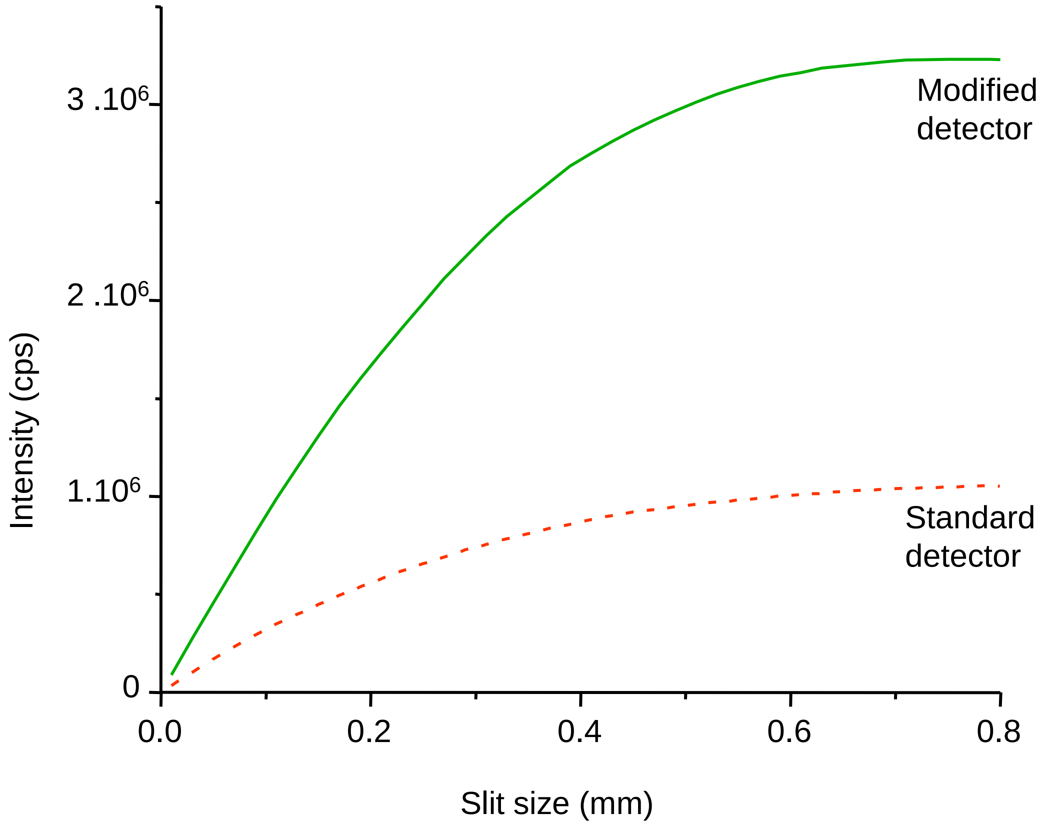


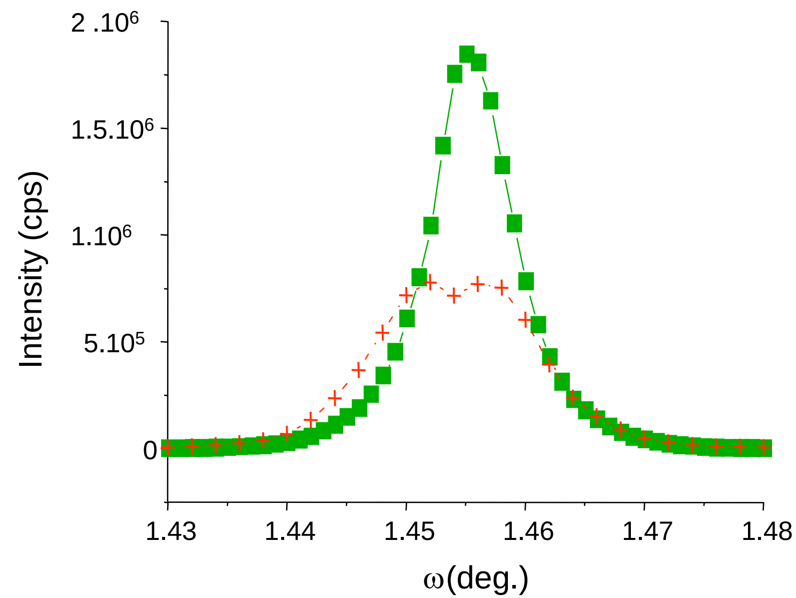
Standard detector

- Saturation count rate: 1 MHz
- Linear region: up to 300 kHz
- Electronic background: 0.2 cps
- Energy range: 4-25 keV

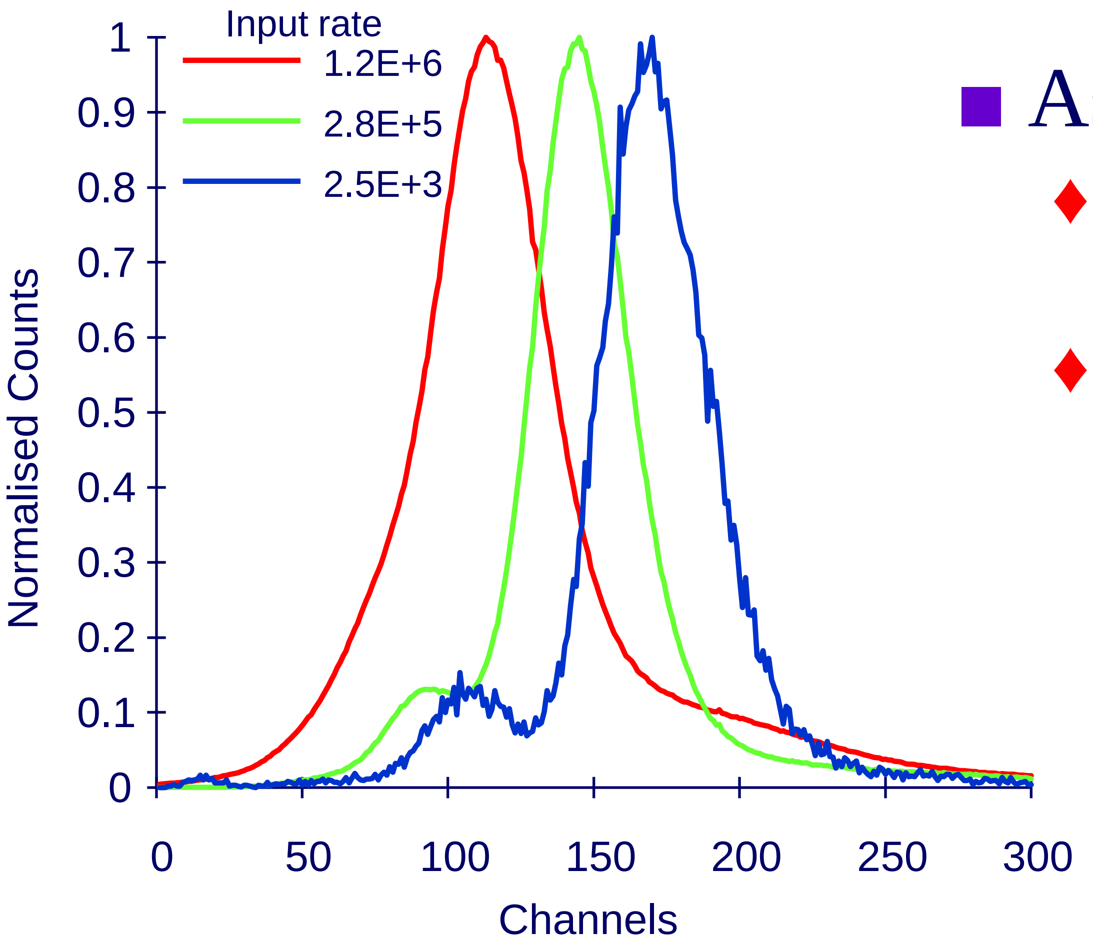
Modified detector

- Saturation count rate: 3 MHz
- Linear region: up to 2 MHz
- Electronic background: 0.7 cps
- Energy range: 4-25 keV





Spectral Peak Shift vs Rate



- As rate rises
  - Spectral resolution deteriorates
  - Note also the K escape feature

Detector Considerations

- Intensity Measurement
  - Uniformity across device
  - Ageing, radiation damage
  - Dynamic Range
  - Linearity of Response
  - Stability
- Spatial Measurement
  - Spatial Resolution
  - Spatial Distortion
  - Parallax
- Energy Measurement
  - Spectral Resolution
  - Linearity of Response
  - Uniformity of Response
  - Stability
- Time Measurement
  - Frame Rate
  - Photon Time Resolution
- Others
  - Size and weight
  - Cost

A Universal Specification?

Wombat

Counting Statistics

- Photons are quantised and hence subject to probabilities
- The Poisson distribution expresses the probability of a number of events, k occurring relative to an expected number, n

$$P(n,k) = \frac{n^k e^{-n}}{k!}$$

- The mean of P(n, k) is n
- The variance of P(n, k) is n
- The standard deviation or error (noise) is $\sqrt{n}$
- If signal = n, then SNR = $n/\sqrt{n} = \sqrt{n}$
- As n increases, SNR improves

Performance Measure - DQE

Perfect detector $SNR_{inc} = \sqrt{N_{inc}} \quad \therefore N_{inc} = SNR^2{}_{inc}$

Real detector $SNR_{Non-ideal} < \sqrt{N_{inc}}$

Can define $N_{photons}$ that describes real SNR

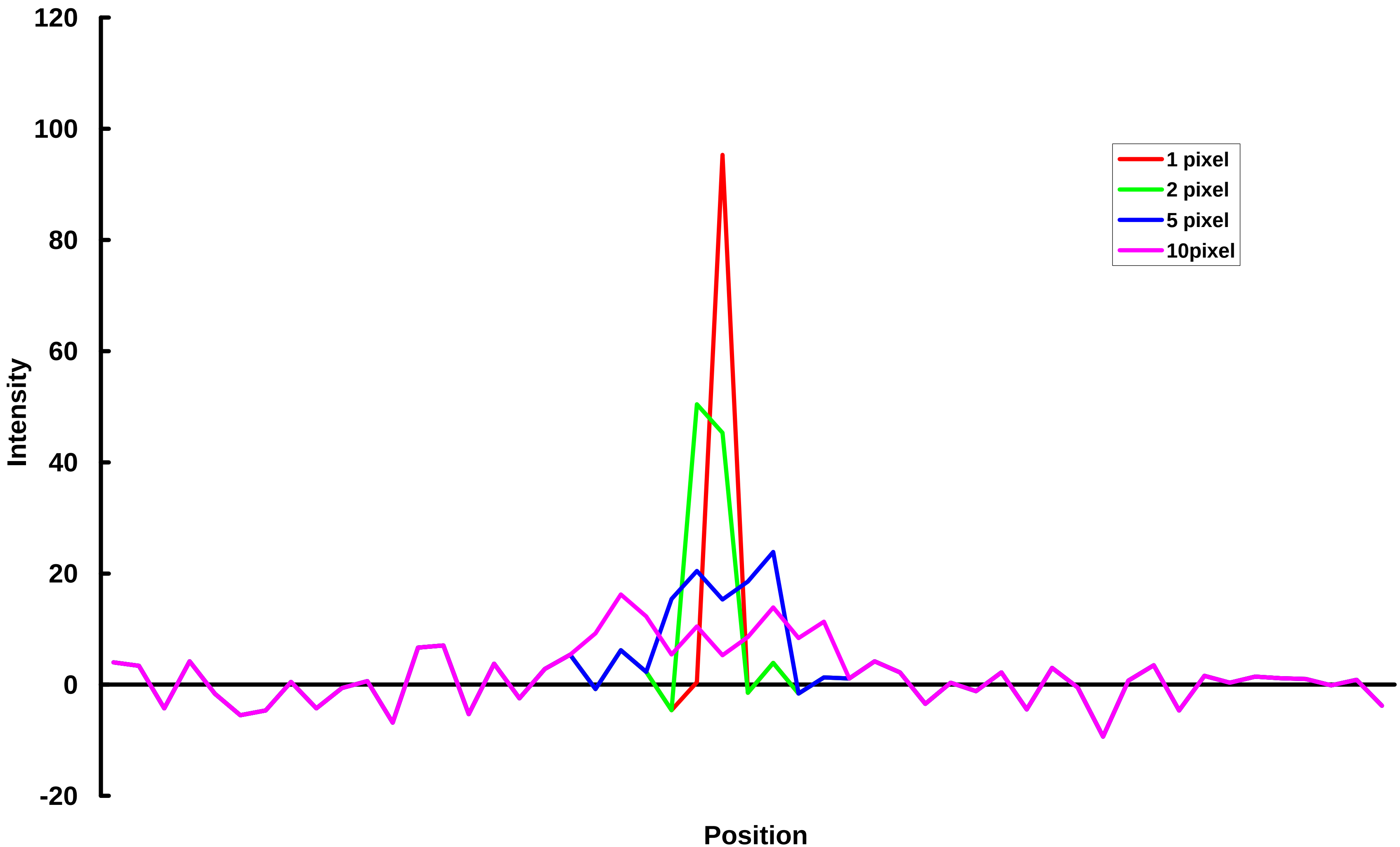
$$NEQ = SNR^2{}_{Non-ideal}$$

Ratio of this to N_{inc} is a measure of efficiency

$$DQE = \frac{NEQ}{N_{inc}} = \frac{SNR^2{}_{Non-ideal}}{SNR_{inc}{}^2}$$

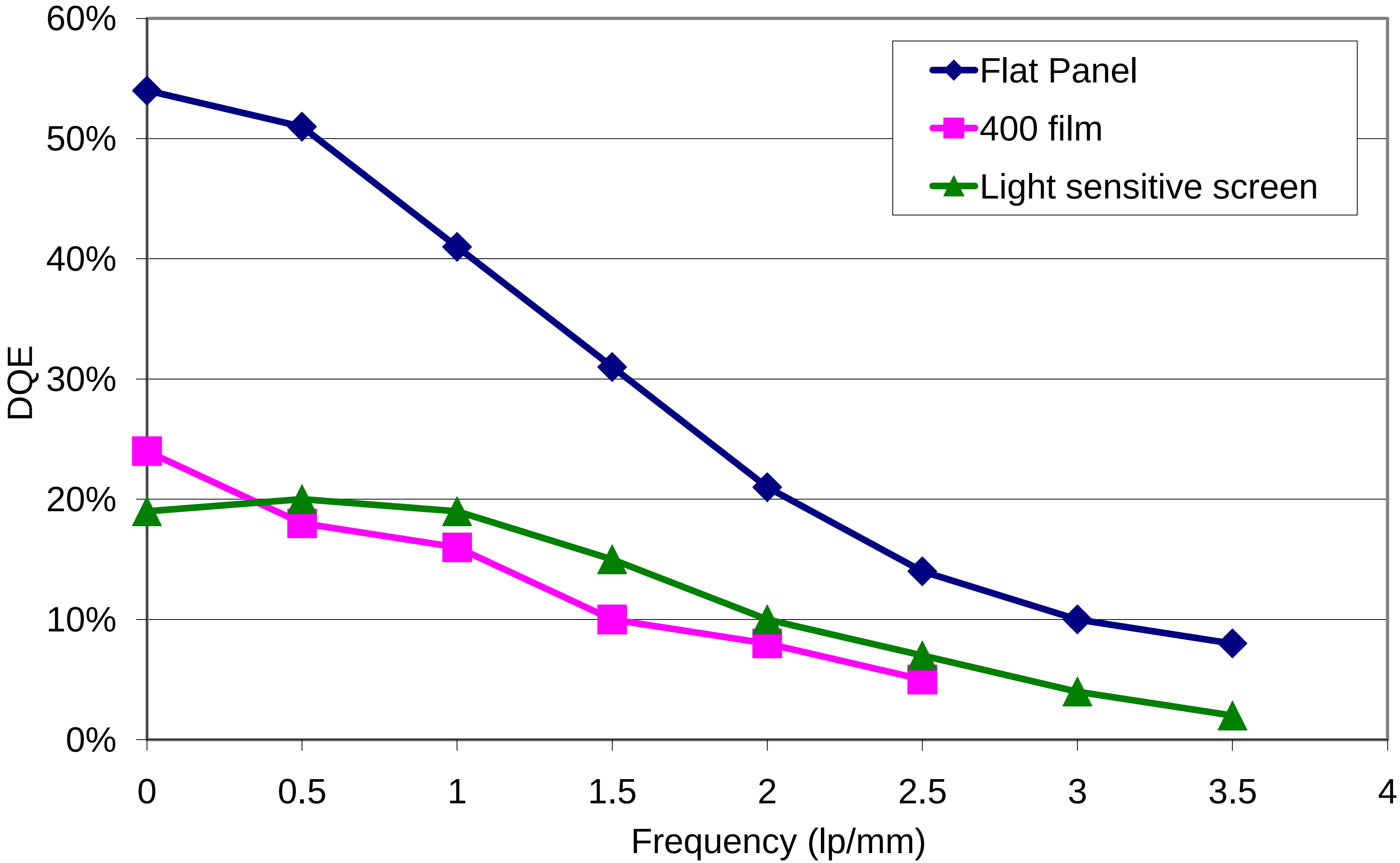
Note that DQE is f(spatial and spectral frequencies)

Effect of Peak Width



DQE Comparison

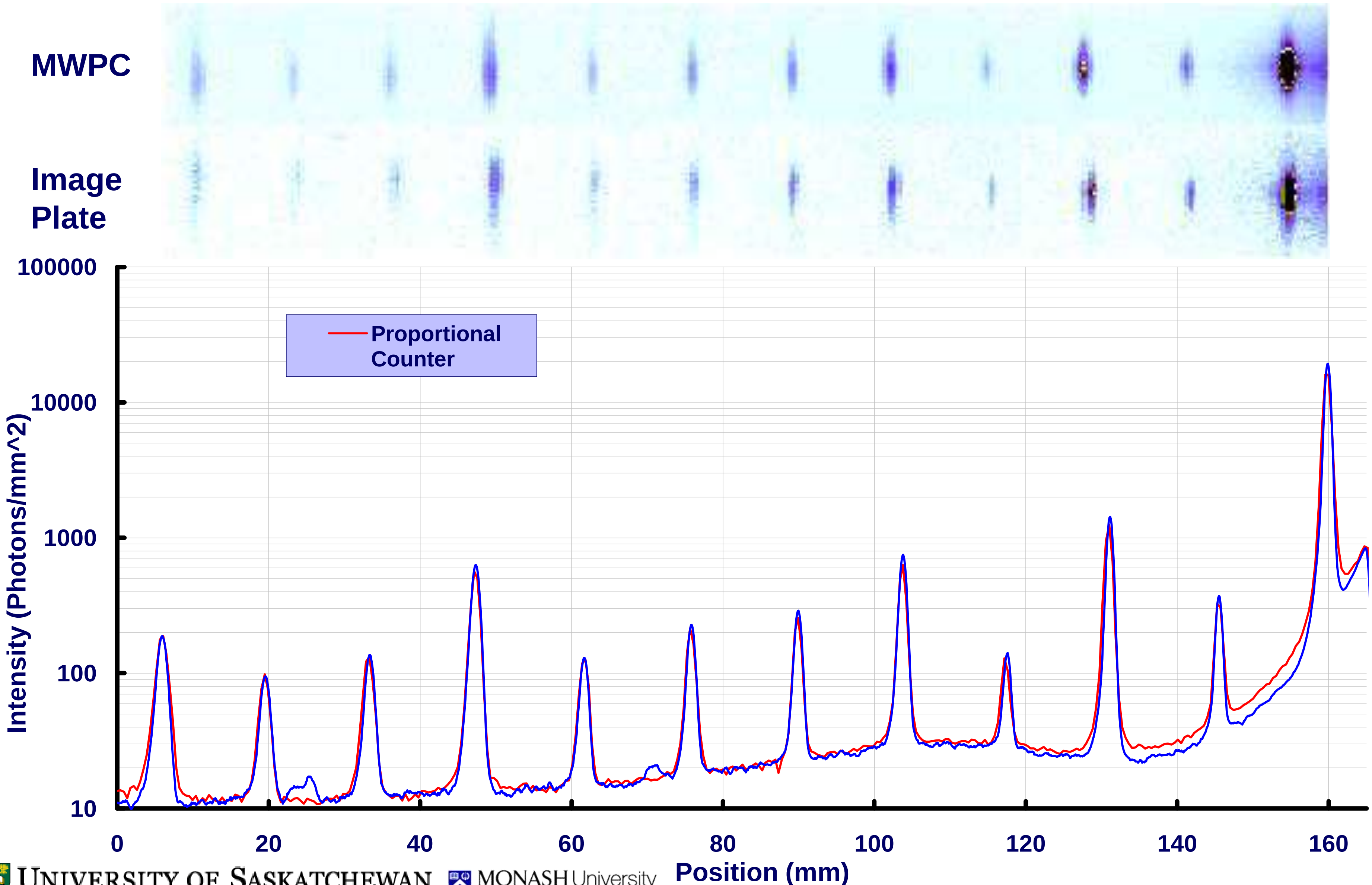
DN-5 beam
2.6μGy



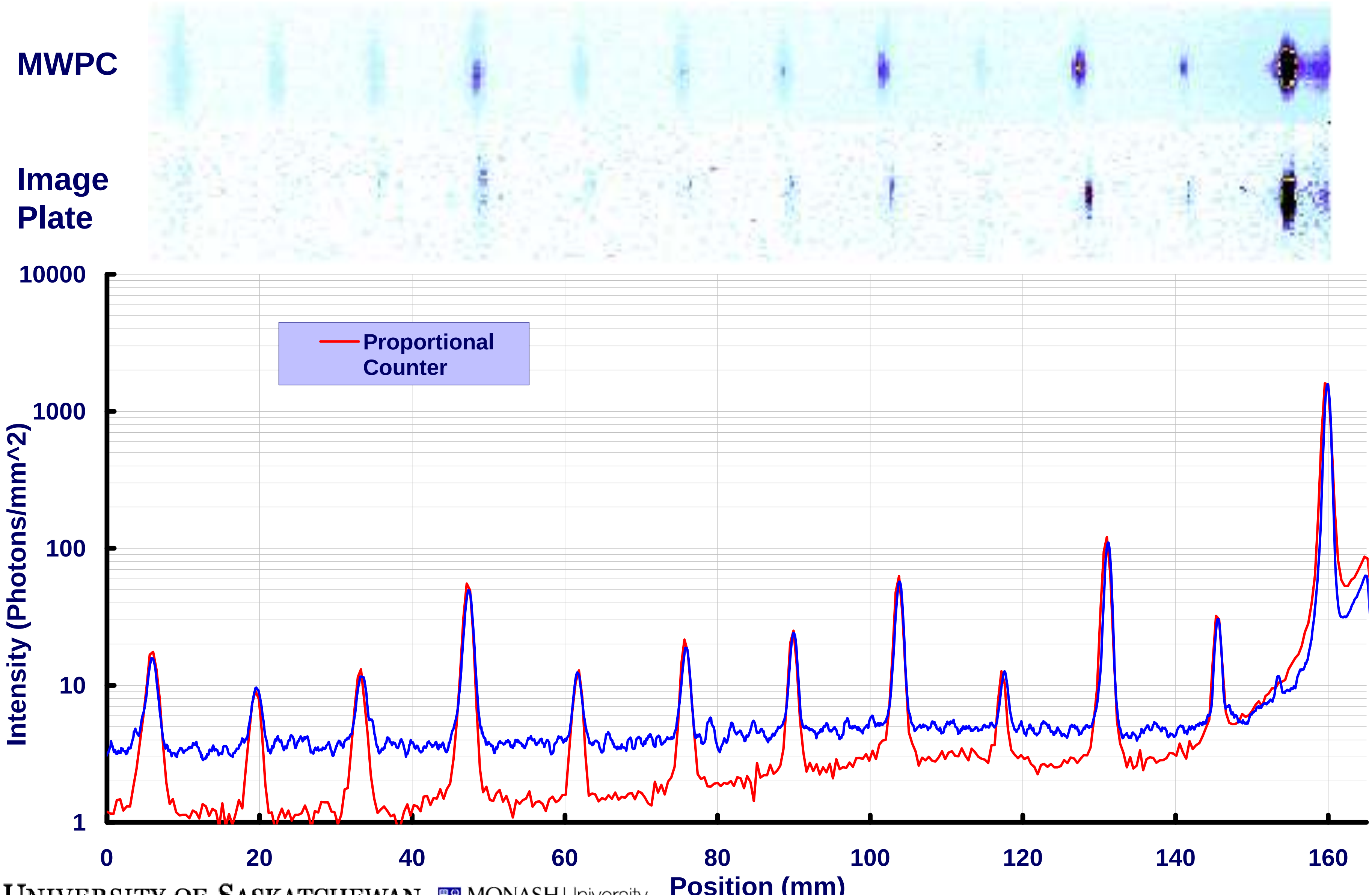
To Count or Not to Count

Tasmanian Devil

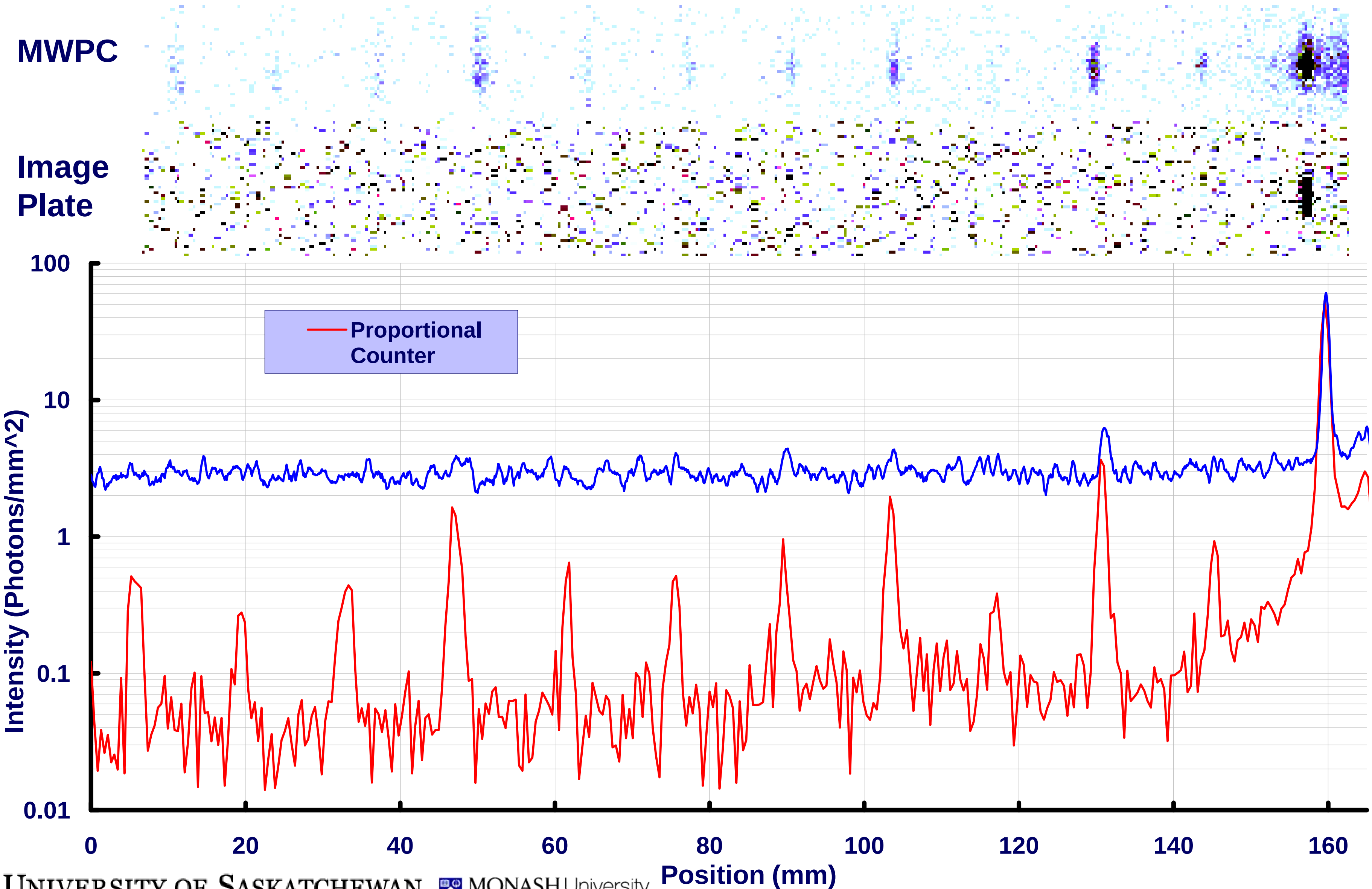
Collagen 100s Exposure



Collagen 10s Exposure

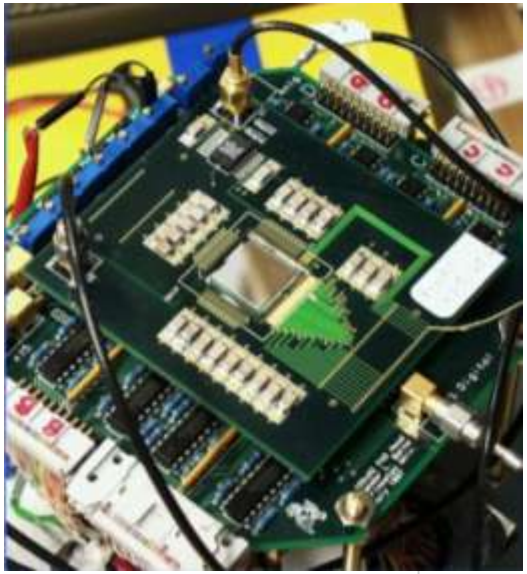


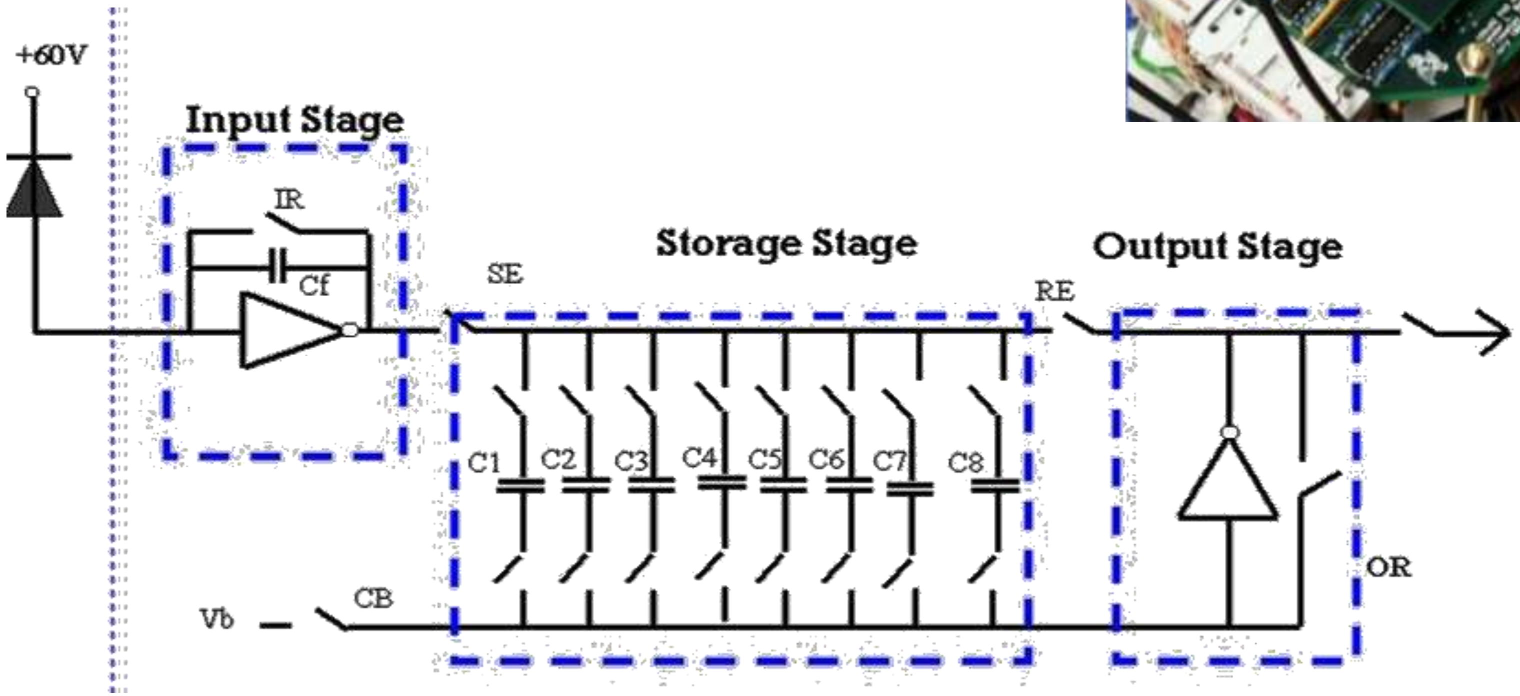
Collagen 0.3s Exposure



Cornell PAD (Integrating)

- Rapid Framing Imager
  - $15 \times 13.8 mm^2$ active area
  - 150µm square pixel
  - Storage for 8 frames
  - Selectable T_{int} down to 1µs
  - Deadtime < 1µs





Diesel Fuel Injection Movie

- Injection
  - Supersonic injection 1350psi Cerium added
  - Chamber 1atm SF_6
  - 10^8-10^9 X-rays/s/pix (6keV)
  - 1.1ms Pulse
- Movie
  - Length 1.3ms
  - Frame length 5.13μs
  - Dead time 2.56μs / frame
  - 168 frames (21 groups of 8)
  - Average 20× to improve S/N
  - Sequence 5×10^4 images

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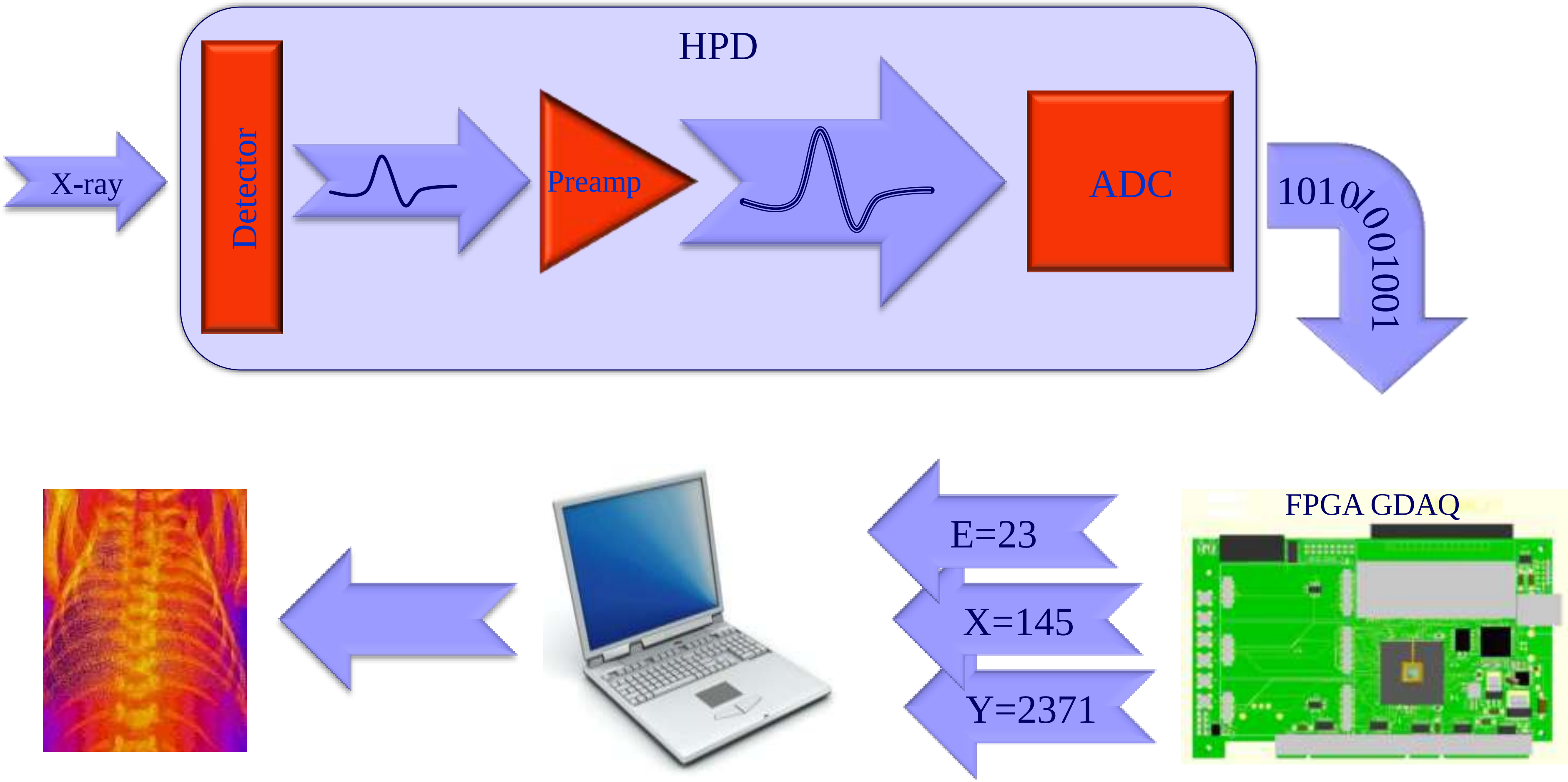
A. MacPhee et al, Science (2002) 295, 1761-1763

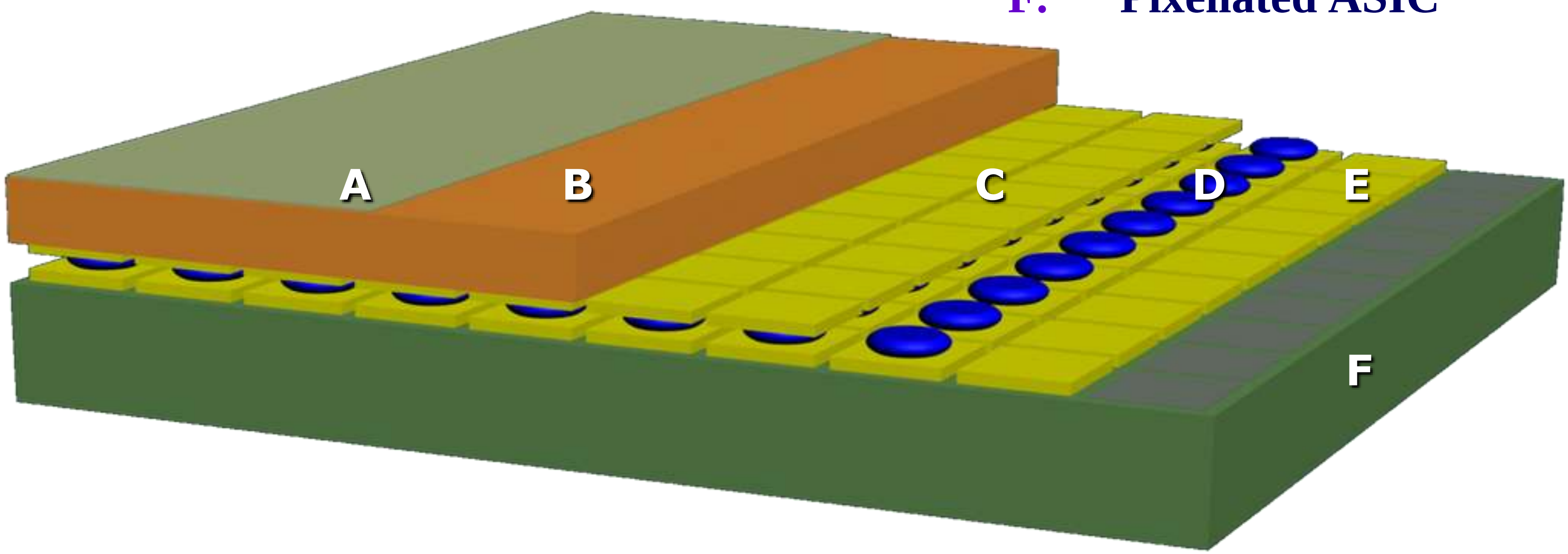
The Future

A Detector System

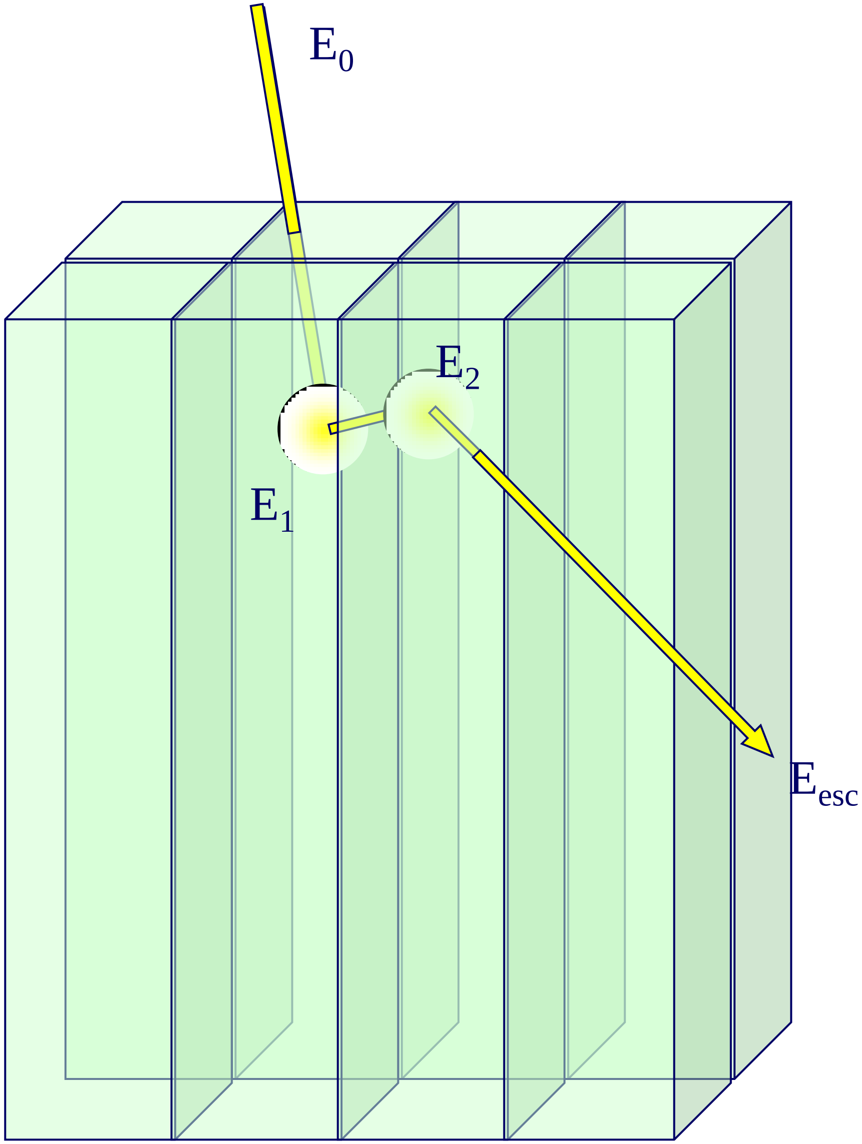


Pixel Array Detector

A. Top electrode
B. Pixellated semiconductor
C. Collection electrodes
D. Bump bonds
E. Input electrode
F. Pixellated ASIC



The Problem of Multiple Scatters



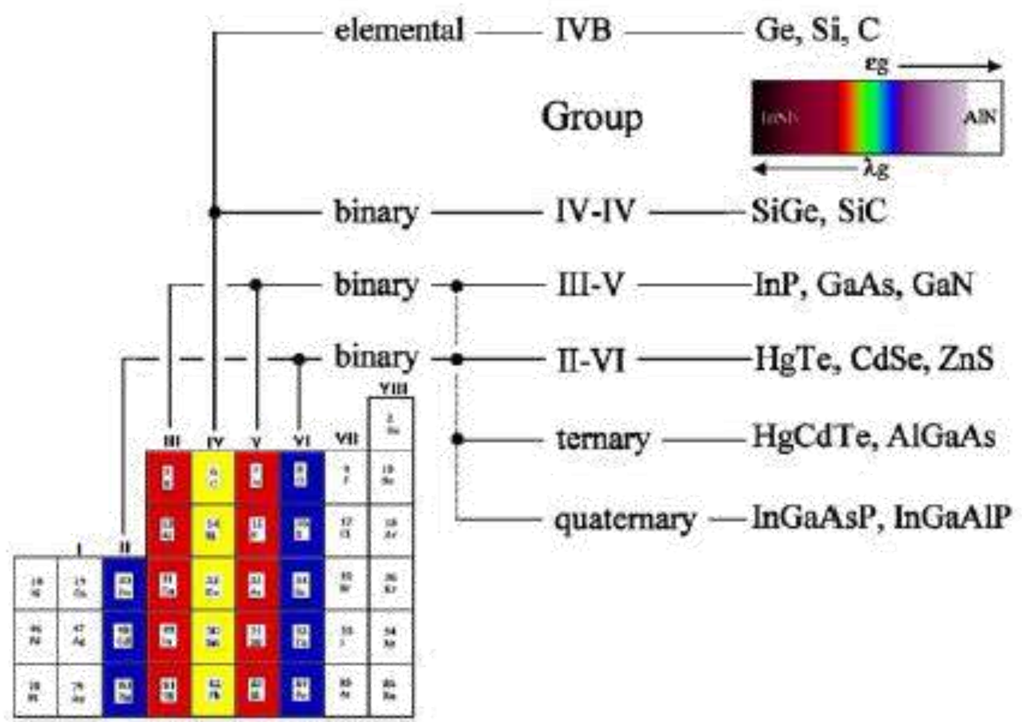
- Need to measure E_0
- $E_0=E_1+E_2+E_{esc}$
- Must be able to detect multiple deposits as single event
- Must minimise E_{esc}

Counting Pixel Detector Problems

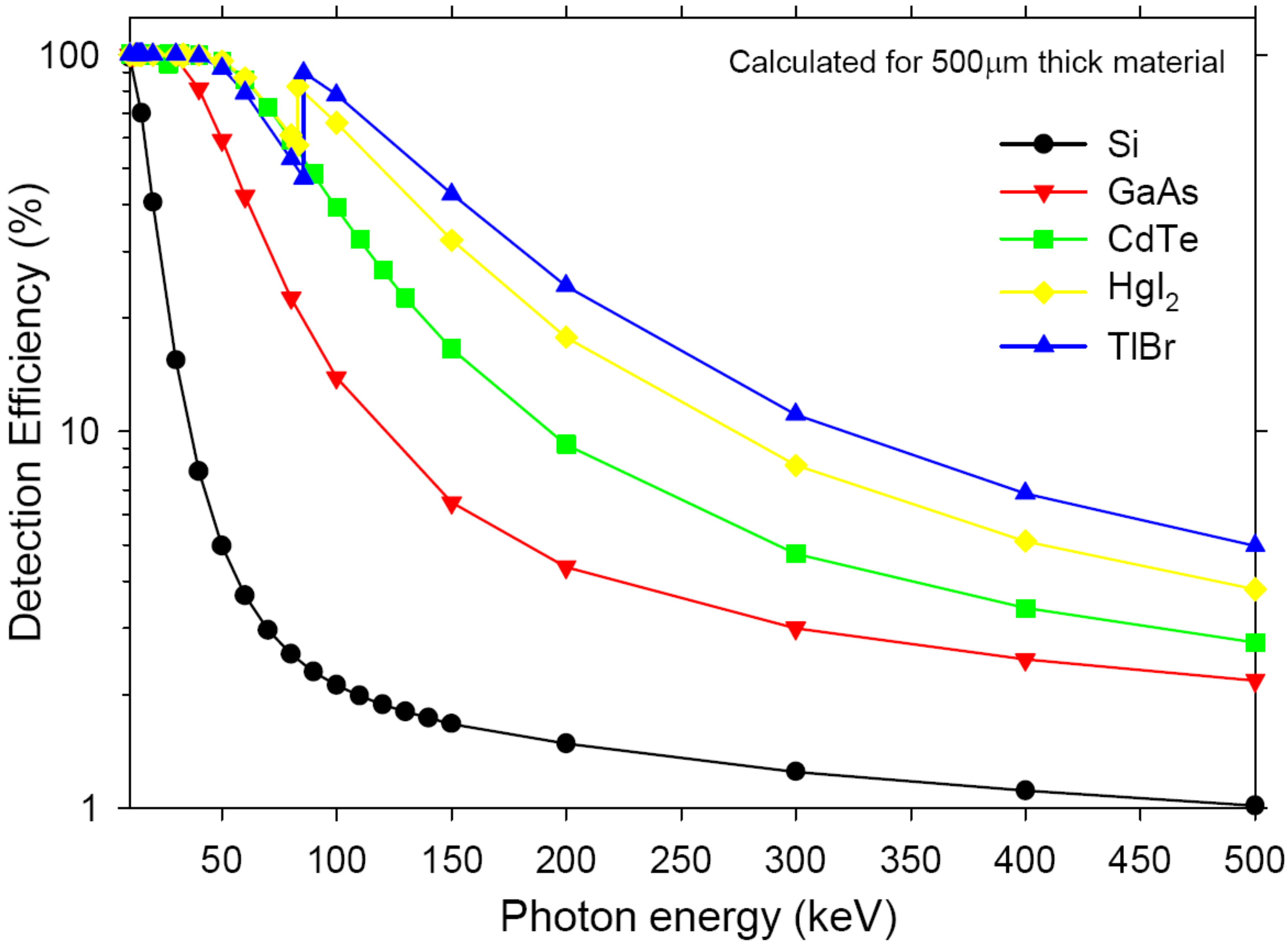
- High power consumption
  - Cooling
- Number of connections
  - Multiplexing
  - Read out time significant
- Limited number of bits in counter
  - Dynamic range issues for diffraction
  - 15bits @ 1Mcps input rate = 30ms frame
  - Read time can be significant
    - Fast read > high power
- Technology not yet good enough for microsecond framing

Available Compound Semiconductors

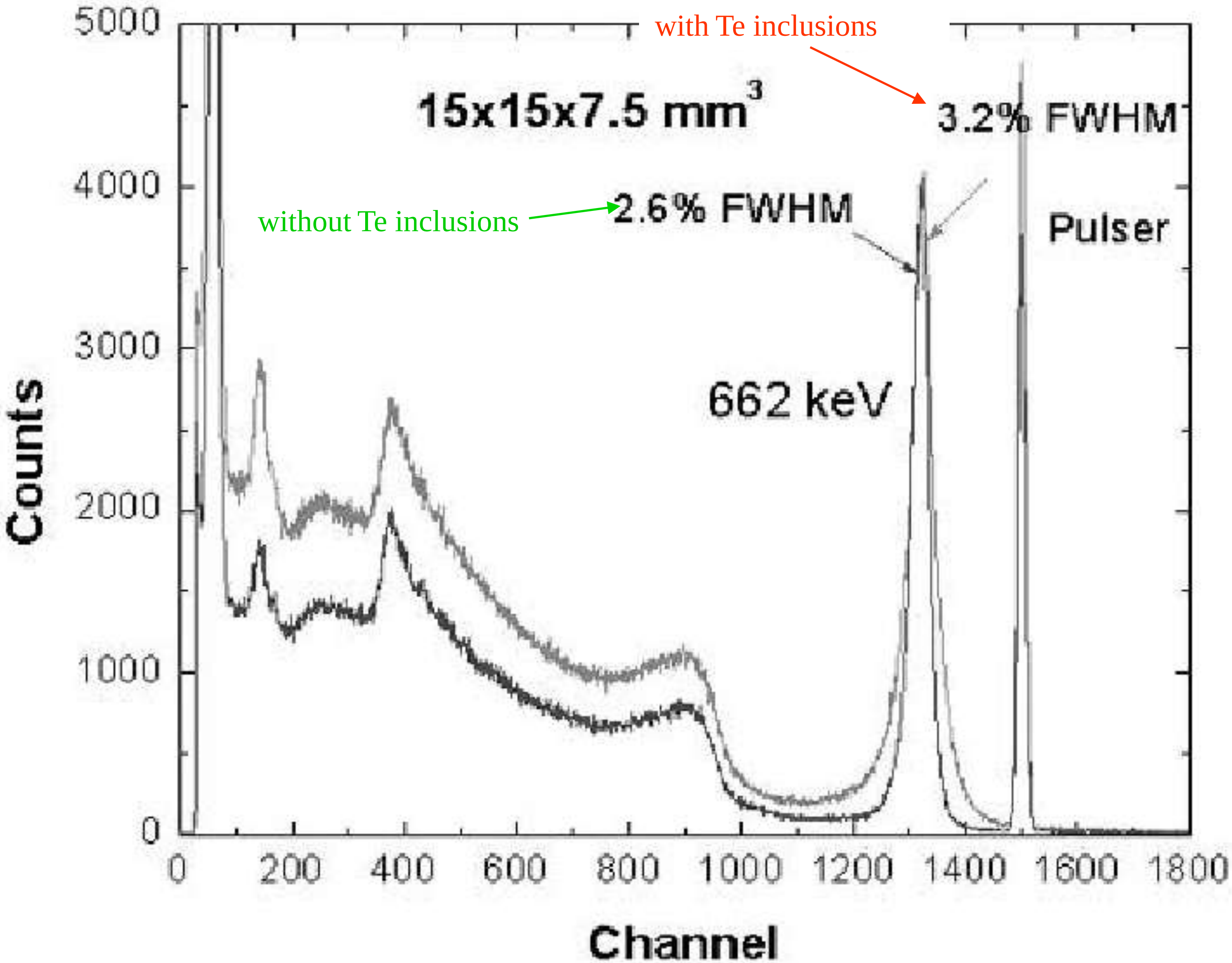
- Predominately CdZnTe, CdTe and GaAs.
- II-VI materials CdTe and CdZnTe cover a suitable range of band gaps:
  - 1.44 eV (CdTe), 1.57 eV (CdZnTe, 10% Zn), 1.64 eV (CdZnTe, 20% Zn)
- Resistivity of CdZnTe is higher than CdTe, hence lower dark current, higher spectroscopic resolution
- Poor hole transport requires electron-sensitive detectors



Absorption Efficiency of Semiconductors



CdZnTe Spectral Resolution



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